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Space-time Fourier transform, convolution and Mustard convolution

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Abstract. In this paper we use the steerable space-time Fourier transform (SFT), and relate the classical convolution of the algebra for space-time $Cl(3, 1)$ -valued signals over the space-time vector space $\mathbb{R}^{3,1}$, with the (equally steerable) Mustard convolution. A Mustard convolution can be expressed in the spectral domain as the point wise product of the SFTs of the factor functions. In full generality do we express the classical convolution of space-time signals in terms of finite linear combinations of Mustard convolutions, and vice versa the Mustard convolution of space-time signals in terms of finite linear combinations of classical convolutions.

Keywords. Convolution, Mustard convolution, space-time Fourier transform, space-time signals, space-time domain, frequency domain.

1. Introduction

The quaternions frequently appear as subalgebras of higher order Clifford geometric algebras [2, 21]. This is for example the case for the Clifford algebra over the space-time vector space [8, 9, 7], which is of prime importance in physics, and in applications where time matters as well (motion in time, video sequences, flow fields, ...). The quaternion subalgebra structure allows to introduce generalizations of the quaternion Fourier transform (QFT) to functions in these higher order Clifford geometric algebras. For example it allows to generalize the QFT to a space-time Fourier transform [10, 13].

Recently it has been shown how the left-sided QFT [4], and the two-sided QFT [20] allow to define Mustard convolutions for which in the spectral domain the QFT of the convolution becomes a simple point wise product of

This paper is dedicated to Rev. Dr. Gideon Jacob, of Trichy, Tamil Nadu, India. Together with his church he founded *MoseMinistries* and saved since 1994 the lives of 89 Indian baby girls [1]. The use of this paper is subject to the *Creative Peace License* [15].

the QFTs of the quaternion signal functions. This paper generalizes the approach of [20] from four-dimensional quaternions to the 16-dimensional Clifford algebra $Cl(3, 1)$, which contains a subalgebra isomorphic to quaternions. Because transferring results from a lower dimensional non-commutative algebra to a higher-dimensional non-commutative algebra is non-trivial, we try to work with sufficient algebraic detail to allow all results to be verified directly.

This paper is organized as follows. Section 2 reviews the Clifford geometric algebra $Cl(3, 1)$ of the space-time vector space $\mathbb{R}^{3,1}$. In particular a subalgebra isomorphic to quaternions is studied, which is generated by the time-vector and the three-dimensional space volume pseudoscalar. Section 3 reviews the (steerable) space-time Fourier transform (SFT) of [10, 13], and newly introduces related exponential-sine Fourier transforms. Section 4 first defines the convolution, and two types of (equally steerable) Mustard convolution for space-time signals in $Cl(3, 1)$ over $\mathbb{R}^{3,1}$. The main results are Theorem 4.3 describing the convolution of space-time signals in terms of the two types of Mustard convolutions, Theorem 4.5 expressing the convolution in terms of only four standard Mustard convolutions, and finally vice versa Theorem 4.6 describing the standard Mustard convolution of space-time signals in terms of eight classical convolutions.

2. Algebra for space-time

The algebra for space-time $Cl(3, 1) = Cl_{3,1} = \mathcal{G}_{3,1} = \mathbb{R}_{3,1}$ is Clifford's geometric algebra of $\mathbb{R}^{3,1}$. In $\mathbb{R}^{3,1}$ we can introduce the following orthonormal vector basis,

$$\{\mathbf{e}_t, \mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3\}, \quad -\mathbf{e}_t^2 = \mathbf{e}_1^2 = \mathbf{e}_2^2 = \mathbf{e}_3^2 = 1. \quad (2.1)$$

In the full blade basis of $Cl(3, 1)$ we thus get three anti-commuting blades that all square to minus one, they are some of the roots of -1 (compare [17]),

$$\mathbf{e}_t^2 = -1, \quad i_3 = \mathbf{e}_1\mathbf{e}_2\mathbf{e}_3, \quad i_3^2 = -1, \quad i_{st} = \mathbf{e}_t\mathbf{e}_1\mathbf{e}_2\mathbf{e}_3, \quad i_{st}^2 = -1, \quad (2.2)$$

and the commutator

$$[\mathbf{e}_t, i_3] = 2\mathbf{e}_t i_3 = 2i_{st}. \quad (2.3)$$

The volume-time subalgebra of $Cl(3, 1)$ generated by these blades is indeed isomorphic to the quaternion algebra [7].

$$\{1, \mathbf{e}_t, i_3, i_{st}\} \longleftrightarrow \{1, \mathbf{i}, \mathbf{j}, \mathbf{k}\} \quad (2.4)$$

This isomorphism allows us now to transfer the quaternionic \pm split (or orthogonal two-dimensional planes split) of [10, 13, 19, 14, 16] to space-time algebra, which turns out to be a very real (physical) space-time split

$$h_{\pm} = \frac{1}{2}(h \pm \mathbf{e}_t h \mathbf{e}_t^*), \quad h = h_+ + h_-, \quad (2.5)$$

where $\mathbf{e}_t^* = i_3$, is the space-time dual of the unit time direction \mathbf{e}_t , i.e.,

$$\mathbf{e}_t^* = \mathbf{e}_t i_{st}^{-1} = -\mathbf{e}_t i_{st} = -\mathbf{e}_t \mathbf{e}_1 \mathbf{e}_2 \mathbf{e}_3 = i_3. \quad (2.6)$$

	space-time algebra basis elements															
inv.	1	e_{23}	e_{31}	e_{12}	e_1	e_2	e_3	i_3	e_t	e_{t23}	e_{t31}	e_{t12}	e_{t1}	e_{t2}	e_{t3}	i_{st}
$e_t()e_t$	-	-	-	-	+	+	+	+	-	-	-	-	+	+	+	+
$i_3()i_3$	-	-	-	-	-	-	-	-	+	+	+	+	+	+	+	+
$i_{st}()i_{st}$	-	-	-	-	+	+	+	+	+	+	+	+	-	-	-	-
$e_t()i_3$	i_{st}	$-e_{t1}$	$-e_{t2}$	$-e_{t3}$	e_{t23}	e_{t31}	e_{t12}	$-e_t$	$-i_3$	e_1	e_2	e_3	$-e_{23}$	$-e_{31}$	$-e_{12}$	-1
$e_t()i_{st}$	$-i_3$	e_1	e_2	e_3	e_{23}	e_{31}	e_{12}	-1	$-i_{st}$	e_{t1}	e_{t2}	e_{t3}	e_{t23}	e_{t31}	e_{t12}	$-e_t$
$i_3()i_{st}$	e_t	e_{t23}	e_{t31}	e_{t12}	$-e_{t1}$	$-e_{t2}$	$-e_{t3}$	$-i_{st}$	1	e_{23}	e_{31}	e_{12}	$-e_1$	$-e_2$	$-e_3$	$-i_3$

TABLE 1. Involutions of space-time algebra $Cl(3, 1)$. If the involution only changes the sign of the element, only the sign is given. Abbreviations: $i_3 = e_{123}$, $i_{st} = e_{t123}$, $e_{t1} = e_t e_1$, $e_{12} = e_1 e_2$, etc.

The time direction e_t determines therefore the complementary three-dimensional physical Euclidean space with pseudoscalar i_3 as well! Their product $i_{st} = e_t i_3$ is the four-dimensional space-time hypervolume pseudoscalar. Note that

$$e_t h i_3 = h_+ - h_-, \quad (2.7)$$

i.e. under the involution map $e_t()i_3$ the h_+ part is invariant, but the h_- part changes sign, which is related to the Coxeter half-turn [3]. See also Table 1.

We further note, that with respect to $f \in \{e_t, i_3, i_{st}\} \subset \mathbb{R}^{3,1}$, every multivector $A \in Cl(3, 1)$ can be split into *commuting* and *anticommuting* parts [17].

Lemma 2.1 (Commuting and anticommuting with $f \in \{e_t, i_3, i_{st}\} \subset \mathbb{R}^{3,1}$ [17]). *Every multivector $A \in Cl(3, 1)$ has, with respect to every $f \in \{e_t, i_3, i_{st}\} \subset \mathbb{R}^{3,1}$, where we note that $f^{-1} = -f$, the unique decomposition denoted by*

$$A_{+f} = \frac{1}{2}(A + f^{-1}Af), \quad A_{-f} = \frac{1}{2}(A - f^{-1}Af)$$

$$A = A_{+f} + A_{-f}, \quad A_{+f}f = fA_{+f}, \quad A_{-f}f = -fA_{-f}. \quad (2.8)$$

Notation 2.2 (Argument reflection). *For a function $h : \mathbb{R}^{3,1} \rightarrow Cl(3, 1)$ and a multi-index $\phi = (\phi_1, \phi_2)$ with $\phi_1, \phi_2 \in \{0, 1\}$ we set*

$$h^\phi = h^{(\phi_1, \phi_2)}(\mathbf{x}) := h((-1)^{\phi_1}t, (-1)^{\phi_2}\vec{x}). \quad (2.9)$$

In (2.5) the involution $e_t()e_t^* = e_t()i_3$, plays an important role. For the quaternion algebra \mathbb{H} this has been studied in [10, 13, 19, 14, 16]. The isomorphism (2.4) immediately allows to transfer these results to the space-time subalgebra $\{1, e_t, i_3, i_{st}\}$. Involutions like $e_t()i_3$, and decompositions like (2.5) and Lemma 2.1, provided the key to the geometric interpretation of the two-sided QFT [19, 14, 16], and are significant and efficient in establishing several types of quaternion signal convolutions both in the spatial, as well as in the spectral domain [20].

We therefore begin our investigation of the convolutions of space-time signals, $h : \mathbb{R}^{3,1} \rightarrow Cl(3, 1)$, also with the study of involutions of $Cl(3, 1)$, using the three square roots of -1 : $\{e_t, i_3, i_{st}\}$. Following Table 1, and giving the 16 element basis set of the algebra for space-time $Cl(3, 1)$ in the first line

the name B , we find the following important basis subsets, spanning eight-dimensional subspaces

$$\begin{aligned}
B_+ &= \{e_t - i_3, (e_t - i_3)e_1, (e_t - i_3)e_2, (e_t - i_3)e_3, \\
&\quad 1 + i_{st}, (1 + i_{st})e_1, (1 + i_{st})e_2, (1 + i_{st})e_3\}, \\
B_- &= \{e_t + i_3, (e_t + i_3)e_1, (e_t + i_3)e_2, (e_t + i_3)e_3, \\
&\quad 1 - i_{st}, (1 - i_{st})e_1, (1 - i_{st})e_2, (1 - i_{st})e_3\}, \\
B_{+e_t} &= \{1, e_{23}, e_{31}, e_{12}, e_t, e_{t23}, e_{t31}, e_{t12}\}, \\
B_{-e_t} &= \{e_1, e_2, e_3, i_3, e_{t1}, e_{t2}, e_{t3}, i_{st}\}, \\
B_{+i_3} &= \{1, e_{23}, e_{31}, e_{12}, e_1, e_2, e_3, i_3\}, \\
B_{-i_3} &= \{e_t, e_{t23}, e_{t31}, e_{t12}, e_{t1}, e_{t2}, e_{t3}, i_{st}\},
\end{aligned} \tag{2.10}$$

where B_{\pm} is defined by (2.5), and $B_{\pm e_t}$, $B_{\pm i_3}$ according to Lemma 2.1. The eight-dimensional plus and minus parts of the algebra $Cl(3, 1)$ arising from the split with (2.5) can also be specified as

$$\begin{aligned}
Cl(3, 1)_+ &= \text{span}[e_t - i_3, (e_t - i_3)\vec{x}, 1 + i_{st}, (1 + i_{st})\vec{y}; \forall \vec{x}, \vec{y} \in \mathbb{R}^3], \\
Cl(3, 1)_- &= \text{span}[e_t + i_3, (e_t + i_3)\vec{x}, 1 - i_{st}, (1 - i_{st})\vec{y}; \forall \vec{x}, \vec{y} \in \mathbb{R}^3].
\end{aligned} \tag{2.11}$$

The following *identities* hold for $m_{\pm} \in Cl(3, 1)_{\pm}$,

$$e^{\alpha e_t} m_{\pm} e^{\beta i_3} = m_{\pm} e^{(\beta \mp \alpha) i_3} = e^{(\alpha \mp \beta) e_t} m_{\pm}. \tag{2.12}$$

Particularly useful cases of (2.12) are $(\alpha, \beta) = (\pi/2, 0)$ and $(0, \pi/2)$:

$$e_t m_{\pm} = \mp m_{\pm} i_3, \quad m_{\pm} i_3 = \mp e_t m_{\pm}. \tag{2.13}$$

Because of (2.3), we have for the product of exponentials

$$\begin{aligned}
e^{\alpha e_t} e^{\beta i_3} &= e^{\beta i_3} e^{\alpha e_t} + [e_t, i_3] \sin(\alpha) \sin(\beta) \\
&= e^{\beta i_3} e^{\alpha e_t} + 2i_{st} \sin(\alpha) \sin(\beta),
\end{aligned} \tag{2.14}$$

which (in the same form for general multivector square roots of -1) has been used in [18] in order to derive a general convolution theorem for Clifford Fourier transformations. Moreover, note that

$$e_t [e_t, i_3] = 2(e_t)^2 i_3 = -2i_3. \tag{2.15}$$

We furthermore note the useful anticommutation relationships

$$e_t [e_t, i_3] = -[e_t, i_3] e_t, \quad i_3 [e_t, i_3] = -[e_t, i_3] i_3, \tag{2.16}$$

and therefore

$$e^{\alpha e_t} [e_t, i_3] = [e_t, i_3] e^{-\alpha e_t}, \quad e^{\beta i_3} [e_t, i_3] = [e_t, i_3] e^{-\beta i_3}. \tag{2.17}$$

And because of the fundamental anticommutation

$$e_t i_3 = -i_3 e_t \Rightarrow e^{\alpha e_t} i_3 = i_3 e^{-\alpha e_t}. \tag{2.18}$$

3. The steerable space-time Fourier transform (SFT)

The steerable space-time Fourier transform maps 16-dimensional space-time algebra functions $h : \mathbb{R}^{3,1} \rightarrow Cl_{3,1}$ to 16-dimensional space-time spectrum functions $\mathcal{F}\{h\} : \mathbb{R}^{3,1} \rightarrow Cl_{3,1}$. It is defined in the following way

$$h \rightarrow \mathcal{F}^{e_t, i_3}\{h\}(\boldsymbol{\omega}) = \mathcal{F}\{h\}(\boldsymbol{\omega}) = \int_{\mathbb{R}^{3,1}} e^{-\mathbf{e}_t t \omega_t} h(\mathbf{x}) e^{-i_3 \vec{x} \cdot \vec{\omega}} d^4 \mathbf{x}, \quad (3.1)$$

with

- space-time vectors $\mathbf{x} = t\mathbf{e}_t + \vec{x} \in \mathbb{R}^{3,1}$, $\vec{x} = x\mathbf{e}_1 + y\mathbf{e}_2 + z\mathbf{e}_3 \in \mathbb{R}^3$
- space-time volume $d^4 \mathbf{x} = dt dx dy dz$
- space-time frequency vectors $\boldsymbol{\omega} = \omega_t \mathbf{e}_t + \vec{\omega} \in \mathbb{R}^{3,1}$, $\vec{\omega} = \omega_1 \mathbf{e}_1 + \omega_2 \mathbf{e}_2 + \omega_3 \mathbf{e}_3 \in \mathbb{R}^3$

Note, that we usually omit the upper indexes showing the special square roots of -1 selected for the transform, as in $\mathcal{F}^{e_t, i_3}\{h\} = \mathcal{F}\{h\}$.

Remark 3.1. *The above SFT is a steerable operator¹ depending on the choice of unit time direction \mathbf{e}_t in the forward light cone of $\mathbb{R}^{3,1}$. In the case of a local inertial frame of reference, the vector \mathbf{e}_t in $\mathbb{R}^{3,1}$ specifies the velocity of the observer.*

Remark 3.2. *The three-dimensional integration part*

$$\int h(\mathbf{x}) e^{-i_3 \vec{x} \cdot \vec{\omega}} d^3 \vec{x}$$

in (3.1) fully corresponds to the Clifford algebra Fourier transform (CFT) in $Cl(3,0)$, compare [11, 12].

The inverse \mathcal{F}^{-1} of the SFT (3.1) is given by

$$h(\mathbf{x}) = \frac{1}{(2\pi)^4} \int_{\mathbb{R}^{3,1}} e^{\mathbf{e}_0 t \omega_t} \mathcal{F}\{h\}(\boldsymbol{\omega}) e^{i_3 \vec{x} \cdot \vec{\omega}} d^4 \boldsymbol{\omega}. \quad (3.2)$$

The \pm split of the QFT can now, via the isomorphism (2.4) of quaternions to the volume-time subalgebra of the space-time algebra, be extended to splitting general space-time algebra multivector functions over $\mathbb{R}^{3,1}$. This leads to the following interesting result [10],

$$\begin{aligned} \mathcal{F}\{h\} &= \mathcal{F}\{h\}_+ + \mathcal{F}\{h\}_- \\ &= \int_{\mathbb{R}^{3,1}} h_+ e^{-i_3(\vec{x} \cdot \vec{\omega} - t \omega_t)} d^4 \mathbf{x} + \int_{\mathbb{R}^{3,1}} h_- e^{-i_3(\vec{x} \cdot \vec{\omega} + t \omega_t)} d^4 \mathbf{x} \\ &= \int_{\mathbb{R}^{3,1}} e^{-\mathbf{e}_0(t \omega_t - \vec{x} \cdot \vec{\omega})} h_+ d^4 \mathbf{x} + \int_{\mathbb{R}^{3,1}} e^{-\mathbf{e}_0(t \omega_t + \vec{x} \cdot \vec{\omega})} h_- d^4 \mathbf{x}. \end{aligned} \quad (3.3)$$

This result shows us that the SFT is identical to a sum of right and left propagating multivector wave packets. We therefore see that these physically important wave packets arise absolutely naturally from elementary purely algebraic considerations.

¹Note that the *steerability* of the closely related general two-sided QFT has been discussed at length in [14, 16].

We further define for later use the following two mixed *exponential-sine* Fourier transforms

$$\mathcal{F}^{e_t, \pm s}\{h\}(\omega) = \int_{\mathbb{R}^{3,1}} e^{-e_t t \omega_t} h(\mathbf{x}) (\pm 1) \sin(-\vec{x} \cdot \vec{\omega}) d^4 \mathbf{x}, \quad (3.4)$$

$$\mathcal{F}^{\pm s, i_3}\{h\}(\omega) = \int_{\mathbb{R}^{3,1}} (\pm 1) \sin(-t \omega_t) h(\mathbf{x}) e^{-i_3 \vec{x} \cdot \vec{\omega}} d^4 \mathbf{x}. \quad (3.5)$$

With the help of

$$\begin{aligned} \sin(-t \omega_t) &= \frac{e_t}{2} (e^{-e_t t \omega_t} - e^{e_t t \omega_t}), \\ \sin(-\vec{x} \cdot \vec{\omega}) &= \frac{i_3}{2} (e^{-i_3 \vec{x} \cdot \vec{\omega}} - e^{i_3 \vec{x} \cdot \vec{\omega}}), \end{aligned} \quad (3.6)$$

we can rewrite the above mixed exponential-sine Fourier transforms in terms of the SFT of (3.1) as

$$\mathcal{F}^{e_t, \pm s}\{h\} = \pm \frac{1}{2} (\mathcal{F}^{e_t, i_3}\{h i_3\} - \mathcal{F}^{e_t, -i_3}\{h i_3\}), \quad (3.7)$$

$$\mathcal{F}^{\pm s, i_3}\{h\} = \pm \frac{1}{2} (\mathcal{F}^{e_t, i_3}\{e_t h\} - \mathcal{F}^{-e_t, i_3}\{e_t h\}). \quad (3.8)$$

We further note the following useful relationships using the argument reflection of Notation 2.2

$$\mathcal{F}^{-e_t, i_3}\{h\} = \mathcal{F}^{e_t, g}\{h^{(1,0)}\} = \mathcal{F}\{h^{(1,0)}\}, \quad \mathcal{F}^{e_t, -g}\{h\} = \mathcal{F}\{h^{(0,1)}\}, \quad (3.9)$$

and similarly

$$\mathcal{F}^{e_t, -s}\{h\} = \mathcal{F}^{e_t, s}\{h^{(0,1)}\}, \quad \mathcal{F}^{-s, i_3}\{h\} = \mathcal{F}^{s, i_3}\{h^{(1,0)}\}. \quad (3.10)$$

4. Convolution and Mustard convolution

We define the *convolution* of two quaternion signals $a, b \in L^1(\mathbb{R}^{3,1}; Cl_{3,1})$ as

$$(a \star b)(\mathbf{x}) = \int_{\mathbb{R}^{3,1}} a(\mathbf{y}) b(\mathbf{x} - \mathbf{y}) d^2 \mathbf{y}, \quad (4.1)$$

provided that the integral exists.

The *Mustard convolution* [22] of two quaternion signals $a, b \in L^1(\mathbb{R}^{3,1}; Cl_{3,1})$ is defined as

$$(a \star_M b)(\mathbf{x}) = \mathcal{F}^{-1}(\mathcal{F}\{a\} \mathcal{F}\{b\}). \quad (4.2)$$

provided that the integral exists.

Remark 4.1. *The Mustard convolution has the conceptual and computational advantage to simply yield as spectrum in the SFT Fourier domain the point wise product of the SFTs of the two signals, just as for the classical complex Fourier transform. On the other hand, by its very definition, the Mustard convolution depends on the choice of the pair e_t, i_3 , of square roots of -1 used in the definition (3.1) of the SFT. The Mustard convolution (4.2) is therefore a steerable operator, depending on the choice of unit time direction e_t in the space-time forward light cone of $\mathbb{R}^{3,1}$. This may be of advantage in applications to special relativistic physics, electromagnetic signal processing, optics, and aero-space navigation.*

We additionally define a further type of (steerable) *exponential-sine* Mustard convolution as

$$(a \star_{Ms} b)(\mathbf{x}) = \mathcal{F}^{-1}(\mathcal{F}^{e_t, s}\{a\}\mathcal{F}^{s, i_3}\{b\}). \quad (4.3)$$

In the following two Subsections we will first express the convolution (4.1) in terms of the Mustard convolution (4.2) and vice versa.

4.1. Expressing the convolution in terms of the Mustard convolution

In [4] Theorem 4.1 on page 584 expresses the classical convolution of two quaternion functions with the help of the general left-sided QFT as a sum of 40 Mustard convolutions. Similar results have been established for the general two-sided QFT in [20]. Based on the isomorphism (2.4) and on the splits of (2.5) and of Lemma 2.1, we generalize these results now to the SFT. Moreover, we use Theorem 5.12 on page 327 of [18], which expresses the convolution of two Clifford signal functions (higher dimensional generalizations of quaternion or space-time functions) in the Clifford Fourier domain with the help of the general two-sided Clifford Fourier transform (CFT), the latter is in turn a generalization of the QFT and SFT to general Clifford algebras with non-degenerate quadratic forms. We restate this theorem here again, specialized for space-time functions and the SFT of (3.1).

Theorem 4.2 (SFT of convolution). *The SFT of the convolution (4.1) of two functions $a, b \in L^1(\mathbb{R}^{3,1}; Cl_{3,1})$ can be expressed as*

$$\begin{aligned} \mathcal{F}\{a \star b\} = & \\ & \mathcal{F}\{a_{+e_t}\}\mathcal{F}\{b_{+i_3}\} + \mathcal{F}^{e_t, -i_3}\{a_{+e_t}\}\mathcal{F}^{e_t, i_3}\{b_{-i_3}\} \\ & + \mathcal{F}^{e_t, i_3}\{a_{-e_t}\}\mathcal{F}^{-e_t, i_3}\{b_{+i_3}\} + \mathcal{F}^{e_t, -i_3}\{a_{-e_t}\}\mathcal{F}^{-e_t, -i_3}\{b_{-i_3}\} \\ & + 2\mathcal{F}^{e_t, s}\{a_{+e_t}\}i_{st}\mathcal{F}^{s, i_3}\{b_{+i_3}\} + 2\mathcal{F}^{e_t, -s}\{a_{+e_t}\}i_{st}\mathcal{F}^{s, i_3}\{b_{-i_3}\} \\ & + 2\mathcal{F}^{e_t, s}\{a_{-e_t}\}i_{st}\mathcal{F}^{-s, i_3}\{b_{+i_3}\} + 2\mathcal{F}^{e_t, -s}\{a_{-e_t}\}i_{st}\mathcal{F}^{-s, i_3}\{b_{-i_3}\}. \end{aligned} \quad (4.4)$$

Note that due to the commutation properties of (3.4) and (3.5) we can place the pseudosalar i_{st} also inside the exponential-sine transform terms as e.g. in

$$\begin{aligned} \mathcal{F}^{e_t, s}\{a_{+e_t}\}i_{st}\mathcal{F}^{s, i_3}\{b_{+i_3}\} &= \mathcal{F}^{e_t, s}\{a_{+e_t}i_{st}\}\mathcal{F}^{s, i_3}\{b_{+i_3}\} \\ &= \mathcal{F}^{e_t, s}\{a_{+e_t}\}\mathcal{F}^{s, i_3}\{i_{st}b_{+i_3}\}. \end{aligned} \quad (4.5)$$

By applying the inverse SFT to (4.4), we can now easily express the convolution of two space-time signals $a \star b$ in terms of only eight Mustard convolutions (4.2) and (4.3).

Theorem 4.3 (Convolution in terms of two types of Mustard convolution). *The convolution (4.1) of two space-time functions $a, b \in L^1(\mathbb{R}^{3,1}; Cl_{3,1})$ can be*

expressed in terms of four Mustard convolutions (4.2) and four exponential-sine Mustard convolutions (4.3) as

$$\begin{aligned}
a \star b &= a_{+e_t} \star_M b_{+i_3} + a_{+e_t}^{(0,1)} \star_M b_{-i_3} + a_{-e_t} \star_M b_{+i_3}^{(1,0)} + a_{-e_t}^{(0,1)} \star_M b_{-i_3}^{(1,0)} \\
&\quad + 2a_{+e_t} \star_{Ms} i_{st} b_{+i_3} + 2a_{+e_t}^{(0,1)} \star_{Ms} i_{st} b_{-i_3} \\
&\quad + 2a_{-e_t} \star_{Ms} i_{st} b_{+i_3}^{(1,0)} + 2a_{-e_t}^{(0,1)} \star_{Ms} i_{st} b_{-i_3}^{(1,0)}. \tag{4.6}
\end{aligned}$$

Remark 4.4. We use the convention, that terms such as $a_{+e_t} \star_{Ms} i_{st} b_{+i_3}$, should be understood with brackets $a_{+e_t} \star_{Ms} (i_{st} b_{+i_3})$, which are omitted to avoid clutter.

Furthermore, applying (3.7) and (3.8), we can expand the terms in (4.4) with exponential-sine transforms into sums of products of SFTs. For example, the first term gives, using $e_t i_{st} = -i_3$,

$$\begin{aligned}
&\mathcal{F}^{e_t, s} \{a_{+e_t}\} i_{st} \mathcal{F}^{s, i_3} \{b_{+i_3}\} \\
&= \frac{1}{4} (\mathcal{F}^{e_t, i_3} \{a_{+e_t} i_3\} - \mathcal{F}^{e_t, -i_3} \{a_{+e_t} i_3\}) (\mathcal{F}^{e_t, i_3} \{e_t i_{st} b_{+i_3}\} - \mathcal{F}^{-e_t, i_3} \{e_t i_{st} b_{+i_3}\}) \\
&= \frac{1}{4} (\mathcal{F}^{e_t, i_3} \{a_{+e_t} i_3\} - \mathcal{F}^{e_t, -i_3} \{a_{+e_t} i_3\}) (\mathcal{F}^{e_t, i_3} \{-i_3 b_{+i_3}\} - \mathcal{F}^{-e_t, i_3} \{-i_3 b_{+i_3}\}) \\
&= \frac{1}{4} (\mathcal{F} \{a_{+e_t} i_3\} \mathcal{F} \{-i_3 b_{+i_3}\} - \mathcal{F} \{a_{+e_t} i_3\} \mathcal{F} \{-i_3 b_{+i_3}^{(1,0)}\} \\
&\quad - \mathcal{F} \{a_{+e_t}^{(0,1)} i_3\} \mathcal{F} \{-i_3 b_{+i_3}\} + \mathcal{F} \{a_{+e_t}^{(0,1)} i_3\} \mathcal{F} \{-i_3 b_{+i_3}^{(1,0)}\}) \\
&= \frac{1}{4} (\mathcal{F} \{a_{+e_t}\} \mathcal{F} \{b_{+i_3}^{(1,0)}\} - \mathcal{F} \{a_{+e_t}\} \mathcal{F} \{b_{+i_3}\} \\
&\quad - \mathcal{F} \{a_{+e_t}^{(0,1)}\} \mathcal{F} \{b_{+i_3}^{(1,0)}\} + \mathcal{F} \{a_{+e_t}^{(0,1)}\} \mathcal{F} \{b_{+i_3}\}), \tag{4.7}
\end{aligned}$$

because

$$\begin{aligned}
\mathcal{F} \{a_{+e_t} i_3\} \mathcal{F} \{-i_3 b_{+i_3}\} &= \mathcal{F} \{a_{+e_t}\} i_3 (-i_3) \mathcal{F} \{b_{+i_3}^{(1,0)}\} = \mathcal{F} \{a_{+e_t}\} \mathcal{F} \{b_{+i_3}^{(1,0)}\}, \\
&\text{etc.} \tag{4.8}
\end{aligned}$$

where we applied (2.18) for the first equality.

By taking the inverse SFT of (4.7) we obtain an identity for expressing a mixed exponential-sine Mustard convolution (4.3) in terms of four standard Mustard convolutions (4.2),

$$\begin{aligned}
&a_{+e_t} \star_{Ms} i_{st} b_{+i_3} \tag{4.9} \\
&= a_{+e_t} \star_M b_{+i_3}^{(1,0)} - a_{+e_t} \star_M b_{+i_3} - a_{+e_t}^{(0,1)} \star_M b_{+i_3}^{(1,0)} + a_{+e_t}^{(0,1)} \star_M b_{+i_3}.
\end{aligned}$$

This now allows us in turn to express the space-time signal convolution purely in terms of standard Mustard convolutions,

$$\begin{aligned}
 a \star b = & \quad (4.10) \\
 & \frac{1}{2} (a_{+e_t} \star_M b_{+i_3}^{(1,0)} + a_{+e_t} \star_M b_{+i_3} - a_{+e_t}^{(0,1)} \star_M b_{+i_3}^{(1,0)} + a_{+e_t}^{(0,1)} \star_M b_{+i_3} \\
 & + a_{+e_t}^{(0,1)} \star_M b_{-i_3}^{(1,0)} + a_{+e_t}^{(0,1)} \star_M b_{-i_3} - a_{+e_t} \star_M b_{-i_3}^{(1,0)} + a_{+e_t} \star_M b_{-i_3} \\
 & + a_{-e_t} \star_M b_{+i_3} + a_{-e_t} \star_M b_{+i_3}^{(1,0)} - a_{-e_t}^{(0,1)} \star_M b_{+i_3} + a_{-e_t}^{(0,1)} \star_M b_{+i_3}^{(1,0)} \\
 & + a_{-e_t}^{(0,1)} \star_M b_{-i_3} + a_{-e_t}^{(0,1)} \star_M b_{-i_3}^{(1,0)} - a_{-e_t} \star_M b_{-i_3} + a_{-e_t} \star_M b_{-i_3}^{(1,0)}).
 \end{aligned}$$

Furthermore, we can combine four pairs of space-time split terms, e.g.,

$$a_{+e_t} \star_M b_{+i_3} + a_{+e_t} \star_M b_{-i_3} = a_{+e_t} \star_M b, \text{ etc.} \quad (4.11)$$

This leaves only twelve terms for expressing a classical convolution in terms of a Mustard convolution,

$$\begin{aligned}
 a \star b = & \\
 & \frac{1}{2} (a_{+e_t} \star_M b_{+i_3}^{(1,0)} + a_{+e_t} \star_M b - a_{+e_t}^{(0,1)} \star_M b_{+i_3}^{(1,0)} + a_{+e_t}^{(0,1)} \star_M b \\
 & + a_{+e_t}^{(0,1)} \star_M b_{-i_3}^{(1,0)} - a_{+e_t} \star_M b_{-i_3}^{(1,0)} \\
 & + a_{-e_t} \star_M b_{+i_3} + a_{-e_t} \star_M b^{(1,0)} - a_{-e_t}^{(0,1)} \star_M b_{+i_3} + a_{-e_t}^{(0,1)} \star_M b^{(1,0)} \\
 & + a_{-e_t}^{(0,1)} \star_M b_{-i_3} - a_{-e_t} \star_M b_{-i_3}). \quad (4.12)
 \end{aligned}$$

Moreover, we can combine with the help of the involution $e_t(i_3)$ of (2.7) four pairs of terms like

$$\begin{aligned}
 a_{+e_t} \star_M b_{+i_3}^{(1,0)} - a_{+e_t} \star_M b_{-i_3}^{(1,0)} &= a_{+e_t} \star_M [b_{+i_3}^{(1,0)} - b_{-i_3}^{(1,0)}] \\
 &= a_{+e_t} \star_M (e_t[b_{+i_3}^{(1,0)} + b_{-i_3}^{(1,0)}]i_3) = a_{+e_t} \star_M e_t b^{(1,0)} i_3, \quad (4.13)
 \end{aligned}$$

where in the final result we omit the round brackets, i.e. we understand $a_{+e_t} \star_M e_t b^{(1,0)} g = a_{+e_t} \star_M (e_t b^{(1,0)} i_3)$. This in turn leaves only eight terms for expressing a classical convolution in terms of Mustard convolutions,

$$\begin{aligned}
 a \star b = & \frac{1}{2} (a_{+e_t} \star_M e_t b^{(1,0)} i_3 + a_{+e_t} \star_M b - a_{+e_t}^{(0,1)} \star_M e_t b^{(1,0)} i_3 + a_{+e_t}^{(0,1)} \star_M b \\
 & + a_{-e_t} \star_M e_t b i_3 + a_{-e_t} \star_M b^{(1,0)} - a_{-e_t}^{(0,1)} \star_M e_t b i_3 + a_{-e_t}^{(0,1)} \star_M b^{(1,0)}). \quad (4.14)
 \end{aligned}$$

Finally, we note, that (4.14) contains pairs of functions $a_{\pm e_t}$ with unreflected and reflected second three-dimensional space vector argument. Adding these pairs leads to even \oplus or odd \ominus symmetry in the second three-dimensional space vector argument. That is, we combine

$$a_{+e_t}^{\oplus} = \frac{1}{2} (a_{+e_t} + a_{+e_t}^{(0,1)}), \quad a_{+e_t}^{\ominus} = \frac{1}{2} (a_{+e_t} - a_{+e_t}^{(0,1)}). \quad (4.15)$$

Remembering the Notation 2.2, the space-time function $a_{+e_t}^{\oplus}$, mapping $\mathbb{R}^{3,1} \rightarrow Cl(3, 1)$, is therefore symmetrized in its three-dimensional space vector argument \vec{x} , whereas $a_{+e_t}^{\ominus}$ is antisymmetrized in its three-dimensional space vector argument \vec{x} .

This finally allows us to write the classical convolution in terms of just four Mustard convolutions.

Theorem 4.5 (Convolution in terms of Mustard convolution). *The convolution (4.1) of two space-time functions $a, b \in L^1(\mathbb{R}^{3,1}; Cl(3, 1))$ can be expressed in terms of four standard Mustard convolutions (4.2) as*

$$a \star b = \frac{1}{2} (a_{+e_t}^{\ominus} \star_M e_t b^{(1,0)} i_3 + a_{+e_t}^{\oplus} \star_M b + a_{-e_t}^{\ominus} \star_M e_t b i_3 + a_{-e_t}^{\oplus} \star_M b^{(1,0)}). \quad (4.16)$$

4.2. Expressing the Mustard convolution in terms of the convolution

Now we will simply write out the Mustard convolution (4.2) and simplify it until only standard convolutions (4.1) remain. In this Subsection we will use the general space-time split of equation (2.5).

We begin by writing the Mustard convolution (4.2) of two space-time functions $a, b \in L^1(\mathbb{R}^{3,1}; Cl(3, 1))$, with space-time vector arguments $\mathbf{x} = t\mathbf{e}_t + \vec{x}$, $\mathbf{y} = t'\mathbf{e}_t + \vec{y}$, and $\mathbf{z} = t''\mathbf{e}_t + \vec{z}$, all in $\mathbb{R}^{3,1}$,

$$\begin{aligned} a \star_M b(\mathbf{x}) &= \frac{1}{(2\pi)^4} \int_{\mathbb{R}^{3,1}} e^{e_t t_1 \omega_t} \mathcal{F}\{a\}(\omega) \mathcal{F}\{b\}(\omega) e^{i_3 \vec{x} \cdot \vec{\omega}} d^2 \omega \\ &= \frac{1}{(2\pi)^4} \int_{\mathbb{R}^{3,1}} e^{-e_t t_1 \omega_t} \int_{\mathbb{R}^{3,1}} e^{-e_t t'_1 \omega_t} a(\mathbf{y}) e^{-i_3 \vec{y} \cdot \vec{\omega}} d^2 \mathbf{y} \\ &\quad \int_{\mathbb{R}^{3,1}} e^{-e_t t''_1 \omega_t} b(\mathbf{z}) e^{-i_3 \vec{z} \cdot \vec{\omega}} d^2 \mathbf{z} e^{i_3 \vec{x} \cdot \vec{\omega}} d^2 \omega \\ &= \frac{1}{(2\pi)^4} \int_{\mathbb{R}^{3,1}} \int_{\mathbb{R}^{3,1}} \int_{\mathbb{R}^{3,1}} e^{e_t (t_1 - t'_1) \omega_t} (a_+(\mathbf{y}) + a_-(\mathbf{y})) e^{-i_3 \vec{y} \cdot \vec{\omega}} \\ &\quad e^{-e_t t''_1 \omega_t} (b_+(\mathbf{z}) + b_-(\mathbf{z})) e^{i_3 (\vec{x} - \vec{z}) \cdot \vec{\omega}} d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega. \end{aligned} \quad (4.17)$$

Next, we use the identities (2.12) in order to shift the inner factor $e^{-i_3 \vec{y} \cdot \vec{\omega}}$ to the left and $e^{-e_t t''_1 \omega_t}$ to the right, respectively. We abbreviate $\int_{\mathbb{R}^{3,1}} \int_{\mathbb{R}^{3,1}} \int_{\mathbb{R}^{3,1}}$ to \iiint .

$$\begin{aligned} a \star_M b(\mathbf{x}) &= \quad (4.18) \\ &= \frac{1}{(2\pi)^4} \iiint e^{e_t (t_1 - t'_1) \omega_t} e^{e_t \vec{y} \cdot \vec{\omega}} a_+(\mathbf{y}) b_+(\mathbf{z}) e^{i_3 t''_1 \omega_t} e^{i_3 (\vec{x} - \vec{z}) \cdot \vec{\omega}} d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\ &+ \frac{1}{(2\pi)^4} \iiint e^{e_t (t_1 - t'_1) \omega_t} e^{e_t \vec{y} \cdot \vec{\omega}} a_+(\mathbf{y}) b_-(\mathbf{z}) e^{-i_3 t''_1 \omega_t} e^{i_3 (\vec{x} - \vec{z}) \cdot \vec{\omega}} d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\ &+ \frac{1}{(2\pi)^4} \iiint e^{e_t (t_1 - t'_1) \omega_t} e^{-e_t \vec{y} \cdot \vec{\omega}} a_-(\mathbf{y}) b_+(\mathbf{z}) e^{i_3 t''_1 \omega_t} e^{i_3 (\vec{x} - \vec{z}) \cdot \vec{\omega}} d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\ &+ \frac{1}{(2\pi)^4} \iiint e^{e_t (t_1 - t'_1) \omega_t} e^{-e_t \vec{y} \cdot \vec{\omega}} a_-(\mathbf{y}) b_-(\mathbf{z}) e^{-i_3 t''_1 \omega_t} e^{i_3 (\vec{x} - \vec{z}) \cdot \vec{\omega}} d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega. \end{aligned}$$

Furthermore, we abbreviate the inner function products as $ab_{\pm\pm}(\mathbf{y}, \mathbf{z}) := a_{\pm}(\mathbf{y})b_{\pm}(\mathbf{z})$, and apply the space-time split of equation (2.5) once again to obtain $ab_{\pm\pm}(\mathbf{y}, \mathbf{z}) = [ab_{\pm\pm}(\mathbf{y}, \mathbf{z})]_{+} + [ab_{\pm\pm}(\mathbf{y}, \mathbf{z})]_{-} = ab_{\pm\pm}(\mathbf{y}, \mathbf{z})_{+} + ab_{\pm\pm}(\mathbf{y}, \mathbf{z})_{-}$. We omit the square brackets and use the convention that the final space-time split indicated by the final \pm index should be performed last.

This allows to further apply (2.12) again in order to shift the factors $e^{\pm i_3 t'_1 \omega_t}$ $e^{i_3(\vec{x}-\vec{z})\cdot\vec{\omega}}$ to the left. We end up with the following eight terms

$$\begin{aligned}
a \star_M b(\mathbf{x}) &= \tag{4.19} \\
&= \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1-t''_1)\omega_t} e^{e_t(\vec{y}-(\vec{x}-\vec{z}))\cdot\vec{\omega}} ab_{++}(\mathbf{y}, \mathbf{z})_+ d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\
&+ \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1+t''_1)\omega_t} e^{e_t(\vec{y}+(\vec{x}-\vec{z}))\cdot\vec{\omega}} ab_{++}(\mathbf{y}, \mathbf{z})_- d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\
&+ \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1+t''_1)\omega_t} e^{e_t(\vec{y}-(\vec{x}-\vec{z}))\cdot\vec{\omega}} ab_{+-}(\mathbf{y}, \mathbf{z})_+ d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\
&+ \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1-t''_1)\omega_t} e^{e_t(\vec{y}+(\vec{x}-\vec{z}))\cdot\vec{\omega}} ab_{+-}(\mathbf{y}, \mathbf{z})_- d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\
&+ \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1-t''_1)\omega_t} e^{e_t(-\vec{y}-(\vec{x}-\vec{z}))\cdot\vec{\omega}} ab_{-+}(\mathbf{y}, \mathbf{z})_+ d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\
&+ \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1+t''_1)\omega_t} e^{e_t(-\vec{y}+(\vec{x}-\vec{z}))\cdot\vec{\omega}} ab_{-+}(\mathbf{y}, \mathbf{z})_- d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\
&+ \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1+t''_1)\omega_t} e^{e_t(-\vec{y}-(\vec{x}-\vec{z}))\cdot\vec{\omega}} ab_{--}(\mathbf{y}, \mathbf{z})_+ d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\
&+ \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1-t''_1)\omega_t} e^{e_t(-\vec{y}+(\vec{x}-\vec{z}))\cdot\vec{\omega}} ab_{--}(\mathbf{y}, \mathbf{z})_- d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega.
\end{aligned}$$

We now only show explicitly how to simplify the second triple integral, the others follow the same pattern.

$$\begin{aligned}
&\frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1+t''_1)\omega_t} e^{e_t(\vec{y}+(\vec{x}-\vec{z}))\cdot\vec{\omega}} [a_+(\mathbf{y})b_+(\mathbf{z})]_- d^2 \mathbf{y} d^2 \mathbf{z} d^2 \omega \\
&= \frac{1}{(2\pi)^4} \iiint e^{e_t(t_1-t'_1+t''_1)\omega_t} d\omega_t \int_{\mathbb{R}^3} e^{e_t(\vec{y}+(\vec{x}-\vec{z}))\cdot\vec{\omega}} d\vec{\omega} [a_+(\mathbf{y})b_+(\mathbf{z})]_- d^2 \mathbf{y} d^2 \mathbf{z} \\
&= \iint \delta(t_1-t'_1+t''_1) \delta(\vec{y}+(\vec{x}-\vec{z})) [a_+(\mathbf{y})b_+(t''_1, \vec{z})]_- d^2 \mathbf{y} d^2 \mathbf{z} \\
&= \int_{\mathbb{R}^2} [a_+(\mathbf{y})b_+(-(t_1-t'_1), \vec{x}+\vec{y})]_- d^2 \mathbf{y} \\
&= \int_{\mathbb{R}^2} [a_+(\mathbf{y})b_+(-(t_1-t'_1), -(\vec{x}-\vec{y}))]_- d^2 \mathbf{y} \\
&= \int_{\mathbb{R}^2} [a_+(\mathbf{y})b_+^{(1,1)}(t_1-t'_1, -\vec{x}-\vec{y})]_- d^2 \mathbf{y} \\
&= [a_+ \star b_+^{(1,1)}(t_1, -\vec{x})]_-. \tag{4.20}
\end{aligned}$$

Note that $a_+ \star b_+^{(1,1)}(t_1, -\vec{x})$ means to first apply the convolution to the pair of functions a_+ and $b_+^{(1,1)}$, and only then to evaluate them with the argument $(t_1, -\vec{x})$. So in general $a_+ \star b_+^{(1,1)}(t_1, -\vec{x}) \neq a_+ \star b_+(-t_1, \vec{x})$. Simplifying the other seven triple integrals similarly we finally obtain the desired decomposition of the Mustard convolution (4.2) in terms of the classical convolution.

Theorem 4.6 (Mustard convolution in terms of standard convolution). *The Mustard convolution (4.2) of two space-time functions $a, b \in L^1(\mathbb{R}^{3,1}; Cl(3, 1))$ can be expressed in terms of eight standard convolutions (4.1) as*

$$\begin{aligned}
 a \star_M b(\mathbf{x}) &= \\
 &= [a_+ \star b_+(\mathbf{x})]_+ + [a_+ \star b_+^{(1,1)}(t_1, -\vec{x})]_- \\
 &\quad + [a_+ \star b_-^{(1,0)}(\mathbf{x})]_+ + [a_+ \star b_-^{(0,1)}(t_1, -\vec{x})]_- \\
 &\quad + [a_- \star b_+^{(0,1)}(t_1, -\vec{x})]_+ + [a_- \star b_+^{(1,0)}(\mathbf{x})]_- \\
 &\quad + [a_- \star b_-^{(1,1)}(t_1, -\vec{x})]_+ + [a_- \star b_-(\mathbf{x})]_-. \tag{4.21}
 \end{aligned}$$

Remark 4.7. *If we would explicitly insert according to (2.12) $a_\pm = \frac{1}{2}(a \pm e_t a i_3)$ and $b_\pm = \frac{1}{2}(b \pm e_t b i_3)$, and similarly explicitly insert the second level space-time split $[\dots]_\pm$, we would obtain up to a maximum of 64 terms. It is therefore obvious how significant and efficient the use of the space-time split (2.5) is in this context.*

Remark 4.8. *The steerability (compare Remark 3.1) of the Mustard convolution (4.2) is seen in Theorem 4.6 in the explicit occurrence of the algebra of space-time split (2.5).*

5. Conclusion

We have introduced the Clifford algebra $Cl(3, 1)$ for space-time $\mathbb{R}^{3,1}$ together with the space-time split, based on the time vector e_t and its dual three-dimensional space volume pseudoscalar $e_t^* = i_3$. In this context we looked in detail at a number of involutions in $Cl(3, 1)$ connected with e_t , i_3 and their product, the space-time hypervolume pseudoscalar $i_{st} = e_t i_3$. Next, we briefly reviewed for space-time Clifford algebra $Cl(3, 1)$ valued signals over $\mathbb{R}^{3,1}$ the steerable space-time Fourier transform, and defined a pair of related exponential-sine type Fourier transforms. This was followed by definitions of the (classical) convolution for space-time signals and two types of steerable Mustard convolutions (with point wise products in the spectral domain). Finally we expressed the convolution in terms of Mustard convolutions (Theorems 4.3 and 4.5), and vice versa the Mustard convolution in terms of classical convolutions in Theorem 4.6.

We expect our results to be relevant for applied mathematics, physics, engineering and navigation, in particular for special relativistic quantum mechanics, optics, electro-dynamics and aero-space navigation. Furthermore, we expect applications in electromagnetic signal transmission and processing. In the convolutions one signal function could be an electromagnetic signal, the other a filter function, window function, continuous mother wavelet, etc.

Acknowledgments

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²At the occasion of John Ryan's 60th birthday it may be in order to reflect about the interdisciplinary context of mathematical research, including life sciences and theology, and thus to put the above Bible quotation in context: *Natural selection has shown insidious imperialistic tendencies. The offering of post-hoc explanations of phenotypic traits by reference to their hypothetical effects on fitness in their hypothetical environments of selection has spread from evolutionary theory to a host of other traditional disciplines: philosophy, psychology, anthropology, sociology, and even to aesthetics and theology. Some people really do seem to think that natural selection is a universal acid, and that nothing can resist its powers of dissolution. However, the internal evidence to back this imperialistic selectionism strikes us as very thin. Its credibility depends largely on the reflected glamour of natural selection which biology proper is said to legitimise. Accordingly, if natural selection disappears from biology, its offshoots in other fields seem likely to disappear as well. This is an outcome much to be desired since, more often than not, these offshoots have proved to be not just post hoc but ad hoc, crude, reductionist, scientistic rather than scientific, shamelessly self-congratulatory, and so wanting in detail that they are bound to accommodate the data, however that data may turn out. So it really does matter whether natural selection is true.*[5]

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