DARK MATTER DENSITY FUNCTION DEPENDING ON INTENSITY OF GRAVITATIONAL FIELD AS UNIVERSAL LAW

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1. ABSTRAT

The purpose this work is to study density function of Dark Matter (D.M.) depending on E, total intensity of gravitational field, in a group of six big spiral galaxies and try to explore a possible Universal law between both magnitudes.

Galaxies selected are the following ones: Milky Way, M31, NGC 3031, NGC 3992, NGC 7331, NGC 2841

The reason why it has been made this work is to check the model of DM proposed by the author in his previous paper [5] Abarca,M.2014 viXra. *Dark Matter model by quantum vacuum*.

Briefly, in that work was exposed a original theory of DM. This new theory defends that DM is generated by the own gravitational field according an unknown quantum gravitational mechanism. In other words, Density of Dark Matter in a point P of space, depend on E, according a Universal law. Therefore if two different points P and Q belonging different galaxies had the same E, intensity of gravitational field, then they should have the same density of DM. In short, If E(P) = E(Q) then $D_{DM}(P) = D_{DM}(Q)$.

In the present paper, for each galaxy it will be calculate E, total intensity of gravitational field, depending on R. This calculus is carry out through the Rotational Curve and the Virial Theorem.

In addition, for each galaxy it is considered a Dark matter density function, which is tabulated depending on R. After, it is plotted DM density depending on E.

Data of eight galactic set are studied statistically and it is concluded that its coefficient correlation is enough high to fit a potential function to DM density depending on E. Statistical process has been carefully detailed through the paper, so I have found plausible to postulate this function as Universal:

Dark matter function Density as Universal law for big galaxies

Density $DM = 1.26 \cdot 10^{\Lambda-6} \cdot E^{\Lambda B}$ where B = 1.74 inside a dominion 100 < E < 6500

Where Unit for $D_{D.M.}$ is $10^{^{-3}} M_{\odot}/pc^{^3}$ or $m M_{\odot}/pc^{^3}$ and Unit for E is Km $^{^2}/$ s $^{^2}/$ Kpc

Big galaxy set considered in this paper have their flat speed region in rotation curve above 200 Km/s.

In chapter 12 it is calculated Local DM density. From Solar parameters $R_{\Theta} = 8 \, Kpc$ and $V_{\Theta} = 220 \, Km/s$ it is got $E_{\Theta} = 6050 \, \mathrm{Km^2/s^2/Kpc}$ and $\varphi_{LOCAL-DM} = 4.8 \, \mathrm{m} \, M_{\Theta}/\mathrm{pc^3} = 0.182 \, \mathrm{GeV/cm^3}$.

Also, in chapter 12 are quoted data about recent Local DM density measures and it is concluded that $\varphi_{LOCAL-DM}$ calculated in this paper is included inside interval defined by $\varphi_{LOCAL-DM}$ measures published in recent papers.

I think that it has been justified properly that results could confirm DM density as universal law, especially if it is considered experimental error in rotation curves and error in calculus of DM density functions.

Unfortunately this law is wrong for intermediate galaxies, whose flat region in rotation curve is under 150 Km/s. Specifically have been studied NGC 3198 (Ursa Major), NGC 2403 (M81 group of galaxies) and M33 (Local group). In these galaxies density of dark matter is more than four times bigger that density of dark matter inside big galaxies although this result has not been published in this paper.

However, I think that these results can not discard this model because DM density function inside an intermediate galaxy may be influenced by gravitational field of a nearby giant galaxy. In other words, gravitational field of a giant neighbour galaxy might generate DM inside an intermediate galaxy placed near a big galaxy.

In addition, a nearby giant galaxy may break significantly the spherical symmetry of gravitational field inside an intermediate galaxy and therefore calculus made through Virial Theorem become wrong.

The author think that it may be worth to study a bigger number of galaxies in order to check or reject results got in this paper because Dark matter nature is one of the most important challenges for astrophysics science nowadays.

2. INTRODUCTION

Method followed to calculate E(P), total gravitational field at point P, has been through Virial theorem. It is a well known method but the problem is that galaxies do not have perfect spherical symmetry and their stars do not have perfect gravitational equilibrium, so this method give a approximate measure of E and his error it will depend on degree of spherical symmetry and gravitational equilibrium of stars inside each galaxy.

Inside a galaxy with spherical symmetry, E(P) at a point P at R distance from galactic centre depend on spin speed of stars at R distance from galactic centre.

According Virial theorem $V^2 = GM(\langle R \rangle / R)$ where $M(\langle R \rangle)$ is total mass inside a sphere with radius R, supposing a mass distribution with spherical symmetry.

As $E = GM(\langle R)/R^2$ it is right conclude that $E = V^2/R$.

As measures of rotation curve are in Km/s and radii are in Kpc, unit of E in this work has been Km²/s²/Kpc which is a bit weird as unit of E but it allows to simplify calculus.

Considering Dark matter, astrophysical researchers measure baryonic mass inside a sphere R, M_{BARIONIC} (R). After they calculate density Dark Matter function which is necessary to reach total mass M(R) to get the speed given by $V^2 = GM(\langle R)/R$.

In the whole paper unit of Dark matter density is 10⁻³Msolar/pc³.

As it is known there are several model of Dark matter density. For Milky Way and M31, author has chosen NFW

profile
$$D_{NFW}(R) = \frac{D_0}{x \cdot (1+x)^2}$$
 where x= R/h. The other four galaxies authors have chosen pseudo-isothermal profile

$$D_{ISO}(R) = \frac{D_0}{1+x^2}$$
 where x= R/R_C.

Authors of rotation curves of galaxies and his data parameter of Dark matter density function will be quoted specifically for each galaxy.

Remember that for the whole paper, his graphs and tables, unit for E is $Km^2/s^2/Kpc$ and unit for Density of dark matter is 10^{-3} Msolar/pc³.

3. GALACTIC SAMPLE

After a meticulous research in arXiv, it has been selected six big galaxies from three different papers. Two galaxies have data from two different papers, because authors give different dark matter parameters for the same galaxy. Unfortunately, I have not found any more papers about rotation curves of big galaxies, because it is a hard work measuring rotation curves. It is a great human and technological achievement to be able to measure rotation curves of galaxies which are lot of Mpc away_i.

Galactic sample are the following ones:

NGC 2841

It is a spiral galaxy in the Ursa Major galaxy cluster. Its distance is 14 Mpc. Aprox. Calculus for this galaxy has been made two times according data provided by two different papers: [2] Randriamampandry & T. Carignan. and [3] Bottema, R.B. & Pestaña, J.L.G.

NGC 7331

It is a spiral galaxy about 12 Mpc away in the Constelation Pegasus. This galaxy is similar in size and structure to Milky Way. Calculus for this galaxy has been made two times according data provided by two different papers: [2] Randriamampandry & T. Carignan. and [3] Bottema, R.B. & Pestaña, J.L.G.

NGC 3992 It is one of the most prominent members of the Ursa Major cluster of galaxies. Its approximate distance is 18 Mpc. Calculus has been made according data from [3] Bottema, R.B. & Pestaña, J.L.G.

NGC 3031

This galaxy also is known as M81 or Bode's galaxy. It is the main galaxy in M81 group of galaxies. It is 3 Mpc away. Calculus has been made according data from [2] Randriamampandry & T. Carignan.

Milky Way

Thousand of stars are possible to see with naked eye. However measures toward galactic centre are difficult because of gas and cosmic dust. Calculus has been made according data from [1] Sofue, Y.

M31 or Andromeda Galaxy

It is the nearest major spiral galaxy to the Milky Way. M31 is 780 Kpc away to our Galaxy. it is the largest galaxy of the Local Group, which also contains the Milky Way, the Triangulum Galaxy (M33), and about 44 other smaller galaxies. Calculus has been made according data from [1] Sofue, Y.

Apart from these six galaxies, I have studied NGC 2998 and NGC 2903 as well. However, I have not considered NGC 2998 because it is 67 Mpc away, and it has an anomalous high density of D.M. I think that a distance so huge may originate lot of errors in measures of rotation curve.

In addition NGC 2903 has been rejected of the galactic sample because it has an outstanding high density of dark matter by unknown reasons. This galaxy has a anomalous behaviour to explain its rotation curve by MOND theory as well.

Milky Way, M31, NGC 3031, NGC 3992, NGC 7331, NGC 2841 are big spiral galaxies. Their flat region rotation curve have speeds from 200 km/s up to 300 Km/s.

Considering their distances:

NGC 2841, NGC 3992, belong to Ursa Major Cluster which is 18 Mpc away.

NGC 7331 is 12 Mpc away in constellation Pegasus.

NGC 3031 is 3 Mpc away and belong to M81 group of galaxies.

M31 is only 770 Kpc away and Milky Way no need presentation.

Therefore it has been covered a wide interval of distances. In addition, galaxies have been chosen from four different galaxy clusters which do not have gravitational interaction each other.

4.GALACTIC CALCULUS

Each galaxy is identified with his name, his cluster and his approximate distance. In addition it is quoted the author

who has measured the rotation curve and has calculate density of dark matter through NFW or pseudo-isothermal profiles.

Through the rotation curve is tabulated the spin speed for a wide number of radii and E, intensity of gravitational field according the formula $E = v^2/R$

In the second column, it is tabulated $X=R/R_c$ where R_c is a parameter of dark matter density of pseudo-isothermal profile. Similarly X=R/h where h is a parameter of dark matter density of NFW profile.

Finally, in the last column it is tabulated dark matter density using a pseudo-isothermal profile $D_{ISO}(R) = \frac{D_0}{1+x^2}$

or NFW profile $D_{NFW}(R) = \frac{D_0}{x \cdot (1+x)^2}$ In the graphic is plotted density of dark matter depending E, through table

data. The same process has been followed for the six galaxies.

Graph of rotation curve which come from paper quoted show several curves. Reader may consult original paper to know directly meaning each curve. In general all graphs have similar curves. Briefly is explained bellow.

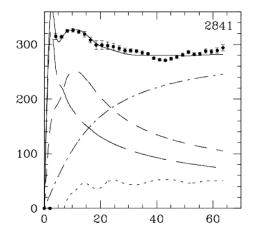
Dots are experimental data and they have a fitted function. The curves with relative maximum correspond to field generated by baryonic mass from bulge and from disk. The curve which is always increasing correspond to dark matter. It is important to emphasise that Dark matter contribution is not important for distances bellow 10 Kpc. I mean that calculus of Dark matter density function might have too much error for distances bellow 10 Kpc. This is the reason why it will be considered gravitational field E bellow 6000 to do statistical analysis in this paper.

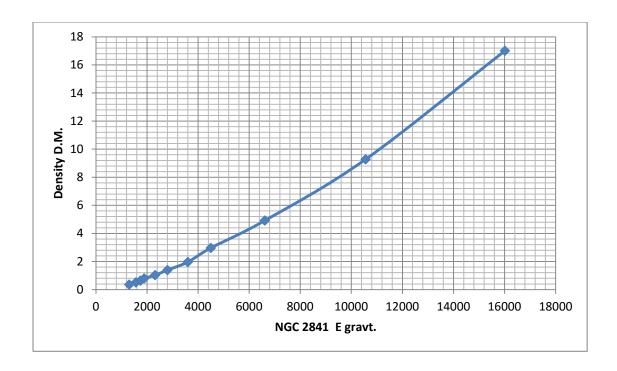
4.1 NGC 2841 Bottema, R.B.

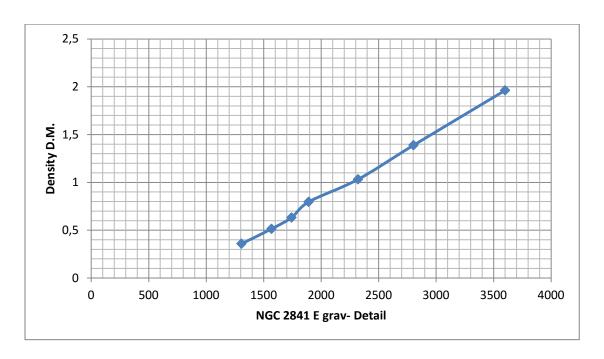
| 1 Name | Cluster | Distance | Author |
|----------|------------|----------|------------------|
| NGC 2841 | Ursa Major | 14 Mpc | [3] Bottema,R.B. |

| Dark matter der | nsity function profile | Isothermal |
|----------------------|---|------------|
| $Rc = 6,39 \pm 0,52$ | 2 Крс | |
| Do = 32 ± 5 | 10 ⁻³ Msolar/pc ³ | |

| Radii | Х | Speed | E gravt. | Density |
|-------|------------|-------|------------|------------|
| 60 | 9,38967136 | 280 | 1306,66667 | 0,35888148 |
| 50 | 7,82472613 | 280 | 1568 | 0,51425169 |
| 45 | 7,04225352 | 280 | 1742,22222 | 0,63249438 |
| 40 | 6,25978091 | 275 | 1890,625 | 0,79631987 |
| 35 | 5,47730829 | 285 | 2320,71429 | 1,03222789 |
| 30 | 4,69483568 | 290 | 2803,33333 | 1,38879955 |
| 25 | 3,91236307 | 300 | 3600 | 1,96239743 |
| 20 | 3,12989045 | 300 | 4500 | 2,96400194 |
| 15 | 2,34741784 | 315 | 6615 | 4,91523484 |
| 10 | 1,56494523 | 325 | 10562,5 | 9,27790752 |
| 6 | 0,93896714 | 310 | 16016,6667 | 17,0062669 |





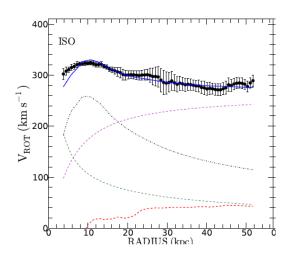


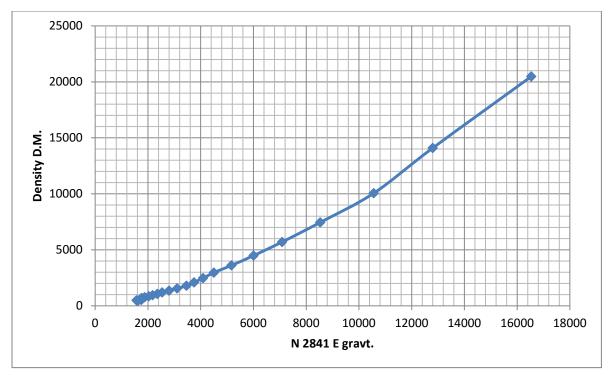
4.2 NGC 2841 Randriamampandry, T.

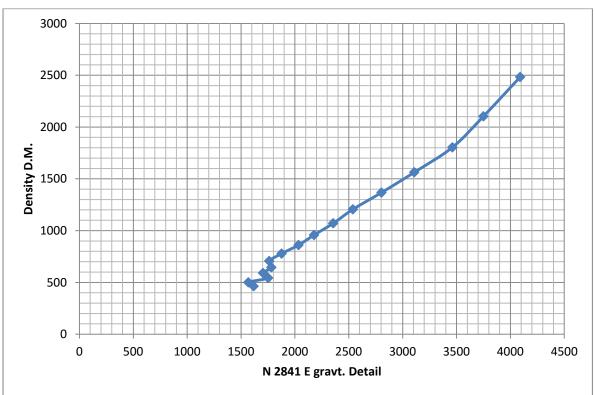
| 2 Name | Cluster | Distance | Author |
|----------|------------|----------|--------------------------|
| NGC 2841 | Ursa Major | 14 Mpc | [2] Randriamampandry, T. |

| Dark matter density function profile Isothermal | |
|---|--|
| $Rc = 5.08 \pm 0.23$ Kpc | |
| Do = $49,06 \pm 3,61 \cdot 10^{-3} \text{ Msolar/pc}^3$ | |

| Radii | Х | Speed | E gravt. | Density |
|-------|------------|-------|------------|------------|
| 6 | 1,18110236 | 315 | 16537,5 | 20,4843185 |
| 8 | 1,57480315 | 320 | 12800 | 14,0976811 |
| 10 | 1,96850394 | 325 | 10562,5 | 10,0635737 |
| 12 | 2,36220472 | 320 | 8533,33333 | 7,45591441 |
| 14 | 2,75590551 | 315 | 7087,5 | 5,70795966 |
| 16 | 3,1496063 | 310 | 6006,25 | 4,49266583 |
| 18 | 3,54330709 | 305 | 5168,05556 | 3,61932196 |
| 20 | 3,93700787 | 300 | 4500 | 2,97332775 |
| 22 | 4,33070866 | 300 | 4090,90909 | 2,4834172 |
| 24 | 4,72440945 | 300 | 3750 | 2,10376956 |
| 26 | 5,11811024 | 300 | 3461,53846 | 1,80400461 |
| 28 | 5,51181102 | 295 | 3108,03571 | 1,56341316 |
| 30 | 5,90551181 | 290 | 2803,33333 | 1,36752347 |
| 32 | 6,2992126 | 285 | 2538,28125 | 1,20599568 |
| 34 | 6,69291339 | 283 | 2355,55882 | 1,0712939 |
| 36 | 7,08661417 | 280 | 2177,77778 | 0,95782709 |
| 38 | 7,48031496 | 278 | 2033,78947 | 0,8613801 |
| 40 | 7,87401575 | 274 | 1876,9 | 0,77872863 |
| 42 | 8,26771654 | 272 | 1761,52381 | 0,7073737 |
| 44 | 8,66141732 | 280 | 1781,81818 | 0,64535521 |
| 46 | 9,05511811 | 280 | 1704,34783 | 0,59111878 |
| 48 | 9,4488189 | 290 | 1752,08333 | 0,54341939 |
| 50 | 9,84251969 | 280 | 1568 | 0,5012506 |
| 52 | 10,2362205 | 290 | 1617,30769 | 0,46379186 |





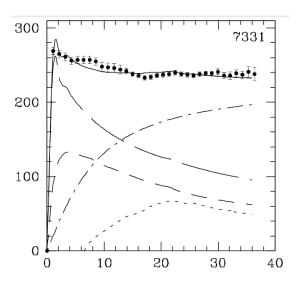


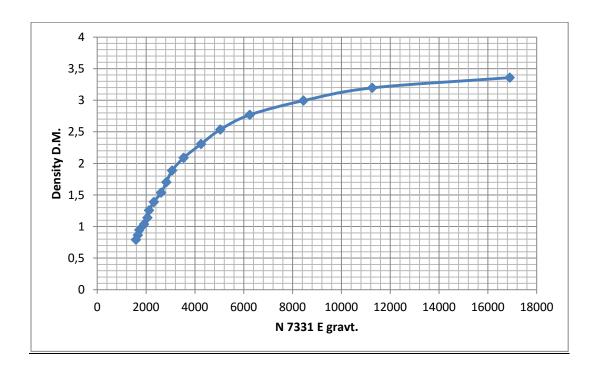
4.3 NGC 7331 Bottema, R.B.

| 3 Name | Constelation | Distance | Author |
|----------|--------------|----------|------------------|
| NGC 7331 | Pegasus | 12 Mpc | [3] Bottema,R.B. |

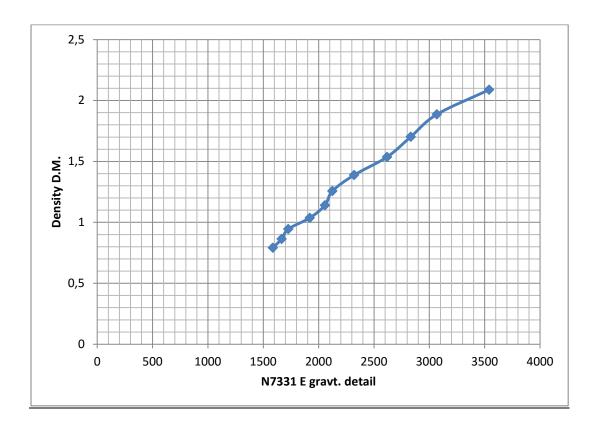
| Dark matter density function profile | Isothermal |
|---|------------|
| Rc = 19,46 ± 1,79 Kpc | |
| Do = $3.5 \pm 0.7 \cdot 10^{-3}$ Msolar/pc ³ | |

| Radii | X | Spin speed | E gravt. | Density |
|-------|------------|------------|------------|------------|
| | | | | |
| 4 | 0,20554985 | 260 | 16900 | 3,35811707 |
| 6 | 0,30832477 | 260 | 11266,6667 | 3,19615975 |
| 8 | 0,41109969 | 260 | 8450 | 2,9940044 |
| 10 | 0,51387461 | 250 | 6250 | 2,76884031 |
| 12 | 0,61664954 | 246 | 5043 | 2,53576028 |
| 14 | 0,71942446 | 244 | 4252,57143 | 2,30631629 |
| 16 | 0,82219938 | 238 | 3540,25 | 2,08829075 |
| 18 | 0,92497431 | 235 | 3068,05556 | 1,88620527 |
| 20 | 1,02774923 | 238 | 2832,2 | 1,70211236 |
| 22 | 1,13052415 | 240 | 2618,18182 | 1,53637824 |
| 24 | 1,23329908 | 236 | 2320,66667 | 1,38832331 |
| 26 | 1,336074 | 235 | 2124,03846 | 1,2566902 |
| 28 | 1,43884892 | 240 | 2057,14286 | 1,13995887 |
| 30 | 1,54162384 | 240 | 1920 | 1,03654439 |
| 32 | 1,64439877 | 235 | 1725,78125 | 0,94491234 |
| 34 | 1,74717369 | 238 | 1666 | 0,8636397 |
| 36 | 1,84994861 | 239 | 1586,69444 | 0,7914416 |





In the whole galactic sample, NGC 7331 is the only one which has a convex curve as it is shown above.

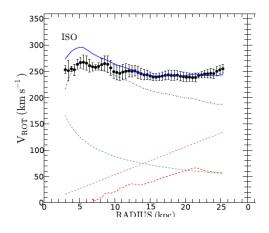


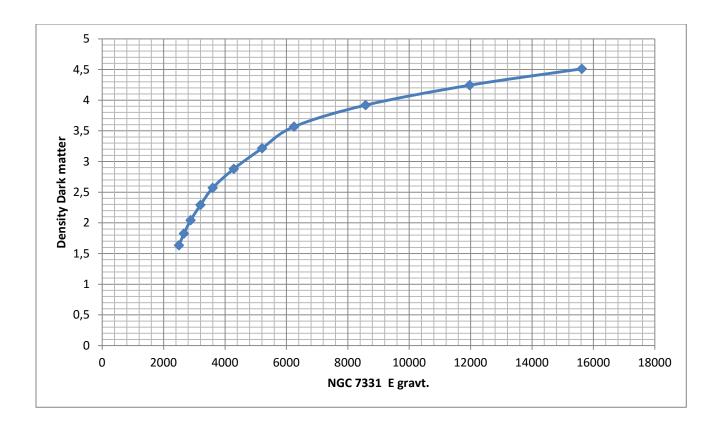
4.4 NGC 7331 Randriamampandry, T.

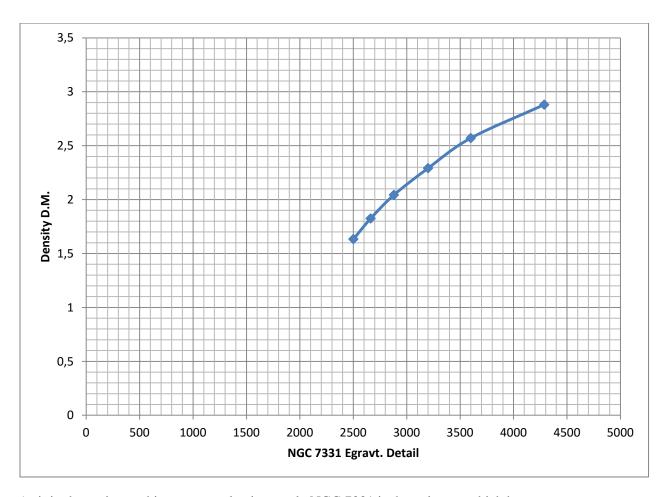
| 4 Name | Constelation | Distance | Author |
|----------|--------------|----------|--------------------------|
| NGC 7331 | Pegasus | 12 Mpc | [2] Randriamampandry, T. |

| Dark matter density function profile Isotermal | | | |
|--|--|--|--|
| $Rc = 17,38 \pm 2,75$ Kpc | | | |
| Do = $4,75 \pm 0.6 10^{-3} \text{Msolar/pc}^3$ | | | |

| Radii | Х | Speed | E gravt. | Density |
|-------|------------|-------|------------|------------|
| 4 | 0,2301496 | 250 | 15625 | 4,51105468 |
| 6 | 0,3452244 | 268 | 11970,6667 | 4,24417922 |
| 8 | 0,46029919 | 262 | 8580,5 | 3,91954503 |
| 10 | 0,57537399 | 250 | 6250 | 3,56859722 |
| 12 | 0,69044879 | 250 | 5208,33333 | 3,21658913 |
| 14 | 0,80552359 | 245 | 4287,5 | 2,88076381 |
| 16 | 0,92059839 | 240 | 3600 | 2,57104001 |
| 18 | 1,03567319 | 240 | 3200 | 2,29178644 |
| 20 | 1,15074799 | 240 | 2880 | 2,04369556 |
| 22 | 1,26582278 | 242 | 2662 | 1,82530324 |
| 24 | 1,38089758 | 245 | 2501,04167 | 1,63405543 |





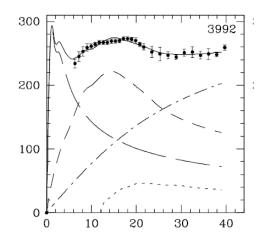


As it is shown in graphic, among galactic sample NGC 7331 is the only one which has a convex curve.

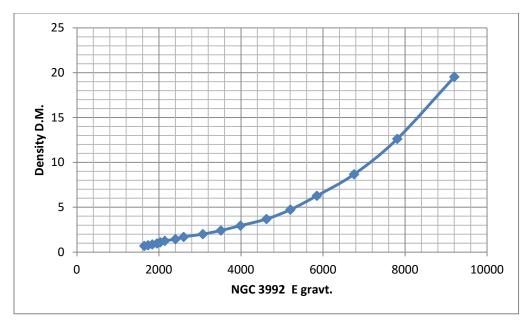
4.5 NGC 3992 [3] Bottema,R.B.

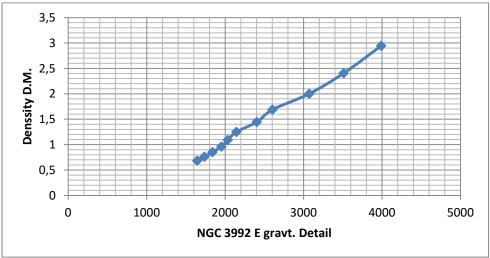
| 5 Name | Cluster | Distance | Author |
|----------|------------|----------|------------------|
| NGC 3992 | Ursa Major | 18,6 Mpc | [3] Bottema,R.B. |

| Dark matter density function profile Isothermal | | |
|--|--|--|
| $Rc = 3,89 \pm 0,34$ Kpc | | |
| Do = $66 \pm 11 \cdot 10^{-3} \text{ Msolar/pc}^3$ | | |



| Radii | Х | Speed | E gravt. | Density |
|-------|------------|-------|------------|------------|
| 38 | 9,76863753 | 250 | 1644,73684 | 0,68446071 |
| 36 | 9,25449871 | 250 | 1736,11111 | 0,76172233 |
| 34 | 8,7403599 | 250 | 1838,23529 | 0,85278048 |
| 32 | 8,22622108 | 250 | 1953,125 | 0,96110841 |
| 30 | 7,71208226 | 247 | 2033,63333 | 1,09133818 |
| 28 | 7,19794344 | 245 | 2143,75 | 1,24975408 |
| 26 | 6,68380463 | 250 | 2403,84615 | 1,44504734 |
| 24 | 6,16966581 | 250 | 2604,16667 | 1,68950155 |
| 22 | 5,65552699 | 260 | 3072,72727 | 2,00091038 |
| 20 | 5,14138817 | 265 | 3511,25 | 2,40578505 |
| 18 | 4,62724936 | 268 | 3990,22222 | 2,944925 |
| 16 | 4,11311054 | 272 | 4624 | 3,68351294 |
| 14 | 3,59897172 | 270 | 5207,14286 | 4,73030202 |
| 12 | 3,0848329 | 265 | 5852,08333 | 6,27603482 |
| 10 | 2,57069409 | 260 | 6760 | 8,67454515 |
| 8 | 2,05655527 | 250 | 7812,5 | 12,6209035 |
| 6 | 1,54241645 | 235 | 9204,16667 | 19,5321256 |



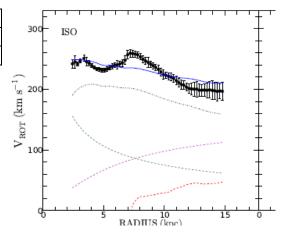


As it will be shown in paragraph 4. The graph of dark matter density has an anomalous behaviour for E=9000 because of data this galaxy. I have decided omit this galaxy in statistical study in paragraph 5.

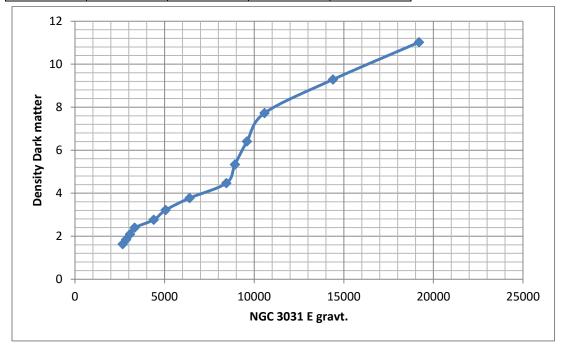
4.6 NGC 3031 Randriamampandry, T.

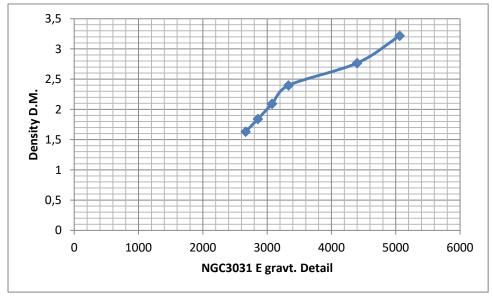
| 6 Name | Group / Constelation | Distance | Author |
|------------------|----------------------|----------|--------------------------|
| NGC 3031 or Bode | Group M81 | 2,5 Mpc | [2] Randriamampandry, T. |

| Dark matter density function profile Isothermal |
|---|
| Rc = 5.34 ± 1.97 Kpc |
| Do = 14, 55 \pm 5,87 10^{-3} Msolar/pc ³ |



| | I | | I | I |
|-------|------------|-------|------------|------------|
| Radii | X | Speed | E gravt. | Density |
| 15 | 2,80898876 | 200 | 2666,66667 | 1,63096946 |
| 14 | 2,62172285 | 200 | 2857,14286 | 1,84163684 |
| 13 | 2,43445693 | 200 | 3076,92308 | 2,09338503 |
| 12 | 2,24719101 | 200 | 3333,33333 | 2,39674673 |
| 11 | 2,05992509 | 220 | 4400 | 2,76543852 |
| 10 | 1,87265918 | 225 | 5062,5 | 3,21732303 |
| 9 | 1,68539326 | 240 | 6400 | 3,77550048 |
| 8 | 1,49812734 | 260 | 8450 | 4,46925924 |
| 7 | 1,31086142 | 250 | 8928,57143 | 5,33410307 |
| 6 | 1,12359551 | 240 | 9600 | 6,40893365 |
| 5 | 0,93632959 | 230 | 10580 | 7,72627421 |
| 4 | 0,74906367 | 240 | 14400 | 9,28834386 |
| 3 | 0,56179775 | 240 | 19200 | 11,0214471 |





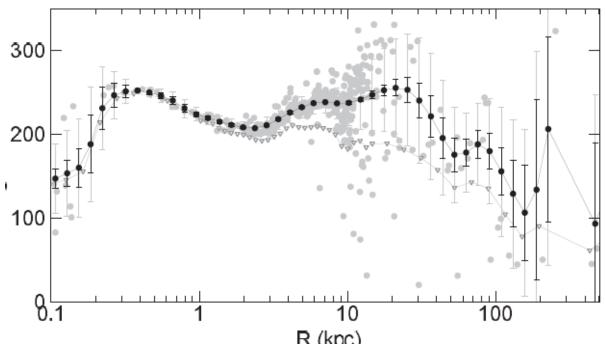
4.7 MILKY WAY [1] Sofue,Y.

| 7 Name | Group of galaxies | Stellar Radius | Author |
|-----------|-------------------|----------------|--------------------|
| Milky Way | Local group | 15 Kpc | [1] Sofue,Y. 2015. |

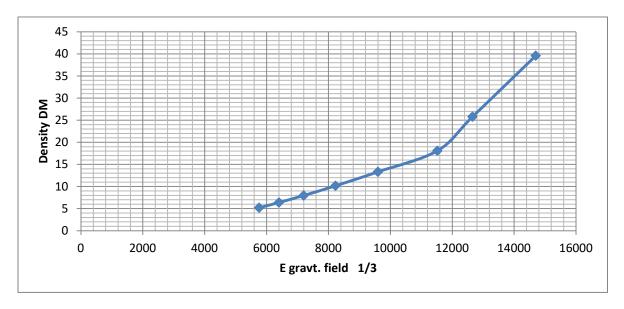
| Dark matter density function profile NFW |
|--|
| h = 10,7 ± 2,9 Kpc |
| Do = $18.2 \pm 7.4 \cdot 10^{-3}$ Msolar/pc ³ Error 40% iii |

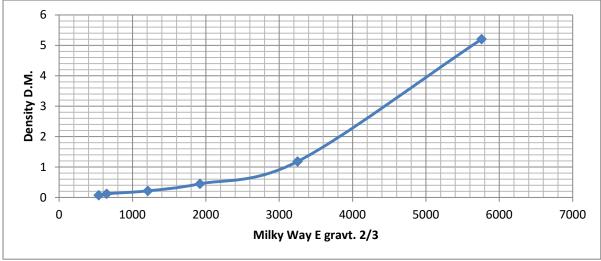
$$D_{NFW}(R) = \frac{D_0}{x \cdot (1+x)^2}$$
 where x=R/h

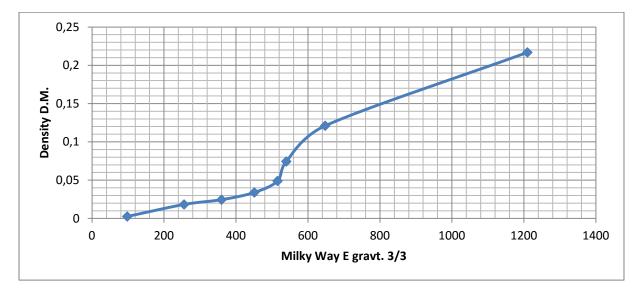
As it will be show in epigraph 5. Density of dark matter in Milky Way is outstandingly high. Considering that $Do = 18.2 \pm 7.4 \, 10^{-3} \, \text{Msolar/pc}^3$ is a parameter with error = 40%, perhaps this error so high might explain its atypical D.M. density calculated.



| Radii | Χ | Speed | E gravt. | Density |
|-------|------------|-------|------------|------------|
| 3 | 0,28037383 | 210 | 14700 | 39,5968221 |
| 4 | 0,37383178 | 225 | 12656,25 | 25,7945562 |
| 5 | 0,46728972 | 240 | 11520 | 18,0906184 |
| 6 | 0,56074766 | 240 | 9600 | 13,3241198 |
| 7 | 0,65420561 | 240 | 8228,57143 | 10,1666565 |
| 8 | 0,74766355 | 240 | 7200 | 7,9698385 |
| 9 | 0,8411215 | 240 | 6400 | 6,3833368 |
| 10 | 0,93457944 | 240 | 5760 | 5,20333791 |
| 20 | 1,86915888 | 255 | 3251,25 | 1,18281269 |
| 30 | 2,80373832 | 240 | 1920 | 0,44865514 |
| 40 | 3,73831776 | 220 | 1210 | 0,2168437 |
| 50 | 4,6728972 | 180 | 648 | 0,12102507 |
| 60 | 5,60747664 | 180 | 540 | 0,07434173 |
| 70 | 6,54205607 | 190 | 515,714286 | 0,04890774 |
| 80 | 7,47663551 | 190 | 451,25 | 0,03387803 |
| 90 | 8,41121495 | 180 | 360 | 0,02442988 |
| 100 | 9,34579439 | 160 | 256 | 0,01819397 |
| 200 | 18,6915888 | 140 | 98 | 0,0025111 |



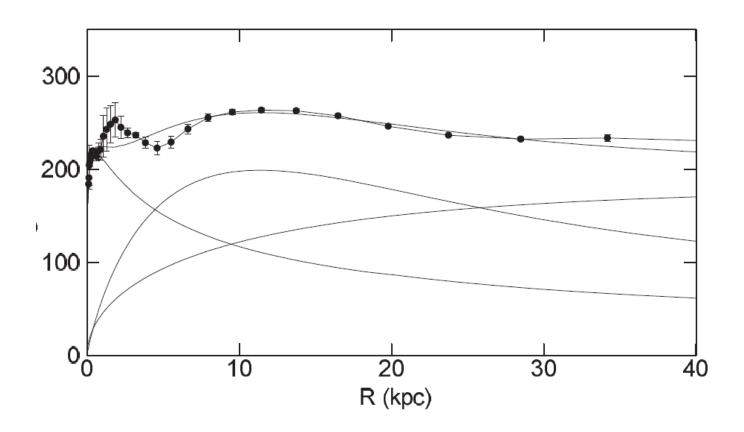


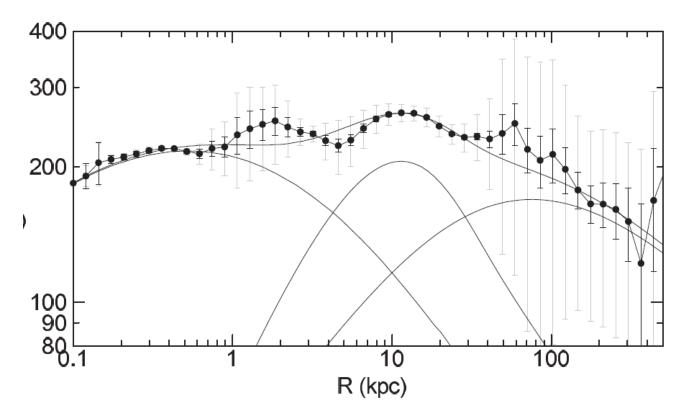


4.8 M31-ANDROMEDA [1] Sofue,Y.

| 8 Name | Group of galaxies | Distance | Author |
|--------|-------------------|----------|---------------------|
| M31 | Local group | 770 Kpc | [1] Sofue, Y. 2015. |

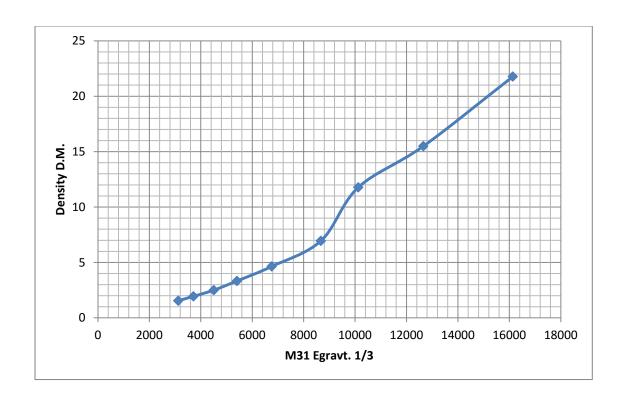
| Dark matter density function profile | NFW |
|--|-----|
| h = 34,6 ± 2,1 Kpc | |
| Do = $2,23 \pm 0,24 \cdot 10^{-3} \text{ Msolar/pc}^3$ | |

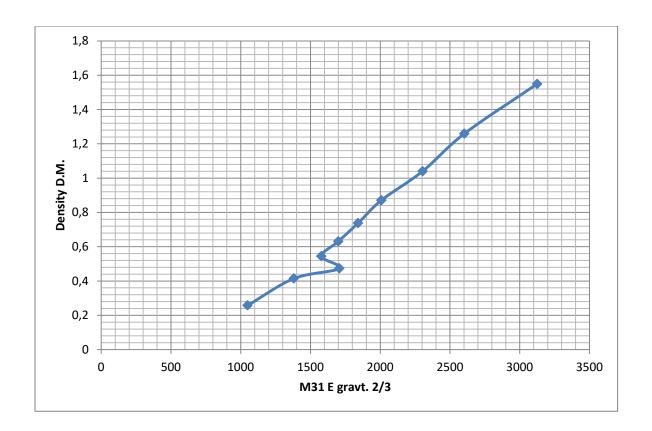


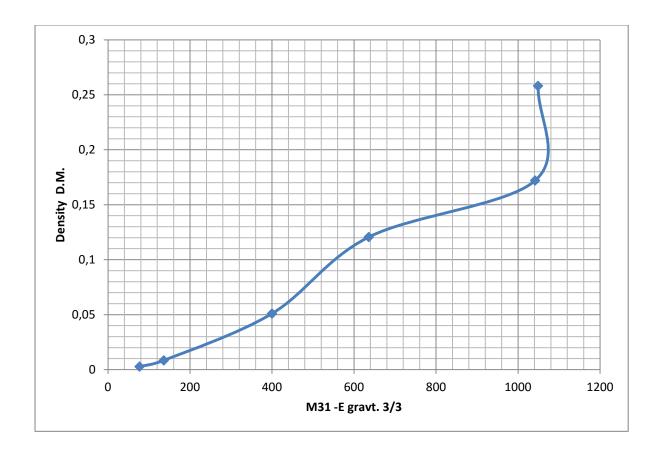


Rotation curve up 400 Kpc. Logarithmic scale for distances and speed.

| Radii | Х | Speed | E gravt. | Density |
|-------|------------|-------|------------|------------|
| 3 | 0,0867052 | 220 | 16133,3333 | 21,7789137 |
| 4 | 0,11560694 | 225 | 12656,25 | 15,4988173 |
| 5 | 0,14450867 | 225 | 10125 | 11,7807458 |
| 7,5 | 0,21676301 | 255 | 8670 | 6,94876628 |
| 10 | 0,28901734 | 260 | 6760 | 4,64369238 |
| 12,5 | 0,36127168 | 260 | 5408 | 3,33105138 |
| 15 | 0,43352601 | 260 | 4506,66667 | 2,503102 |
| 17,5 | 0,50578035 | 255 | 3715,71429 | 1,94455246 |
| 20 | 0,57803468 | 250 | 3125 | 1,54923706 |
| 22,5 | 0,65028902 | 242 | 2602,84444 | 1,25915277 |
| 25 | 0,72254335 | 240 | 2304 | 1,04016116 |
| 27,5 | 0,79479769 | 235 | 2008,18182 | 0,87099822 |
| 30 | 0,86705202 | 235 | 1840,83333 | 0,73781396 |
| 32,5 | 0,93930636 | 235 | 1699,23077 | 0,63125489 |
| 35 | 1,01156069 | 235 | 1577,85714 | 0,54481196 |
| 37,5 | 1,08381503 | 253 | 1706,90667 | 0,47383961 |
| 40 | 1,15606936 | 235 | 1380,625 | 0,41494976 |
| 50 | 1,44508671 | 229 | 1048,82 | 0,25812046 |
| 60 | 1,73410405 | 250 | 1041,66667 | 0,1720282 |
| 70 | 2,02312139 | 211 | 636,014286 | 0,12060679 |
| 100 | 2,89017341 | 200 | 400 | 0,05098507 |
| 200 | 5,78034682 | 165 | 136,125 | 0,00839165 |
| 300 | 8,67052023 | 152 | 77,0133333 | 0,00275017 |







5. STATISTICAL ANALYSIS OF EIGHT GALACTIC DATA SET - 8GDS -

As eight galactic data set it will be used frequently some times it will be called 8GDS.

Bellow are collected data from eight galactic set got in previous epigraph. Dominion extend from 1500 < E < 15000 because curve rotation data extend from inner region of bulge to outskirt region of galactic disks. Number beside each galaxy shows author who come from rotation curve data.

| | NGC | NGC | NGC | Milky | | NGC | NGC | NGC |
|----------|---------|---------|---------|--------------|--------|---------|---------|---------|
| E gravt. | 3031[2] | 7331[2] | 2841[3] | , Way [1] | M31[1] | 2841[2] | 7331[3] | 3992[3] |
| 1500 | | | 0,47 | 0,3 | 0,5 | 0,5 | 0,72 | 0,6 |
| 2000 | | | 0,85 | 0,5 | 0,85 | 0,86 | 1,1 | 1 |
| 3000 | 2 | 2,15 | 1,5 | 0,95 | 1,48 | 1,5 | 1,84 | 1,95 |
| 4000 | 2,6 | 2,75 | 2,4 | 2,3 | 2,2 | 2,4 | 2,2 | 2,95 |
| 5000 | 3,2 | 3,2 | 3,4 | 3,9 | 2,95 | 3,5 | 2,5 | 4,5 |
| 6000 | 3,6 | 3,5 | 4,3 | 5,5 | 3,9 | 4,5 | 2,7 | 6,5 |
| 7000 | 4 | 3,75 | 5,3 | 7,5 | 5 | 5,6 | 2,85 | 9,5 |
| 8000 | 4,3 | 3,85 | 6,4 | 9,5 | 6 | 7 | 2,93 | 13,5 |
| 9000 | 5,5 | 4 | 7,4 | 12 | 8 | 8 | 3,03 | 18,5 |
| 10000 | 7,2 | 4,1 | 8,6 | 14 | 11,5 | 9 | 3,1 | |
| 11000 | 8 | 4,17 | 9,8 | 16,5 | 13,3 | 10,6 | 3,15 | |
| 12000 | 8,4 | 4,25 | 11,2 | 20,5 | 14,5 | 12,6 | 3,2 | |
| 13000 | 8,8 | 4,333 | 12,6 | 28 | 16 | 14,5 | 3,24 | |
| 14000 | 9,2 | 4,42 | 14 | 35 | 18 | 16,1 | 3,28 | |
| 15000 | 9,6 | 4,5 | 15,4 | 41 | 19,5 | 17,9 | 3,32 | |

| | | Standard | Variation |
|----------|-------------------------|------------|-------------|
| E gravt. | $\overline{D}_{\it DM}$ | Deviation | coefficient |
| 1500 | 0,515 | 0,12776932 | 0,24809577 |
| 2000 | 0,86 | 0,18574176 | 0,21597879 |
| 3000 | 1,67125 | 0,36381443 | 0,21769001 |
| 4000 | 2,475 | 0,25248762 | 0,1020152 |
| 5000 | 3,39375 | 0,56592706 | 0,16675567 |
| 6000 | 4,3125 | 1,12742794 | 0,26143256 |
| 7000 | 5,4375 | 2,01877902 | 0,37126971 |
| 8000 | 6,685 | 3,2179885 | 0,4813745 |
| 9000 | 8,30375 | 4,64788645 | 0,55973343 |
| 10000 | 8,21428571 | 3,56868466 | 0,43444857 |
| 11000 | 9,36 | 4,40104208 | 0,4701968 |
| 12000 | 10,6642857 | 5,57383648 | 0,52266384 |
| 13000 | 12,4961429 | 7,77034091 | 0,62181915 |
| 14000 | 14,2857143 | 9,93243871 | 0,69527071 |
| 15000 | 15,8885714 | 11,8026963 | 0,74284188 |

Results shown above from statistical analysis of galactic data allow conclude quickly that Density of DM not depend on E, when gravitational field is increasing because standard deviation and variation coefficient are too high. Studying rotational curves I have concluded that Density of DM has too much error for radius under 10 Kpc which correspond E= 6000 aproximately. So I have decided to restrict dominion of E in order to get reliable results in statistical analysis.

6. FUNCTION FITTED TO EIGHT GALACTIC DATA SET - 8GDS - FOR E BELLOW 6000

It is clearly shown in previous figure of data cloud that point dispersion increase with E. I think, the reason why this happen is that the more near the bulge the more is dark matter negligible versus baryonic matter and error in calculus of dark matter density is bigger.

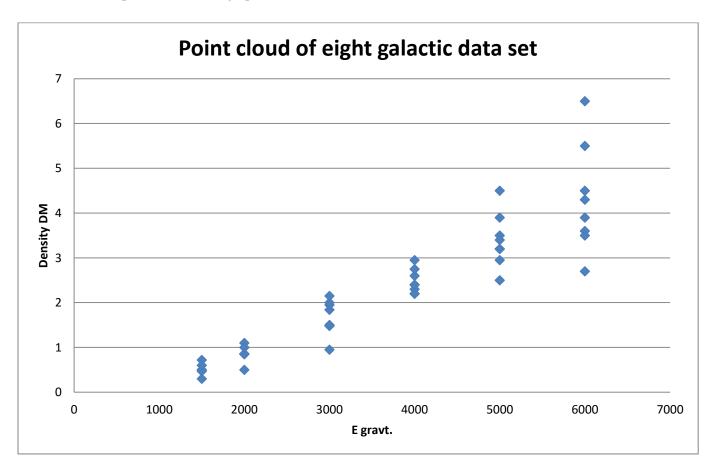
According the model proposed by the author, [4] Abarca,M.2014 dispersion of point cloud only is due to experimental error because the author defend a Universal law for Density of dark matter. In other words D.M. density does not depend on the galaxy but the E. Therefore point dispersion must decrease when contribution of dark matter increase because the more DM proportion is not negligible the less experimental error will be produced. That is precisely what happen in figures of cloud points.

I have decided to consider E< 6000 because point dispersion is a great deal lower that dispersion point at above table.

Looking for data table of galaxies studied, it is right to see that if radius is bigger than radius tabulated bellow, then E < 6000 for each galaxy.

| Galaxy | NGC 2841 | NGC 7331 | NGC 3031 | Milky Way | M31 | NGC 3992 |
|--------|------------|-----------|-----------|-----------|------------|----------|
| Radius | R > 15 Kpc | R >10 Kpc | R > 9 Kpc | R > 9 Kpc | R > 10 Kpc | R > 12 |

And looking for in original papers [1], [2], [3] it is possible to check that for R bigger than values showed in table above for each galaxy, contribution of dark matter is about 50%. So at bigger distances errors in calculus of dark matter decrease. Therefore point dispersion decreases when DM proportion increasing when it is considered bigger distances. As it is possible to see in graphic.



Bellow is table of data set when E < 6000

| | NGC | NGC | NGC | Milky | | NGC | NGC | NGC |
|----------|--|---------|---------|---------|--------|---------|---------|---------|
| | 3031[2] | 7331[2] | 2841[3] | Way [1] | M31[1] | 2841[2] | 7331[3] | 3992[3] |
| E gravt. | Density of Dark Matter in different galaxies | | | | | | | |
| 1500 | | | 0,47 | 0,3 | 0,5 | 0,5 | 0,72 | 0,6 |
| 2000 | | | 0,85 | 0,5 | 0,85 | 0,86 | 1,1 | 1 |
| 3000 | 2 | 2,15 | 1,5 | 0,95 | 1,48 | 1,5 | 1,84 | 1,95 |
| 4000 | 2,6 | 2,75 | 2,4 | 2,3 | 2,2 | 2,4 | 2,2 | 2,95 |
| 5000 | 3,2 | 3,2 | 3,4 | 3,9 | 2,95 | 3,5 | 2,5 | 4,5 |
| 6000 | 3,6 | 3,5 | 4,3 | 5,5 | 3,9 | 4,5 | 2,7 | 6,5 |

Table bellow show maximum radius of each galaxy measuring its rotation curve. At that distance E is the lowest value.

| Galaxies | N 3031 [2] | N 7331 [2] | N 2841 [3] | Milky Way | M31 [1] | N 2841 [2] | N 7331 [3] | N 3992 [3] |
|----------|------------|------------|------------|-----------|---------|------------|------------|------------|
| Lowest E | 2700 | 3000 | 1500 | 1500 | 1500 | 1500 | 1500 | 1500 |
| Radius | 15 Kpc | 25 Kpc | 50 Kpc | 35 Kpc | 35 Kpc | 50 Kpc | 36 Kpc | 40 Kpc |

Table bellow shows statistical parameters of dark matter density: average mean, standard deviation and variation coefficient.

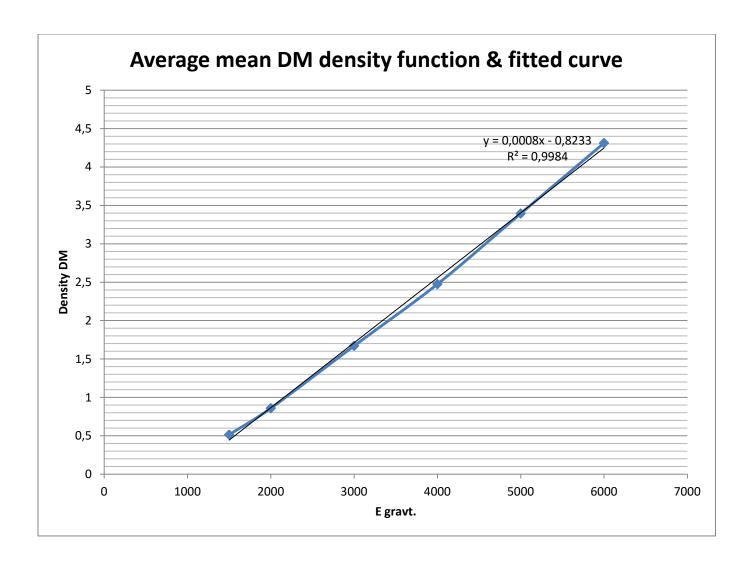
| Average | Standart | Variation | |
|---------|------------|-------------|---------|
| mean | Deviation | Coefficient | E gravt |
| 0,515 | 0,12776932 | 0,24809577 | 1500 |
| 0,86 | 0,18574176 | 0,21597879 | 2000 |
| 1,67125 | 0,34300753 | 0,20524011 | 3000 |
| 2,475 | 0,23804761 | 0,09618085 | 4000 |
| 3,39375 | 0,53356115 | 0,15721875 | 5000 |
| 4,3125 | 1,06294925 | 0,24648099 | 6000 |

It is remarkable that variation coefficient is always under 0,25.

6.1 LINEAR FIT TO AVERAGE MEAN OF DM DENSITY FUNCTION

In graph bellow is plotted average mean of DM Density depending on E.

| E | Avg mean | | |
|--------|------------|--|--|
| gravt. | DM density | | |
| 1500 | 0,515 | | |
| 2000 | 0,86 | | |
| 3000 | 1,6713 | | |
| 4000 | 2,475 | | |
| 5000 | 3,394 | | |
| 6000 | 4,312 | | |



| Linear function fitted to Average data of DM density \overline{D}_{DM} = 0,000845 · E -0,8236 and R=0.99919 | | | | | | | | |
|---|------|------|------|------|------|------|--|--|
| E gravt. | 1500 | 2000 | 3000 | 4000 | 5000 | 6000 | | |
| Den. | 0,44 | 0,87 | 1,71 | 2,56 | 3,40 | 4,25 | | |

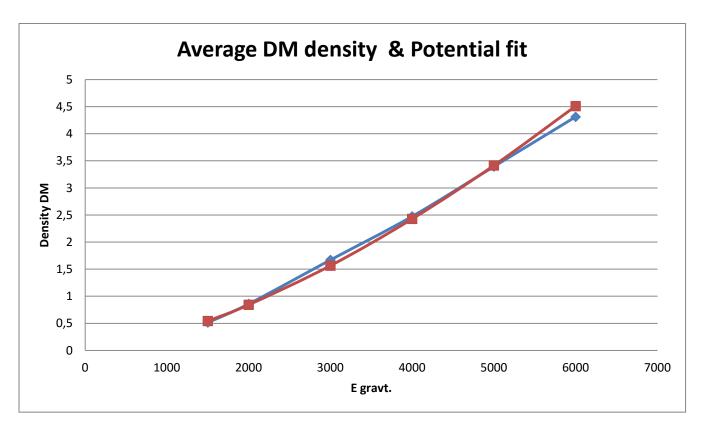
This linear fit function is only right inside dominion 1500 < E < 6000. For example this function become negative when E=974. It is needed to look for another more suitable fit function because it is necessary to wide dominion for E and DM density should go to cero when E goes to cero. Therefore potential fit is the most suitable function to fit DM density.

6.2 POTENTIAL FIT TO AVERAGE MEAN OF DM DENSITY FUNCTION

Physically it is right a potential fit because Density DM goes to 0 when E goes to 0, whereas linear fit above produces Density DM negative when E < 955 which has not physical meaning. Therefore a linear fit is only possible for reduced interval inside dominion.

Data table produces this potential fit function \overline{D}_{DM} =A·E^B where A= 7,686·10⁻⁶ and B= 1,5268 and R= 0,9985

| E | Avg mean | |
|--------|------------|---------------|
| gravt. | DM density | Potential fit |
| 1500 | 0,515 | 0,54319594 |
| 2000 | 0,86 | 0,84277759 |
| 3000 | 1,6713 | 1,56519738 |
| 4000 | 2,475 | 2,42842992 |
| 5000 | 3,394 | 3,41419076 |
| 6000 | 4,312 | 4,51005366 |

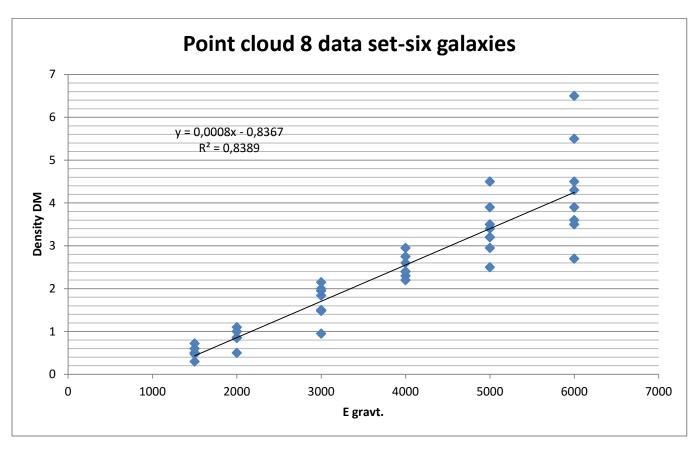


6.3 FUNCTIONAL FIT TO POINT CLOUD OF EIGHT GALACTIC DATA SET - 8GDS -

In this epigraph it will be fitted the point cloud by linear and potential function.

| | NGC | NGC | NGC | Milky | | NGC | NGC | NGC |
|----------|---------|---------|---------|---------|--------|---------|---------|---------|
| E gravt. | 3031[2] | 7331[2] | 2841[3] | Way [1] | M31[1] | 2841[2] | 7331[3] | 3992[3] |
| 1500 | | | 0,47 | 0,3 | 0,5 | 0,5 | 0,72 | 0,6 |
| 2000 | | | 0,85 | 0,5 | 0,85 | 0,86 | 1,1 | 1 |
| 3000 | 2 | 2,15 | 1,5 | 0,95 | 1,48 | 1,5 | 1,84 | 1,95 |
| 4000 | 2,6 | 2,75 | 2,4 | 2,3 | 2,2 | 2,4 | 2,2 | 2,95 |
| 5000 | 3,2 | 3,2 | 3,4 | 3,9 | 2,95 | 3,5 | 2,5 | 4,5 |
| 6000 | 3,6 | 3,5 | 4,3 | 5,5 | 3,9 | 4,5 | 2,7 | 6,5 |

6.3.1 LINEAR FIT

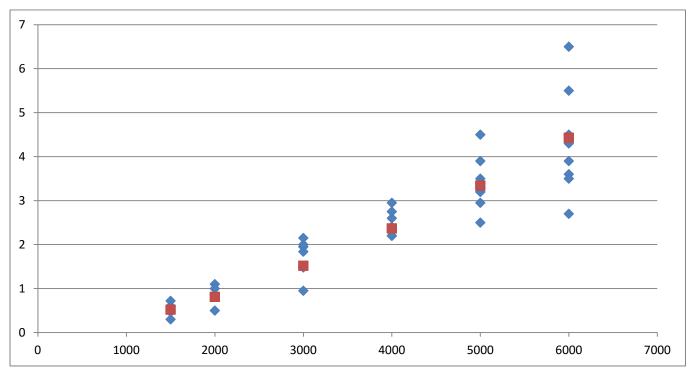


| Linear function fitted to DM density DM Density = 0,0008 · E -0,8367 and R=0,916 | | | | | | |
|--|--------|------|------|------|------|------|
| E gravt. | 1500 | 2000 | 3000 | 4000 | 5000 | 6000 |
| Den. | 0,3633 | 0,76 | 1,56 | 2,36 | 3,16 | 3,96 |

As you can see this fitted function is very similar to done with average density DM. \overline{D}_{DM} = 0,000845 · E -0,8236

6.3.2 POTENTIAL FIT

The point cloud is plotted in figure with diamond blue points and potential fitted results are plotted with square red points.



| DM Density from eight galactic data set DM Density = $6.5 \cdot 10^{-6} \cdot E^{B}$ where B= 1,544 and R= 0,9562 | | | | | | |
|---|--------------------------|------|------|------|------|------|
| E gravt. 1500 2000 3000 4000 5000 6000 | | | | | | |
| Density DM | 0,521 | 0,81 | 1,52 | 2,37 | 3,34 | 4,43 |
| | Dominion 1500 < E < 6000 | | | | | |

Discussion

Firstly it could be seen that correlation coefficient is 0,04 higher than linear fit coefficient. Secondly, this fitted function is very similar to done in epigraph 6.2 for \overline{D}_{DM} =A·E^B where A= 7,686·10⁻⁶ and B= 1,5268 and R= 0,9985.

Hyphotesis about big galaxies guarantee that gravitational field is not disturbed significantly by galactic neighbour. I think that this law must be true everywhere. The problem is that inside either intermediate or dwarf galaxies rotational curve can not show the real gravitational field because neighbour big galaxies disturb field. I will explain this idea again in epigraph 11 which is dedicate to intermediate galaxies.

In the following epigraph it will be studied Milky Way and M31 whose rotation curves extend beyond 200 Kpc. Thanks the remarkable paper [1] Sofue, Y.2015 it will be possible to check Density DM function inside a wide dominion of E. Particularly, dominion will be extend thus 100 < E < 6500.

7. FUNCTION FITTED TO GRAND ROTATION CURVE OF M31

| 8 Name | Group of galaxies | Distance |
|--------|-------------------|----------|
| M31 | Local group | 770 Kpc |

| Dark matter density function profile | NFW - [1] Sofue,Y. |
|--|--------------------|
| h = 34,6 ± 2,1 Kpc | |
| Do = $2,23 \pm 0,24 \cdot 10^{-3} \text{ Msolar/pc}^3$ | |

| | | 1 |
|------------|------------|-------|
| E gravt. | Density DM | Radii |
| 16133,3333 | 21,7789137 | 3 |
| 12656,25 | 15,4988173 | 4 |
| 10125 | 11,7807458 | 5 |
| 8670 | 6,94876628 | 7,5 |
| 6760 | 4,64369238 | 10 |
| 5408 | 3,33105138 | 12,5 |
| 4506,66667 | 2,503102 | 15 |
| 3715,71429 | 1,94455246 | 17,5 |
| 3125 | 1,54923706 | 20 |
| 2602,84444 | 1,25915277 | 22,5 |
| 2304 | 1,04016116 | 25 |
| 2008,18182 | 0,87099822 | 27,5 |
| 1840,83333 | 0,73781396 | 30 |
| 1699,23077 | 0,63125489 | 32,5 |
| 1577,85714 | 0,54481196 | 35 |
| 1706,90667 | 0,47383961 | 37,5 |
| 1380,625 | 0,41494976 | 40 |
| 1048,82 | 0,25812046 | 50 |
| 1041,66667 | 0,1720282 | 60 |
| 636,014286 | 0,12060679 | 70 |
| 400 | 0,05098507 | 100 |
| 136,125 | 0,00839165 | 200 |
| 77,0133333 | 0,00275017 | 300 |

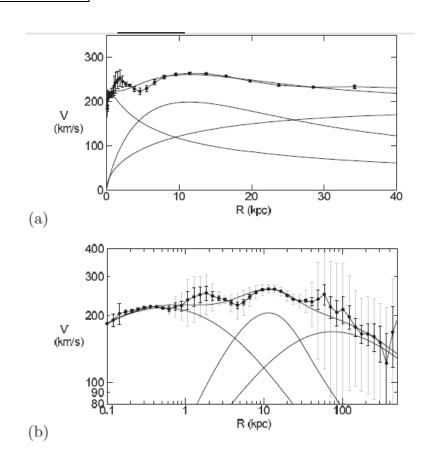


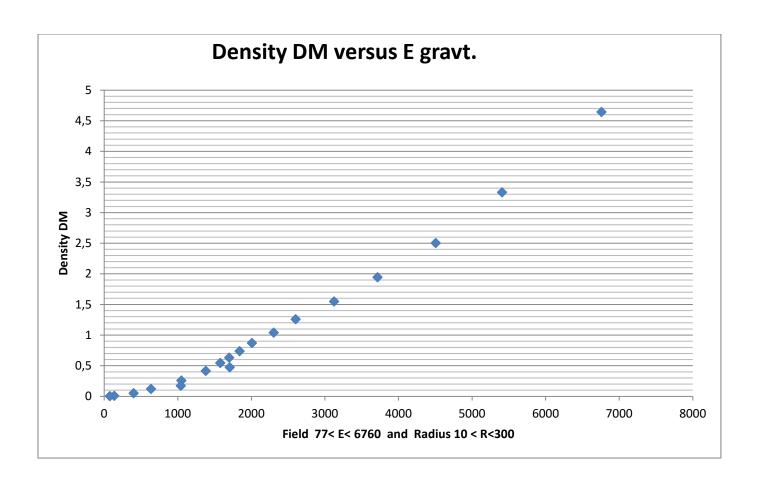
Fig. 5. Least- χ^2 fit of the GRC of M31 by the bulge, disk and dark halo components in (a) linear and (b) logarithmic scalings.

| Above | Grand Rotation Curve of M31 | [1] Sofue, Y. 2015. |
|-------|-----------------------------|---------------------|

Inside graphic (a) there are four curves: higher curve is the fitted to experimental data, the curve decreasing belong to bulge, the curve which has a maximum at 10 Kpc belong to disk and curve always increasing belong to dark matter.

(b) they are the same curves but in logarithmic scale.

Looking the graph at 10 Kpc it is right that dark matter curve is so high as bulge curve, they get 100 Km/s whereas disk curve gets almost 200 Km/s. Therefore it is possible conclude that for R=10 Kpc dark matter is a not negligible proportion of total matter. Looking for table E=6760 when E=10 Kpc. So it will be considered rotation curve for E=10 Kpc. This way error associated calculus of DM density will be reduced.

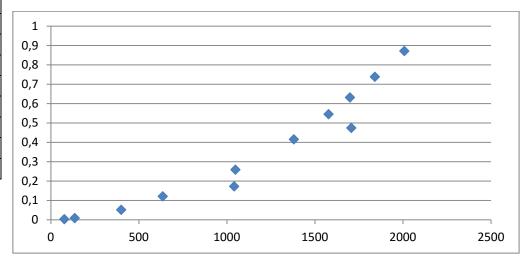


7.1 ANOMALOUS EXPERIMENTAL DATA

| E gravt. | Density DM |
|------------|------------|
| 2008,18182 | 0,87099822 |
| 1840,83333 | 0,73781396 |
| 1699,23077 | 0,63125489 |
| 1577,85714 | 0,54481196 |
| 1706,90667 | 0,47383961 |
| 1380,625 | 0,41494976 |
| 1048,82 | 0,25812046 |
| 1041,66667 | 0,1720282 |
| 636,014286 | 0,12060679 |
| 400 | 0,05098507 |
| 136,125 | 0,00839165 |
| 77,0133333 | 0,00275017 |

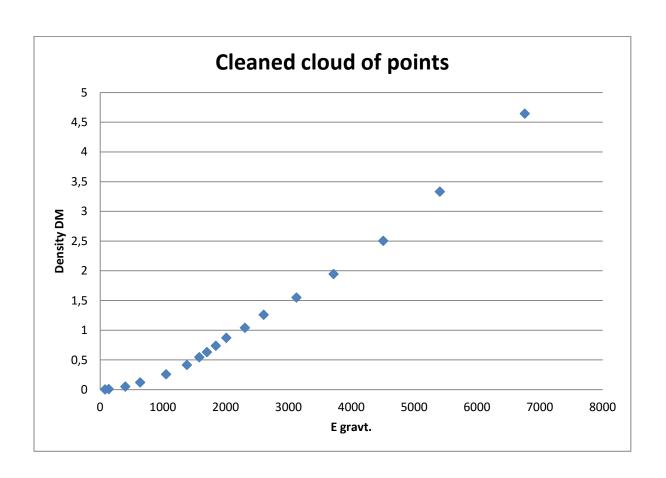
Looking graph it is obvious that points P(1041,8;0,1720) and P(1706,9;0,4738) are anomalous so it will be ignored to fit a curve over the cloud points.

This anomalous values may be originated either experimental error or local perturbations of gravitational field.



| "Cleaned cloud of Points" | | | | |
|---------------------------|------------|-------|--|--|
| | | | | |
| E grav. | Density DM | Radii | | |
| 6760 | 4,64369238 | 10 | | |
| 5408 | 3,33105138 | 12,5 | | |
| 4506,66667 | 2,503102 | 15 | | |
| 3715,71429 | 1,94455246 | 17,5 | | |
| 3125 | 1,54923706 | 20 | | |
| 2602,84444 | 1,25915277 | 22,5 | | |
| 2304 | 1,04016116 | 25 | | |
| 2008,18182 | 0,87099822 | 27,5 | | |
| 1840,83333 | 0,73781396 | 30 | | |
| 1699,23077 | 0,63125489 | 32,5 | | |
| 1577,85714 | 0,54481196 | 35 | | |
| 1380,625 | 0,41494976 | 40 | | |
| 1048,82 | 0,25812046 | 50 | | |
| 636,014286 | 0,12060679 | 70 | | |
| 400 | 0,05098507 | 100 | | |
| 136,125 | 0,00839165 | 200 | | |
| 77,0133333 | 0,00275017 | 300 | | |

As it is shown in graph this cloud of point allow fit a smooth curve with a high correlation coefficient as it will be seen.



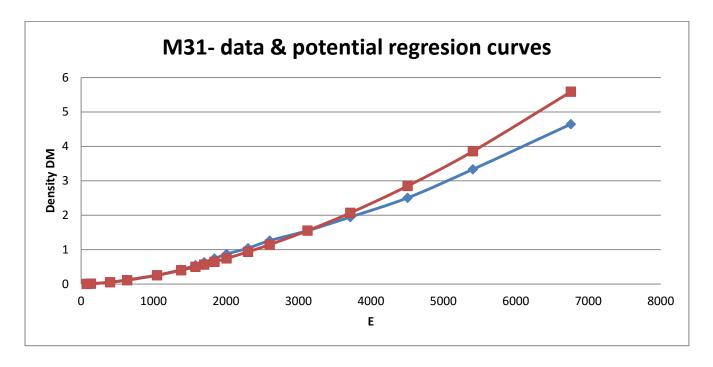
7.2 POTENTIAL FIT TO M31

| E gravt. | Density DM | Fitted function |
|------------|------------|-----------------|
| 6760 | 4,64369238 | 5,59E+00 |
| 5408 | 3,33105138 | 3,86E+00 |
| 4506,66667 | 2,503102 | 2,85E+00 |
| 3715,71429 | 1,94455246 | 2,07E+00 |
| 3125 | 1,54923706 | 1,55E+00 |
| 2602,84444 | 1,25915277 | 1,14E+00 |
| 2304 | 1,04016116 | 9,35E-01 |
| 2008,18182 | 0,87099822 | 7,44E-01 |
| 1840,83333 | 0,73781396 | 6,44E-01 |
| 1699,23077 | 0,63125489 | 5,64E-01 |
| 1577,85714 | 0,54481196 | 4,98E-01 |
| 1380,625 | 0,41494976 | 3,99E-01 |
| 1048,82 | 0,25812046 | 2,53E-01 |
| 636,014286 | 0,12060679 | 1,10E-01 |
| 400 | 0,05098507 | 5,10E-02 |
| 136,125 | 0,00839165 | 8,51E-03 |
| 77,0133333 | 0,00275017 | 3,31E-03 |

Graph above suggest that potential fit is suitable, in addition in epigraph 6.1 were pointed out a physics reason that justify potential fit.

Made statistical calculus through data from left table, it has been got the fitted function Density DM = $2,43\cdot10^{-6}\cdot E^B$ where B= 1,661 and R= 0,9984

Blue curve plot calculated data and red curve plot fitted function.



Bellow are tabulated a selected set of dominion which will be compared with Milky Way data in next epigraph.

| DM Density from M31 DM Density = $2,43 \cdot 10^{-6} \cdot E^{B}$ where B= $1,661$ and correlation coefficient R= $0,99847$ | | | | | | |
|---|---------------|-------|------|------|------|------|
| E gravt. | 100 | 500 | 1500 | 3000 | 4500 | 6400 |
| Den. | 0,0051 | 0,074 | 0,46 | 1,45 | 2,84 | 5,10 |
| Dominion | 77 < E < 6760 | | | | | |

As M31 and Milky Way are the biggest galaxies in Local group, according the theory I proposed in paper [4] Abarca,M.2014, it could be considered as a good approximation that radii of both galaxies are 385 Kpc. Which is half its distance. So 300 Kpc is a distance where field of M31 dominates fully over field of galactic neighbours. Therefore

gravitational field of M31 would be the only main responsible of DM generation at 300 Kpc of distance because the intergalactic field would be negligible.

8. FUNCTION FITTED TO GRAND ROTATION CURVE OF MILKY WAY

| 7 Name | Group of galaxies | Stellar Radius |
|-----------|-------------------|----------------|
| Milky Way | Local group | 15 Kpc |

| Dark matter density function prof | ile NFW [1] Sofue,Y. 2015. |
|---|----------------------------|
| $h = 10.7 \pm 2.9 \text{ Kpc}$ | |
| Do = $18,2 \pm 7,4 \cdot 10^{-3} \text{ Msolar/pc}^3$ | Error 40% iji |

Graphs bellow Extended Rotation curve of Milky Way [1] Sofue, Y. 2015.

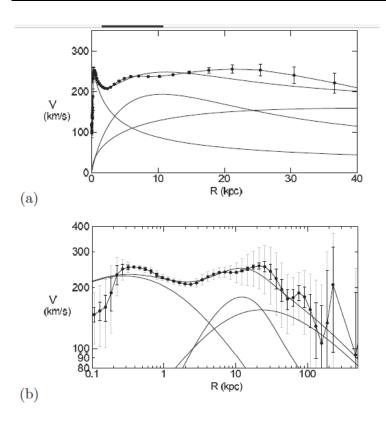
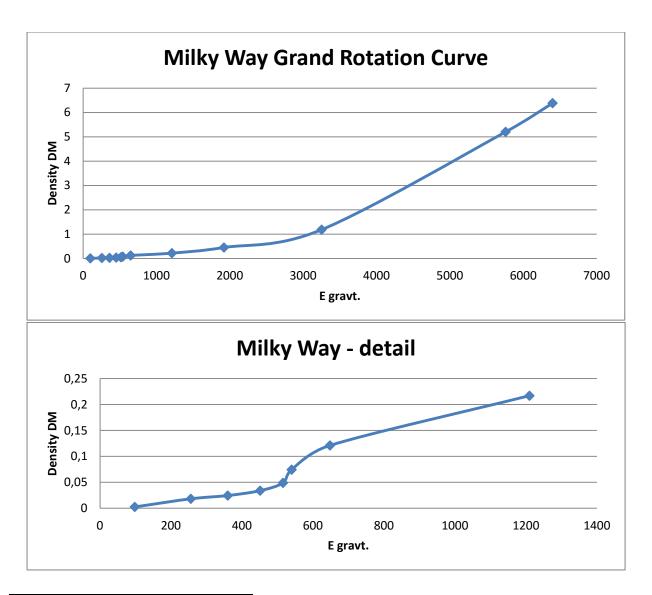


Fig. 6. Same as figure 5, but for the revised new GRC of Milky Way using the same procedure as for M31.

Inside graphic (a) there are four curves: higher curve is the fitted to experimental data, the curve decreasing belong to bulge, the curve which has a maximum at 10 Kpc belong to disk and curve always increasing belong to dark matter. (b) they are the same curves but in logarithmic scale.

Looking the graph at 10 Kpc it is right that dark matter curve is 150 % higher than bulge curve, it gets 125 Km/s whereas disk curve gets almost 200 Km/s. Therefore it is possible conclude that for R=10 Kpc dark matter is a not negligible proportion of total matter.

Looking for table it is got E=6400 when R=9 Kpc. So it will be considered rotation curve for R>9Kpc . This way error associated calculus of DM density will be reduced.



| E gravt. | Density | Radii Kpc |
|---------------------|-----------------------|-----------|
| 14700 | 39,5968221 | 3 |
| 12656,25 | 25,7945562 | 4 |
| 11520 | 18,0906184 | 5 |
| 9600 | 13,3241198 | 6 |
| 8228,57143 | 10,1666565 | 7 |
| 7200 | 7,9698385 | 8 |
| 6400 | 6,3833368 | 9 |
| 5760 | 5,20333791 | 10 |
| 3251,25 | 1,18281269 | 20 |
| 1920 | 0,44865514 | 30 |
| 1210 | 0,2168437 | 40 |
| 648 | 0,12102507 | 50 |
| 540 | 0,07434173 | 60 |
| 515,714286 | 0,04890774 | 70 |
| 451,25 | 0,03387803 | 80 |
| 360 | 0,02442988 | 90 |
| 256 | 0,01819397 | 100 |
| 98 | 0,0025111 | 200 |

Some values have been crossed out because they belong to inner region of galaxy where there is too much error in DM density.

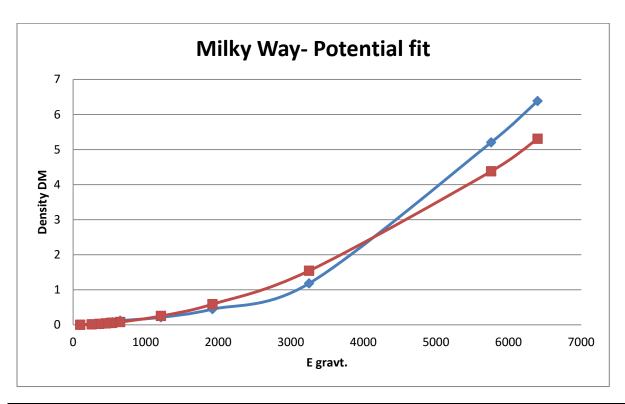
8.1 POTENTIAL FIT IN MILKY WAY

| E gravt. | Density DM | Fitted Function |
|------------|------------|-----------------|
| 6400 | 6,3833368 | 5,30664603 |
| 5760 | 5,20333791 | 4,37911057 |
| 3251,25 | 1,18281269 | 1,54349132 |
| 1920 | 0,44865514 | 0,59074903 |
| 1210 | 0,2168437 | 0,25455545 |
| 648 | 0,12102507 | 0,08151873 |
| 540 | 0,07434173 | 0,05846263 |
| 515,714286 | 0,04890774 | 0,05375742 |
| 451,25 | 0,03387803 | 0,04214013 |
| 360 | 0,02442988 | 0,02791217 |
| 256 | 0,01819397 | 0,01499051 |
| 98 | 0,0025111 | 0,00260275 |

As dominion is a great deal wider than galaxies in 8GDS it is not right to consider linear fit method so it will be got only potential fit like M31.

In addition, as was explained in epigraph 6.1, by physics reason it is suitable a DM density function that goes to cero when E goes to cero.

In graph bellow, blue curve represents DM density calculated and red curve represents potential fit.



| DM Density from Milky Way DM Density = 6,09·10 ⁻⁷ ·E ^B where B= 1,8234 and correlation coefficient R= 0,9953 | | | | | | |
|--|--------------------------------------|--------|-------|-------|------|------|
| E gravt. | E gravt. 100 500 1500 3000 4500 6400 | | | | | |
| Den. | 0,0027 | 0,0508 | 0,376 | 1,333 | 2,79 | 5,30 |
| Dominion 100 < E < 6400 | | | | | | |

It is clear that there is differences between results of fitted function of Milky Way and fitted function got from eight galactic data set which is written bellow.

| DM Density | DM Density from eight galactic data set - 8GDS DM. Density = $6.5 \cdot 10^{-6} \cdot E^{B}$ where B= 1,544 and R= 0,9562 | | | | | | | |
|--|---|------|------|------|------|------|------|------|
| E gravt. | 1500 | 2000 | 3000 | 4000 | 4500 | 5000 | 6000 | 6400 |
| Den. | 0,521 | 0,81 | 1,52 | 2,37 | 2,84 | 3,34 | 4,43 | 4,89 |
| Dominion $1500 < E < 6000$ Unit for $D_{D.M.}$ is $10^{-3} M_{\Theta}/pc^3$ and unit for E is $Km^2/s/Kpc$. | | | | | | | | |

There are several reason that could explain that difference.

Firstly [1] Sofue, Y. 2015. has shown in his paper Do = $18.2 \pm 7.4 \cdot 10^{-3}$ Msolar/pc³ which is a error of 40% iii

Other authors who study Milky Way state that measures of rotation curves inside Milky Way are particularly difficulty because they have to measure star speed from Earth, which orbits at 30 Km/s around the Sun which is a star whose vector velocity is not well known. Even scalar speed of Sun has an error about 10% that is 220 \pm 20 km/s.

This is the Milky Way paradox. Despite the fact we live inside Milky Way, errors in rotation curve measures are too bigj

9. COMPARISON BETWEEN D.M. DENSITY FITTED TO MILKY WAY TO M31 AND TO THE EIGHT GALACTIC DATA SET

9.1 COMPARISON BETWEEN D.M. DENSITY FITTED TO MILKY WAY AND TO M31

| DM Density from Milky Way DM Density = $6.09 \cdot 10^{-7} \cdot E^{B}$ where B= 1.8234 and correlation coefficient R= 0.9953 | | | | | | |
|---|--------------------------------------|--------|-------|-------|------|------|
| E gravt. | E gravt. 100 500 1500 3000 4500 6400 | | | | | |
| Den. | 0,0027 | 0,0508 | 0,376 | 1,333 | 2,79 | 5,30 |
| Dominion 100 < E < 6400 | | | | | | |

| DM Density from M31 DM Density = $2,43\cdot10^{-6}\cdot E^{B}$ where B= 1,661 and correlation coefficient R= 0,99847 | | | | | | |
|--|--------------------------------------|-------|------|------|------|------|
| E gravt. | E gravt. 100 500 1500 3000 4500 6400 | | | | | |
| Den. | 0,0051 | 0,074 | 0,46 | 1,45 | 2,84 | 5,10 |
| Dominion 77 < E < 6760 | | | | | | |

| E gravt. | Fit MW | Fit M31 | (M31-MW)*100/M31 |
|----------|------------|------------|------------------|
| 100 | 0,00270042 | 5,10E-03 | 4,71E+01 |
| 300 | 0,02001771 | 0,0316304 | 3,67E+01 |
| 600 | 0,07084554 | 0,1000266 | 2,92E+01 |
| 900 | 0,14838751 | 0,19615667 | 2,44E+01 |
| 1200 | 0,2507325 | 0,31631976 | 2,07E+01 |
| 1500 | 0,37663125 | 0,45824091 | 1,78E+01 |
| 2000 | 0,63639924 | 0,73895348 | 1,39E+01 |
| 3000 | 1,33295182 | 1,44912109 | 8,02E+00 |
| 4000 | 2,25230781 | 2,3368343 | 3,62E+00 |
| 5000 | 3,38324509 | 3,38528667 | 6,03E-02 |
| 6000 | 4,71750691 | 4,58263746 | -2,94E+00 |
| 6500 | 5,45880737 | 5,23426081 | -4,29E+00 |

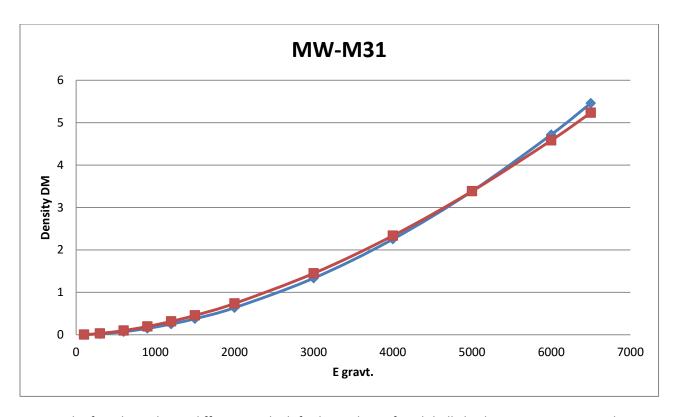
As table shows, relative difference between both curves increase when field decrease.

It is obvious that a 47 % of relative difference when E=100 is too high. However it is necessary to remember that NFW model in Milky Way has parameter Do = $18.2 \pm 7.4 \ 10^{-3} \ Msolar/pc^3$ which has an error of 40%.

In conclusion, perhaps differences between Milky Way and M31 might be explained by experimental error in calculus of Dark matter density.



Blue curve belong to Milky Way and Red curve belong to M31.



Despite the fact that relative difference is high for low values of E, globally both curves are very similar ii

9.2 COMPARISON BETWEEN DM DENSITY FITTED TO MILKY WAY TO M31 AND TO THE 8GDS

Bellow are shown the three functions fitted.

| DM Density from Milky Way DM Density = $6.09 \cdot 10^{-7} \cdot E^{B}$ where B= 1.8234 and correlation coefficient R= 0.9953 | | | | | | | | |
|---|---------------------------|--------|-------|-------|------|------|--|--|
| E gravt. | 100 | 500 | 1500 | 3000 | 4500 | 6400 | | |
| Den. | 0,0027 | 0,0508 | 0,376 | 1,333 | 2,79 | 5,30 | | |
| Dor | Dominion $100 < E < 6400$ | | | | | | | |

| DM Density function from M31 DM Density = $2,43 \cdot 10^{-6} \cdot E^{B}$ where B= $1,661$ and correlation coefficient R= $0,99847$ | | | | | | | | |
|--|------------------------------------|-------|------|------|------|------|--|--|
| E gravt. | gravt. 100 500 1500 3000 4500 6400 | | | | | | | |
| Den. | 0,0051 | 0,074 | 0,46 | 1,45 | 2,84 | 5,10 | | |
| Dominion | Dominion 77 < E < 6760 | | | | | | | |

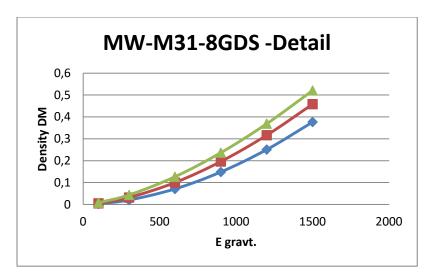
| DM Density function from eight galactic data set Density DM = $6.5 \cdot 10^{-6} \cdot E^{B}$ where B=1,544 and R= 0,9562 | | | | | | | |
|---|-------|------|------|------|------|------|--|
| E gravt. 1500 2000 3000 4000 5000 6000 | | | | | | | |
| Den. | 0,521 | 0,81 | 1,52 | 2,37 | 3,34 | 4,43 | |
| Dominion 1500 < E < 6000 | | | | | | | |

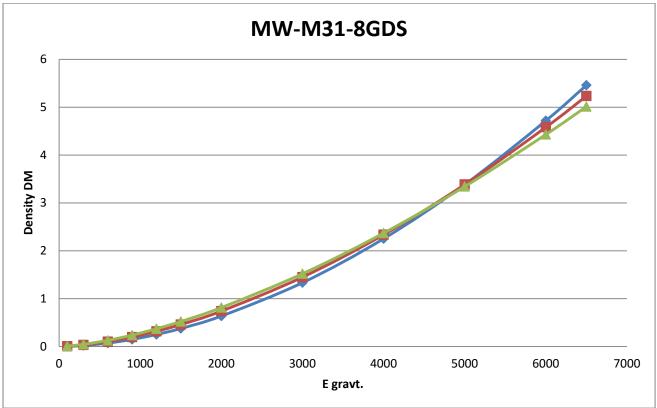
Although experimental data on eight galactic data set belong to dominion 1500 < E < 6000. I am going to extend his dominion to compare the three fitted function over dominion 100 < E < 6500 and cross out data of 8GDS for E< 1500 because these data have not experimental background.

| | | | | Rel. Diff. in % | Rel. Diff. in % | Rel. Diff. in % |
|----------|------------|------------|---------------------|-----------------|---------------------|---------------------|
| E gravt. | Fit MW | Fit M31 | Fit 8GDS | MW& M31 | MW& 8GDS | M31&8GDS |
| 100 | 0,00270042 | 5,10E-03 | 7,96E-03 | 4,71E+01 | 6,61E+01 | 3,59E+01 |
| 300 | 0,02001771 | 0,0316304 | 4,34E-02 | 3,67E+01 | 5,39E+01 | 2,71E+01 |
| 600 | 0,07084554 | 0,1000266 | 1,27E-01 | 2,92E+01 | 4,40E+01 | 2,10E+01 |
| 900 | 0,14838751 | 0,19615667 | 2,37E-01 | 2,44E+01 | 3,73E+01 | 1,71E+01 |
| 1200 | 0,2507325 | 0,31631976 | 3,69E-01 | 2,07E+01 | 3,21E+01 | 1,43E+01 |
| 1500 | 0,37663125 | 0,45824091 | 5,21E-01 | 1,78E+01 | 2,77E+01 | 1,20E+01 |
| 2000 | 0,63639924 | 0,73895348 | 8,12E-01 | 1,39E+01 | 2,17E+01 | 9,03E+00 |
| 3000 | 1,33295182 | 1,44912109 | 1,52E+00 | 8,02E+00 | 1,23E+01 | 4,61E+00 |
| 4000 | 2,25230781 | 2,3368343 | 2,37E+00 | 3,62E+00 | 4,91E+00 | 1,34E+00 |
| 5000 | 3,38324509 | 3,38528667 | 3,34E+00 | 6,03E-02 | -1,21E+00 | -1,27E+00 |
| 6000 | 4,71750691 | 4,58263746 | 4,43E+00 | -2,94E+00 | -6,50E+00 | -3,45E+00 |
| 6500 | 5,45880737 | 5,23426081 | 5,01E+00 | -4,29E+00 | -8,90E+00 | -4,42E+00 |

Discussion of results

As table shows, relative differences increase when field decrease. It is important to remark that relative differences between MW and M31 or 8GDS are bigger that ones between M31 and 8GDS. I think that this fact might be explained by experimental error in calculus of DM inside Milky Way. Remember that Do = $18,2\pm7,4$ 10^{-3} Msolar/pc³ so error is 40%. Also it is important to remark that inside dominion 1500 < E < 6500 relative difference between M31 and 8GDS are bellow 12 % which is even lower that error of parameters of DM model calculated.





Blue curve belong to Milky Way. Red curve belong to M31. Green curve belong to 8GDS.

Globally it is possible to state that three curve are very similar.

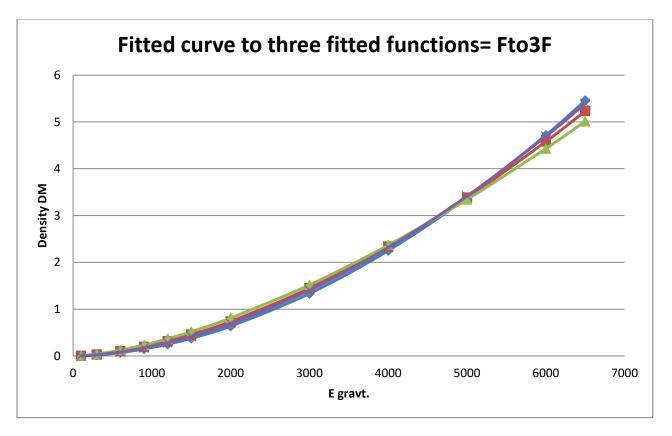
10.DARK M. DENSITY FUNCTION DEPENDING ON E AS UNIVERSAL LAW FOR BIG GALAXIES

Differences found between curves fitted to Milky Way, M31 and 8GDS are enough small to think that it could be caused by experimental errors in galactic rotation curves and dark matter density function depend on E according a Universal law. It is known that experimental errors in rotation curves of galaxies are important because of difficult measures.

According data collected in this paper, the simplest way to calculate an average curve would be to fit a function to fitted functions calculated to Milky Way, to M31 and to 8GDS. Bellow is shown the function found which abbreviate is called Fto3F. Fto3F means Fit curve which fit to 3 fitted curves.

| Funct | ion Fto3F | Density _{DA} | $_{M} = 1,257 \cdot 10^{-6} \cdot E^{B}$ | where B= 1,739 | and R= 0,9975 | | | |
|-------------------------|-----------|-----------------------|--|----------------|---------------|--------|--|--|
| E gravt. | 100 | 500 | 1500 | 3000 | 4500 | 6500 | | |
| Den. | 0,003787 | 0,06226 | 0,42093 | 1,4056 | 2,8456 | 5,3948 | | |
| Dominion 100 < E < 6500 | | | | | | | | |

| | | | | Fit to 3 Fit = |
|----------|------------|------------|----------|----------------|
| E gravt. | Fit MW | Fit M31 | Fit 8GDS | Fto3F |
| 100 | 0,00270042 | 5,10E-03 | 7,96E-03 | 3,79E-03 |
| 300 | 0,02001771 | 0,0316304 | 4,34E-02 | 2,56E-02 |
| 600 | 0,07084554 | 0,1000266 | 1,27E-01 | 8,55E-02 |
| 900 | 0,14838751 | 0,19615667 | 2,37E-01 | 1,73E-01 |
| 1200 | 0,2507325 | 0,31631976 | 3,69E-01 | 2,86E-01 |
| 1500 | 0,37663125 | 0,45824091 | 5,21E-01 | 4,21E-01 |
| 2000 | 0,63639924 | 0,73895348 | 8,12E-01 | 6,94E-01 |
| 3000 | 1,33295182 | 1,44912109 | 1,52E+00 | 1,41E+00 |
| 4000 | 2,25230781 | 2,3368343 | 2,37E+00 | 2,32E+00 |
| 5000 | 3,38324509 | 3,38528667 | 3,34E+00 | 3,42E+00 |
| 6000 | 4,71750691 | 4,58263746 | 4,43E+00 | 4,69E+00 |
| 6500 | 5,45880737 | 5,23426081 | 5,01E+00 | 5,39E+00 |



Blue curve belong to Milky Way. Red curve belong to M31. Green curve belong to 8GDS.

Magenta curve belong to fitted function which fit the fitted functions to Milky Way, to M31 and to 8GDS.

As reader can check it is not possible to differentiate magenta curve of blue curve because both are very similar.

Differences relatives between curves of MW, M31, 8GDS and Fto3F are tabulated bellow in order to check it.

| Relat | Relative Differences between fit curves to MW, M31, 8GDS and Fto3F which is the "average curve" | | | | | | | | | | |
|--------|---|------------|---------------------|--------------|--------------|--------------|----------------------|--|--|--|--|
| E | | | | | Rel. Diff. % | Rel. Diff. % | Rel. Diff. % | | | | |
| gravt. | Fit MW | Fit M31 | Fit 8GDS | Fit to 3 Fit | MW & Fto3F | M31 & Fto3F | 8GDS & Fto3F | | | | |
| 100 | 0,00270042 | 5,10E-03 | 7,96E-03 | 3,79E-03 | 2,87E+01 | -3,47E+01 | -1,10E+02 | | | | |
| 300 | 0,02001771 | 0,0316304 | 4,34E-02 | 2,56E-02 | 2,18E+01 | -2,35E+01 | -6,95E+01 | | | | |
| 600 | 0,07084554 | 0,1000266 | 1,27E-01 | 8,55E-02 | 1,71E+01 | -1,70E+01 | -4,80E+01 | | | | |
| 900 | 0,14838751 | 0,19615667 | 2,37E-01 | 1,73E-01 | 1,43E+01 | -1,33E+01 | -3,68E+01 | | | | |
| 1200 | 0,2507325 | 0,31631976 | 3,69E-01 | 2,86E-01 | 1,22E+01 | -1,08E+01 | -2,93E+01 | | | | |
| 1500 | 0,37663125 | 0,45824091 | 5,21E-01 | 4,21E-01 | 1,05E+01 | -8,86E+00 | -2,38E+01 | | | | |
| 2000 | 0,63639924 | 0,73895348 | 8,12E-01 | 6,94E-01 | 8,34E+00 | -6,43E+00 | -1,70E+01 | | | | |
| 3000 | 1,33295182 | 1,44912109 | 1,52E+00 | 1,41E+00 | 5,17E+00 | -3,10E+00 | -8,08E+00 | | | | |
| 4000 | 2,25230781 | 2,3368343 | 2,37E+00 | 2,32E+00 | 2,85E+00 | -7,94E-01 | -2,16E+00 | | | | |
| 5000 | 3,38324509 | 3,38528667 | 3,34E+00 | 3,42E+00 | 1,02E+00 | 9,57E-01 | 2,20E+00 | | | | |
| 6000 | 4,71750691 | 4,58263746 | 4,43E+00 | 4,69E+00 | -5,08E-01 | 2,37E+00 | 5,62E+00 | | | | |
| 6500 | 5,45880737 | 5,23426081 | 5,01E+00 | 5,39E+00 | -1,19E+00 | 2,98E+00 | 7,09E+00 | | | | |

Discussion of data table

In the last and fourth columns some data have been crossed out because curve of 8GDS were obtained inside dominion 1500 < E < 6000 so data E < 1500 have not experimental background.

Except data belong to E=100 and E=300 the relative differences are bellow 17 % and relative differences inside 1500 < E < 6500 interval are mainly bellow 10%.

Results suggest to think that dark matter density depending on E could be a Universal law because relative differences between functions fitted to Milky Way, M31 and 8GDS are similar to experimental error to measure the galactic rotation curves. Therefore relative difference might be explained by experimental error in measures rotation curves.

Bellow have been collected DM density function, its characteristic parameters and its calculated errors.

| Parameters of DM density function fitted by the authors quoted | | | | | | |
|--|--|--|--|--|--|--|
| | | | | | | |
| 1 - NGC 2841 - Bottema | 5 - NGC 3992 - Bottema | | | | | |
| Dark matter density function profile Isothermal | Dark matter density function profile Isothermal | | | | | |
| Rc = $6,39 \pm 0,52$ Kpc | Rc = 3.89 ± 0.34 Kpc | | | | | |
| Do = 32 ± 5 10 ⁻³ Msolar/pc ³ | Do = $66 \pm 11 \cdot 10^{-3} \text{ Msolar/pc}^3$ | | | | | |
| 2 - NGC 2841 - Randriamampandry Dark matter density function profile Isothermal | 6 - NGC 3031 - Randriamampandry | | | | | |
| Rc = 5.08 ± 0.23 Kpc | Dark matter density function profile Isothermal | | | | | |
| Do = $49,06 \pm 3,61 \cdot 10^{-3} \text{ Msolar/pc}^3$ | Rc = 5,34 \pm 1,97 Kpc Do = 14,55 \pm 5,87 10^{-3} Msolar/pc ³ | | | | | |
| 3 - NGC 7331 - Bottema | 7 -Milky Way - Sofue. | | | | | |
| Dark matter density function profile Isothermal | Dark matter density function profile NFW | | | | | |
| Rc = $19,46 \pm 1,79$ Kpc | h = 10,7 ± 2,9 Kpc | | | | | |
| Do = $3.5 \pm 0.7 \cdot 10^{-3}$ Msolar/pc ³ | Do = $18.2 \pm 7.4 \cdot 10^{-3} \text{ Msolar/pc}^3$ Error 40% iii | | | | | |
| 4 - NGC 7331 - Randriamampandry | 8 - M31 - Sofue | | | | | |
| Dark matter density function profile Isothermal | Dark matter density function profile NFW | | | | | |
| Rc = 17,38 ± 2,75 Kpc | h = 34,6 ± 2,1 Kpc | | | | | |
| Do = $4,75 \pm 0.6 10^{-3} \text{Msolar/pc}^3$ | $Do = 2,23 \pm 0,24 \ 10^{-3} \ Msolar/pc^3$ | | | | | |

As it is right to see, most of the times parameters Do has 15% of error or even higher. Except Milky Way whose Do has 40% of error;

In addition, another way to explain differences in DM density function in different galaxies might be that galaxies does not verify symmetry hypothesis rigorously so Virial theorem cannot calculate real gravitational field with total accuracy.

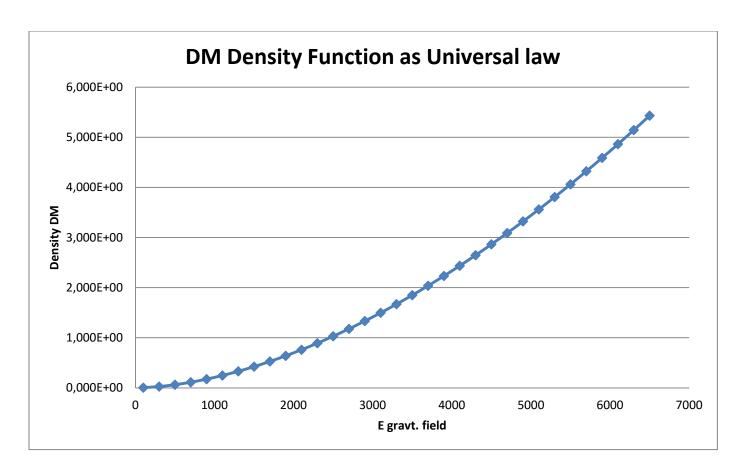
Another way to explain those difference might be that according my previous paper [4] Abarca,M.2014, dark matter is a quantum gravity effect so a classical treatment only could aspire to be a good approximation of a deeper quantum gravity theory.

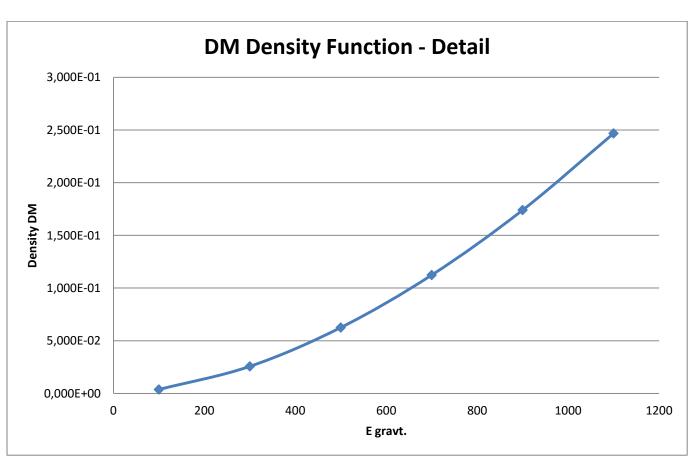
Taking in consideration this ideas I dare to postulate dark matter density function depending on E as universal law.

At this point, it is obvious that candidate to function as Universal law should be the "average" curve between MW, M31 and 8GDS.

| | Dark matter density function as Universal law for big galaxies | | | | | | | | | |
|----------|--|-----|------|------|------|------|--|--|--|--|
| | Density $_{DM} = 1,26 \cdot 10^{-6} \cdot E^{B}$ where B= 1,74 | | | | | | | | | |
| | Where Unit for $D_{D.M.}$ is $10^{-3} M_{\odot}/pc^3$ and Unit for E is $Km^2/s/Kpc$ | | | | | | | | | |
| E gravt. | 100 | 500 | 1500 | 3100 | 4500 | 6500 | | | | |
| Den. | | | | | | | | | | |
| | Dominion 100 < E < 6500 | | | | | | | | | |

| DM Density Function as Universal Law | | | | | | | |
|---|------------|--|--|--|--|--|--|
| E gravt. | DM Density | | | | | | |
| 100 | 3,81E-03 | | | | | | |
| 300 | 2,57E-02 | | | | | | |
| 500 | 6,26E-02 | | | | | | |
| 700 | 1,12E-01 | | | | | | |
| 900 | 1,74E-01 | | | | | | |
| 1100 | 2,47E-01 | | | | | | |
| 1300 | 3,30E-01 | | | | | | |
| 1500 | 4,23E-01 | | | | | | |
| 1700 | 5,26E-01 | | | | | | |
| 1900 | 6,39E-01 | | | | | | |
| 2100 | 7,60E-01 | | | | | | |
| 2300 | 8,91E-01 | | | | | | |
| 2500 | 1,03E+00 | | | | | | |
| 2700 | 1,18E+00 | | | | | | |
| 2900 | 1,33E+00 | | | | | | |
| 3100 | 1,50E+00 | | | | | | |
| 3300 | 1,67E+00 | | | | | | |
| 3500 | 1,85E+00 | | | | | | |
| 3700 | 2,04E+00 | | | | | | |
| 3900 | 2,23E+00 | | | | | | |
| 4100 | 2,44E+00 | | | | | | |
| 4300 | 2,65E+00 | | | | | | |
| 4500 | 2,86E+00 | | | | | | |
| 4700 | 3,09E+00 | | | | | | |
| 4900 | 3,32E+00 | | | | | | |
| 5100 | 3,56E+00 | | | | | | |
| 5300 | 3,81E+00 | | | | | | |
| 5500 | 4,06E+00 | | | | | | |
| 5700 | 4,32E+00 | | | | | | |
| 5900 | 4,59E+00 | | | | | | |
| 6100 | 4,86E+00 | | | | | | |
| 6300 | 5,14E+00 | | | | | | |
| 6500 | 5,43E+00 | | | | | | |





11. EXPRESSION OF DM DENSITY FUNCTION THROUGH INTERNATIONAL SYSTEM OF UNITS AND ANOTHER UNITS

In epigraph 10 has been got Density D.M. Density function using a particular system of units. In this epigraph it will be got the same function expressing its magnitudes through International System of Units.

| Dark matter | Dark matter density function as Universal law for big galaxies – Paper Units | | | | | | | | | |
|-------------|---|-----|------|------|------|------|--|--|--|--|
| Dens | Density $_{DM} = 1,26 \cdot 10^{-6} \cdot E^{B}$ where $B = 1,74$ | | | | | | | | | |
| | Where Unit for $D_{D.M.}$ is $10^{-3} M_{\Theta}/pc^3$ and Unit for E is $Km^2/s^2/Kpc$ | | | | | | | | | |
| E gravt. | 100 | 500 | 1500 | 3000 | 4500 | 6500 | | | | |
| Den. | Den. 0,003787 0,06226 0,42093 1,4056 2,8456 5,3948 | | | | | | | | | |
| | Dominion 100 < E < 6500 | | | | | | | | | |

11.1 CONVERSION OF PAPER UNITS INTO INTERNATIONAL SYSTEM

Astronomical data $M_{\Theta} = 1,989 \cdot 10^{30} \text{ Kg}$ $1 \text{ pc} = 3,086 \cdot 10^{16} \text{ m} \rightarrow 1 \text{ Kpc} = 3,086 \cdot 10^{19} \text{ m}$

In this paper unit for $D_{D.M.}$ is $10^{-3} M_{\odot}/pc^3 = 6,768 \cdot 10^{-23} \text{ Kg/m}^3$ using astronomical data above.

Unit for E is $Km^2/s^2/Kpc = 3,24 \cdot 10^{-14}$ m/s² using astronomical data above.

As a consequence Dominion go from 100 < E < 6500 (paper Units) to $3.24 \cdot 10^{-12} < E < 2.11 \cdot 10^{-10}$ (I.S.)

11.2 DM DENSITY FUNCTION THROUGH INTERNATIONAL SYSTEM OF UNITS

Density $_{DM} = A_{PAPER\ UNITS} \cdot E^B$ where $A_{PAPER\ UNITS} = 1,26 \cdot 10^{-6}$ and B = 1,74 is adimensional

Density DM (I.S. of Units) =
$$A_{I.S.} \cdot E^B$$
 where $A_{I.S.} = \frac{1,26 \cdot 10^{-6} \cdot 6,768 \cdot 10^{-23}}{\left(3,24 \cdot 10^{-14}\right)^{1.74}} = 2,526 \cdot 10^{-5}$ is constant A in I.S.

of units.

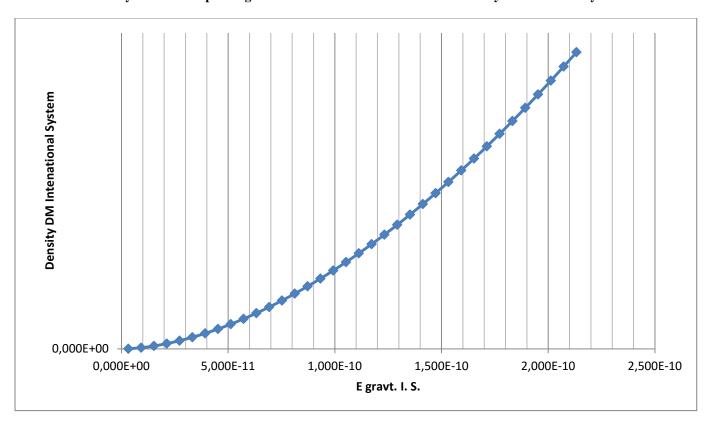
So $A_{LS.}$ is approximately 20 times bigger than A in Units of paper, exactly $A_{LS.} = 20,05 \cdot A_{PAPER\;UNITS}$

| Dark matter den | Dark matter density function as Universal law for big galaxies – International System of Units | | | | | | | | | |
|--|--|-----------------------|------------|-----------------------|------------------------|-------------------------|--|--|--|--|
| Density $_{DM} = 2,526 \cdot 10^{-5} \cdot E^{B}$ where $B = 1,74$ | | | | | | | | | | |
| | Where Unit for Density $D_{D.M.}$ is Kg/m^3 and Unit for E is m/s^2 | | | | | | | | | |
| E paper Units | 100 | 500 | 1500 | 3000 | 4500 | 6500 | | | | |
| E gravt. I.S. | $3,24\cdot10^{-12}$ | $1,62 \cdot 10^{-11}$ | 4,86.10-11 | $9,72 \cdot 10^{-11}$ | $1,458 \cdot 10^{-10}$ | 2,106·10 ⁻¹⁰ | | | | |
| Density _{DM} I.S. | Density _{DM} I.S. $2,58 \cdot 10^{-25}$ $4,24 \cdot 10^{-24}$ $2,86 \cdot 10^{-23}$ $9,57 \cdot 10^{-23}$ $1,94 \cdot 10^{-22}$ $3,67 \cdot 10^{-22}$ | | | | | | | | | |
| | Dominion $3,24 \cdot 10^{-12} < E < 2,11 \cdot 10^{-10}$ | | | | | | | | | |

It is clear that International System Units are not suitable to work with because measures are too tiny.

For example Excel cannot support these numbers so tiny. In graph bellow has been plotted DM Density function above in its dominion $3.24 \cdot 10^{-12} < E < 2.11 \cdot 10^{-10}$, according units in I.S.

Dark Matter Density Function Depending on Gravitational Field as Universal Law by International System Units.



However I.S. Units are needed to check results this paper with results which come from other sources.

In table bellow I have inserted a raw which shows equivalence of DM density in numbers of electrons/m³.

| Dark matter density function as Universal law for big galaxies – International System of Units | | | | | | | |
|--|-----------------------|------------------------|------------------|-----------------------|--------------------------|-------------------------|--|
| Density $_{DM} = 2,526 \cdot 10^{-5} \cdot E^{B}$ where $B = 1,74$ | | | | | | | |
| Where Unit for Density $D_{D.M.}$ is Kg / m ³ and Unit for E is m/s ² | | | | | | | |
| E paper Unit | 100 | 500 | 1500 | 3000 | 4500 | 6500 | |
| E gravt. I.S. | $3,24\cdot10^{-12}$ | 1,62·10 ⁻¹¹ | 4,86.10-11 | | $1,458 \cdot 10^{-10}$ | 2,106·10 ⁻¹⁰ | |
| Density _{DM} I.S. | $2,58 \cdot 10^{-25}$ | $4,24 \cdot 10^{-24}$ | 2,86.10-23 | $9,57 \cdot 10^{-23}$ | 1,94 · 10 ⁻²² | 3,67·10 ⁻²² | |
| Nº electrons/m ³ | $2.8 \cdot 10^5$ | $4,65 \cdot 10^6$ | $3,1 \cdot 10^7$ | $1.05 \cdot 10^8$ | $2,1 \cdot 10^{8}$ | 4.10^{8} | |
| Dominion $3.24 \cdot 10^{-12} < E < 2.11 \cdot 10^{-10}$ | | | | | | | |

11.3 EXPRESSION OF DM DENSITY FUNCTION IN GeV/cm³

As GeV/cm³ is a typical unit for Density of D.M. I have made this conversion of units. Reader can check this conversion knowing that $10^{-3} M_{\odot}/\text{pc}^3 = 0.038 \text{ GeV/cm}^3$

| Dark matter density function as Universal law for big galaxies – Unit DM Density GeV/cm ³ | | | | | | | |
|--|-----------|-----------------------|--------|--------|--------|--------|--|
| Density $_{DM} = 4,788 \cdot 10^{-8} \cdot E^{B}$ where $B = 1,74$ | | | | | | | |
| Where Unit for D _{D.M.} is GeV/cm ³ and Unit for E is Km ² /s ² /Kpc | | | | | | | |
| E gravt. | 100 | 500 | 1500 | 3000 | 4500 | 6500 | |
| Den. | 1,44.10-4 | 2,38·10 ⁻³ | 0,0161 | 0,0537 | 0,1088 | 0,2064 | |
| Dominion 100 < E < 6500 | | | | | | | |

12. CHECKING DM DENSITY FUNCTION WITH RECENTS MEASURES OF LOCAL DM DENSITY

Most common values of solar parameters are $R_{\Theta}=8\, Kpc$ and $V_{\Theta}=220\, Km/s$. As $E=V^2/R$ it is right to get $E_{\Theta}=6050\, Km^2/s^2/$ Kpc.

Using Density $_{\rm DM-SUN}=1,26\cdot 10^{-6}\cdot {\rm E}^{1,74}$ in paper units it is right to get Density $_{\rm DM-SUN}=4,8\cdot 10^{-3}\,M_{\odot}/{\rm pc}^3=4,8~{\rm m}\,M_{\odot}/{\rm pc}^3=0,182~{\rm GeV/cm}^3$

According this author [9] Pijushpani Bhattacharjee.2014 values for Sun are $R_{\Theta} = 8.3 \, \text{Kpc} \ V_{\Theta} = 244 \, \text{Km/s}$. These results show that even currently R_{Θ} and V_{Θ} are known with a wide experimental error.

Using such values, $E_{\Theta} = 7173 \text{ Km}^2/\text{s}^2/\text{ Kpc}$ so Density $_{\text{DM-SUN}} = 6,45 \cdot 10^{-3} M_{\Theta}/\text{pc}^3 = 6,45 \text{ m} M_{\Theta}/\text{pc}^3 = 0,245 \text{ GeV/cm}^3$

These values of DM density are very close to experimental measures.

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For example, bellow there are three recent paper about Local DM density measures.

Table 1. The best fit values of the parameters.

$$\begin{array}{|c|c|c|c|}\hline Parameter & Value \\ \hline \hline $\Sigma_{\star} \ (M_{\odot} \ pc^{-2})$ & $40.5^{+7.1}_{-6.6}$ \\ $\rho_{\rm DM} \ (M_{\odot} \ pc^{-3})$ & $0.0159^{+0.0047}_{-0.0057}$ \\ $z_{\rm h} \ (\rm pc)$ & 588^{+151}_{-192} \\ $h \ (\rm pc)$ & 293.8 ± 0.5 \\ \hline \end{array}$$

[6] Qiran Xia et al.2015
$$\varphi_{DM-LOCAL} = 0.0159 \pm 0.005 M_{\odot}/pc^3$$

[7] McKee, C.F.2015 $\varphi_{DM-LOCAL} = 0.013 \pm 0.003 M_{\odot}/\text{pc}^3$

In his remarkable paper the author - [8] J. I. Read.2014- presents a review of ancient and recent measures of local DM density which is summarized in table bellow.

| Label | Reference | Description | Sampling | $\rho_{\rm dm} \ [{\rm M}_{\odot} {\rm pc}^{-3}]$ | $\rho_{\rm dm}~[{\rm GeVcm^{-3}}$ | | |
|--|--------------------------|-----------------|-------------------|--|--|--|--|
| a) Local measures (ρ_{dm}) | | | | | | | |
| Kapteyn | Kapteyn (1922) | _ | _ | 0.0076 | 0.285 | | |
| Jeans | Jeans (1922) | _ | _ | 0.051 | 1.935 | | |
| Oort | Oort (1932) | _ | _ | 0.0006 ± 0.0184 | 0.0225 ± 0.69 | | |
| Hill | Hill (1960) | _ | _ | -0.0054 | -0.202 | | |
| Oort | Oort (1960) | _ | _ | 0.0586 ± 0.015 | 2.2 ± 0.56 | | |
| Bahcall | Bahcall (1984a) | _ | _ | 0.033 ± 0.025 | 1.24 ± 0.94 | | |
| Bienayme [†] | Bienayme et al. (1987) | _ | _ | 0.006 ± 0.005 | 0.22 ± 0.187 | | |
| KG [†] | Kuijken & Gilmore (1991) | _ | _ | 0.0072 ± 0.0027 | 0.27 ± 0.102 | | |
| Bahcall | Bahcall et al. (1992) | _ | _ | 0.033 ± 0.025 | 1.24 ± 0.94 | | |
| Creze | Creze et al. (1998) | _ | _ | -0.015 ± 0.015 | -0.58 ± 0.56 | | |
| HF^{\dagger} | Holmberg & Flynn (2000b) | _ | _ | 0.011 ± 0.01 | 0.4 ± 0.375 | | |
| HF^{\dagger} | Holmberg & Flynn (2004) | _ | _ | 0.0086 ± 0.0027 | 0.324 ± 0.1 | | |
| Bienayme | Bienaymé et al. (2006) | _ | _ | 0.0059 ± 0.005 | 0.51 ± 0.56 | | |
| Latest measurements | | | | | | | |
| MB12 | Moni Bidin et al. (2012) | CSF | 412 | 0.00062 ± 0.001 | 0.023 ± 0.042 | | |
| | | | | $[0 \pm 0.001]$ | $[0 \pm 0.042]$ | | |
| BT12 | Bovy & Tremaine (2012) | CSF | 412 | 0.008 ± 0.003 | 0.3 ± 0.11 | | |
| G12 | Garbari et al. (2012) | VC | 2×10^3 | $0.022^{+0.015}_{-0.013}$ | $0.85^{+0.57}_{-0.5}$ $0.33^{+0.26}_{-0.075}$ | | |
| G12* | Garbari et al. (2012) | $VC + \Sigma_b$ | 2×10^{3} | $0.0087^{+0.007}_{-0.002}$ | $0.33^{+0.26}_{-0.075}$ | | |
| S12 | Smith et al. (2012) | CSF | 10^{4} | 0.005 [no error] | 0.19 | | |
| | | | | [0.015] | [0.57] | | |
| Z13 | Zhang et al. (2013) | CSF | 10^{4} | 0.0065 ± 0.0023 | 0.25 ± 0.09 | | |
| BR13 | Bovy & Rix (2013) | CSF + MAP | 10^{4} | 0.006 ± 0.0018 | 0.22 ± 0.07 | | |
| | | | | $[0.008 \pm 0.0025]$ | $[0.3 \pm 0.094]$ | | |
| b) Global measures assuming spherical symmetry ($\rho_{dm,ext}$) | | | | | | | |
| S10 | Salucci et al. (2010) | NP | _ | 0.011 ± 0.004 | 0.43 ± 0.15 | | |
| CU10 | Catena & Ullio (2010) | NFW; SP | _ | 0.0103 ± 0.00072 | 0.385 ± 0.027 | | |
| WB10 | Weber & de Boer (2010) | NFW/ISO; WP | _ | 0.005 - 0.01 | 0.2 - 0.4 | | |
| I11 | Iocco et al. (2011) | gNFW; WP; ML | _ | 0.005 - 0.015 | 0.2 - 0.56 | | |
| M11 | McMillan (2011) | NFW; SP | _ | 0.011 ± 0.0011 | 0.4 ± 0.04 | | |

Although I recommend to read the paper, briefly I comment that measures of group a) are made through local methods which involve measures of stars few hundred parsecs away. Local measures use the vertical kinematics of stars near the Sun called `tracers'. Group of measures b) involve global measures which extrapolate $\varphi_{DM-EXTR}$ from the rotation curve of Galaxy

DISCUSSION

Despite the fact that currently there are lot of groups of astrophysicist working in Local DM with sophisticated instruments, it is clear that differences between different measures are important.

By other side result of DM density got through the DM density function is similar to low values of DM density measures.

Therefore it can be concluded that DM density function is compatible with current measures of Local DM density.

13.DARK MATTER DENSITY FUNCTION INSIDE INTERMEDIATE GALAXIES

The same method used by big spiral galaxies was used for intermediate galaxies which spin speed in flat region is under 150 km/s

Particularly were studied rotation curves of NGC 3198 (Ursa Major), NGC 2403 (M81 group of galaxies) published by [2] Randriamampandry, T. and [3] Bottema, R.B. In addition was studied M33-Triangle galaxy- using rotation curve published by [5] Corbelli, E which is a satellite galaxy of M31 and its distance to the Earth is under 2 Mpc.

In these galaxies density of dark matter is more than four times bigger that density of dark matter inside big galaxies although these results has not been published in this paper.

The reason to not publish these results is that the author is looking for a Universal law for Density of dark matter therefore apparently these results destroy the paper objective.

However, I think that these results can not discard this model because the gravitational field inside an intermediate galaxy may be influenced by its nearby giant galaxy. In addition, a nearby giant galaxy may break significantly the spherical symmetry of gravitational field inside an intermediate galaxy and therefore calculus made through Virial Theorem become wrong.

I encourage the reader to consult the previous paper the author [4] Abarca, M.2014 to understand better the reasons why behaviour of density of D.M. inside intermediate galaxies cannot reject density of D.M. depending on E as a Universal law.

According the theory exposed in that paper could be better understood the following example of DM inside intermediate galaxies.

M33 -The triangle galaxy- is 220 Kpc away to M31. As I defend in [4] Abarca,M.2014 that M31 halo radius is 375 Kpc, it is obvious that M33 is placed inside M31 halo. As it is known M31 is a big galaxy twin of Milky Way. Therefore, total amount of DM in M33 come not only from DM generated by gravitational field of M33 but also from Density of DM generates by M31. Therefore is logical that Density of DM measured in M33 is bigger that density DM calculated by the law got from eight data set because this law is for big galaxies and inside this galaxies, dark matter only come from mechanism of generation of its own gravitational field.

For this reason it should be a Universal law the following statement:

Density of DM on intermediate galaxies, which are inside halo a big galaxy should be bigger that density of DM on big galaxies.

14. CONCLUSION

The objective the paper has been try to justify that DM density function depend on E according a Universal law.

According my previous paper [4] Abarca,M.2014, DM is generated by gravitational field, so I thought that it was needed to study big galaxies because inside their disks, gravitational field may be measured through rotational curve of a galaxy. In other words, inside a big galaxy it is known that gravitational field is generated by its own mass and gravitational influence of galactic neighbours is negligible.

After a long search in arXiv looking for papers published during last five years only I found four papers with rotational curves published.

Apart from dwarf an intermediate galaxies only there is eight big galaxies, whose spin speed in flat region is higher than 200 Km/s. By reasons explained in epigraph 3, I had to reject two galaxies although I found two galaxies studied by two authors simultaneously, so finally I have worked with eight galactic data set -8GDS-. Specifically the galaxies are Milky Way, M31, NGC 3031, NGC 3992, NGC 7331, NGC 2841.

After a long statistical study carefully detailed through the paper I have calculated an "average" curve which fit all data set studied whose correlation coefficient is bigger than 0,997, so I have found plausible to postulate this function as Universal:

| Dark matter density function as Universal law for big galaxies – flat speed region in rot. curve above 200 Km/s - | | | | | | | |
|---|----------|---------|---------|--------|--------|--------|--|
| Density $_{DM} = 1,26 \cdot 10^{-6} \cdot E^{B}$ where $B = 1,74$ | | | | | | | |
| Where Unit for $D_{D.M.}$ is $10^{-3} M_{\Theta}/pc^3$ and Unit for E is $Km^2/s^2/Kpc$ | | | | | | | |
| E gravt. | 100 | 500 | 1500 | 3000 | 4500 | 6500 | |
| Den. | 0,003787 | 0,06226 | 0,42093 | 1,4056 | 2,8456 | 5,3948 | |
| Dominion 100 < E < 6500 | | | | | | | |

As a previous result was fit a linear function to DM density of 8GDS inside dominion 1500 < E < 6000.

| Linear function fitted to DM Density DM Density = 0,0008 · E -0,8367 and correlation R = 0,916 | | | | | | |
|--|--------|------|------|------|------|------|
| E gravt. | 1500 | 2000 | 3000 | 4000 | 5000 | 6000 |
| Den. | 0,3633 | 0,76 | 1,56 | 2,36 | 3,16 | 3,96 |

It is right to get that a linear function is an approximation of a potential function for a restricted area of dominion.

According theory defended by the author in [4] Abarca, M.2014, DM is generated by the gravitational field so when E goes to cero DM density should go to cero as well. Therefore potential fit is the most suitable function.

In chapter 12 it is calculated Local DM density in vicinity of Sun. From Solar parameters $R_{\Theta} = 8 \, Kpc$ and $V_{\Theta} = 220 \, Km/s$ it is got $E_{\Theta} = 6050 \, \mathrm{Km^2/s^2/Kpc}$ and $\varphi_{LOCAL-DM} = 4.8 \, \mathrm{m} \, M_{\Theta}/\mathrm{pc^3} = 0.182 \, \mathrm{GeV/cm^3}$. This value has been compared with data which come from recent Local DM density measures and it has been concluded that $\varphi_{LOCAL-DM}$ calculated in this paper is included inside interval defined by $\varphi_{LOCAL-DM}$ measures published in recent papers. Therefore it can be concluded that DM density function is compatible with current measures of Local DM density.

Apart from the six big galaxies I have studied rotation curves of NGC 3198 (Ursa Major), NGC 2403 (M81 group of galaxies) and M33 -Triangle galaxy which are galaxies whose spin speed in flat region are under 150 Km/s and they belong to cluster of galaxies dominates by gravitational field of big galaxies.

In chapter 13 were pointed out that experimental results for intermediate an dwarf galaxies show that DM density is a great deal bigger than inside big galaxies. However also were pointed out reasons that in opinion the author disallow

this results as an ultimate counter-example to reject that Density of DM depend on gravitational field E according a Universal law. In addition, reasons developed in this epigraph lead to postulate a new Universal law:

DM density on intermediate galaxies near a big galaxy should have a bigger DM density than DM density in big galaxies because DM density inside intermediate galaxies it is not only generated by its own gravitational field but also by its nearby giant galaxy.

The objective this paper has been to check the model of dark matter proposed by the author in [4] Abarca,M.2014. I think that it has been justified properly that results could confirm DM density as universal law. In my opinion statistical differences in different galaxies might be explained by experimental error in rotation curves and error in calculus of DM density functions.

The author think that it may be worth to study a bigger number of galaxies in order to check if new results could confirm or reject theory of Dark matter proposed by the author because Dark matter nature is one of the most important challenge in astrophysics science nowadays.

15. BIBLIOGRAPHIC REFERENCES

- [1] Sofue,Y.2015,arXiv:1504.05368v1 Dark halos of M31 and the Milky Way.
- [2] Randriamampandry, T. Carignan, C.2014.arXiv:1401.5619v1 Galaxy Mass Models: MOND versus Dark Matter Halos
- [3] Bottema,R.B. Pestaña,J.LG.2015,arXiv:1501.06424v1

 The distribution of dark and luminous matter inferred from extended rotation curves
- [4] Abarca,M.2014,viXra:1410.0200

 Dark matter model by quantum vacuum
- [5] Corbelli,E. et al. 2014,arXiv:1409.2665v2

 Dynamical signatures of a Λ CDM-halo and the distribution of the baryons in M33
- [6] Qiran Xia et al.2015.arXiv:1510.06810v1

Determining the local dark matter density with LAMOST data

- [7] Christopher F. McKee.2015.arXiv: 1509.05334v1 Stars, Gas, and Dark Matter in the Solar Neighborhood
- [8] J. I. Read.2014. arXiv:1404.1938v2 *The Local Dark Matter Density*
- [9] Pijushpani Bhattacharjee.2014.arXiv: 1310.2659v3 Rotation curve of the milky way out to 200 Kpc