

Plausibility of Earth Once Having a Thick Atmosphere – Examining the Rate of Impact Cratering

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Abstract

Theories abound as to how dinosaurs and other prehistoric creatures could have grown to such immense sizes, inconsistent with the spectrum of sizes for today's creatures and Earth's living conditions. Some focus directly on changes in the governing physics of the universe, such as a different gravitational constant. Some postulate that, rather than this difference, the earlier Earth experienced lower gravity due to differences in its size and mass. The majority focus on biological and aerodynamical anomalies that may have prevailed to explain these gargantuan sizes. This paper focuses on the latter group, offering an independent means by which to test the hypothesis that a (much) thicker atmosphere provided the buoyancy needed by these creatures to exist on land. This means is astronomical, an examination of possible differences in the rate of impact cratering on Earth due to atmospheric differences. With the Earth's atmosphere allegedly experiencing eras of much greater thickness than current, and alternating between these "thick" and "thin" atmospheric eras, it is postulated that, in addition to the biological and aerodynamical anomalies, a difference in the cratering rate from meteor impacts on Earth should be evident. Thicker atmosphere would "burn up" more meteors, reducing the cratering rate when compared to that during thinner atmospheric eras. This paper explores this, using the cratering rate from meteor impacts on the Moon as a "control" since it has no atmosphere to attenuate meteors but also is in Earth's orbital vicinity and should have experienced a nearly equivalent rate of meteor influx per unit surface area.

Key Words: Earth, Moon, Atmosphere, Volcanism, Carbon Dioxide, Dinosaurs, Meteors, Impact
Cratering

1. Introduction

Some dinosaurs (and other prehistoric “leviathans”) were inexplicably large, especially in light of today’s spectrum of creature sizes. Various theories to “explain” how they could have functioned given such sizes have been postulated. Some focus on postulates that the gravitational constant was lower, such that Earth’s gravity would have been lower, or a varying size of the Earth may explain the paradox. Others pursue biological arguments, with connections to aerodynamics, for an explanation. We will not consider the first set, but rather focus on the second as being the more plausible. After reviewing the arguments for the biological/aerodynamical postulates, we examine an independent means of ascertaining the plausibility of these, both of which contend that Earth had a much thicker atmosphere in the past. For that independent means, we select an astronomical approach, namely examination of possible differences in the cratering rates due to meteor impacts on the Earth during “thicker” and “thinner” atmosphere eras, representing eras of greater and lesser attenuation (“burn up”) of incoming meteors, thereby affecting the cratering rate per unit surface area on Earth relative to what has been experienced on the geologically and climatologically dead Moon. Since the Moon is in the same orbital neighborhood as the Earth, it should have experienced the same meteor influx per unit surface area over the same eras.

Much of the material in Section 2, especially regarding dinosaur physiology, is provided only as background to the thick atmosphere theories, i.e., this material is not necessarily used in the analysis for cratering rates due to meteor impacts. The reader interested only in the latter may skip to the last paragraph in Section 2.

2. Two Prominent Theories for Thick Earth Atmosphere

Two prominent theories supporting the proposition that Earth has previously experienced (much) thicker atmospheric conditions are examined. Both focus on biological and aerodynamic arguments regarding dinosaur and other prehistoric creatures having sizes incongruously large when viewed in terms of how they could possibly exist today.

2.1. Levenspiel, Fitzgerald and Pettit

In “Earth’s Atmosphere before the Age of Dinosaurs,” Levenspiel, Fitzgerald and Pettit state: [1]

... [T]he giant flying creatures of the dinosaur age could only fly if the atmospheric pressure was much higher than it is now: at least 3.7–5.0 bar. If this is so, it raises several interesting questions. For example, how did the atmosphere get to that pressure 100–65 million years ago (Mya)? What was the pressure before that? And how did it drop down to today’s 1 bar? Although we have no definite answers to these questions, let us put forth reasonable possible explanations.

What was the air pressure for the 97% of Earth’s life before the age of dinosaurs? We have three possible alternatives, as shown in Figure 1.¹

- The pressure could have been at 1 bar throughout Earth’s earlier life, risen to 4–5 bar ~100 Mya (just at the time when the giant fliers needed it), and then returned to 1 bar (curve A).

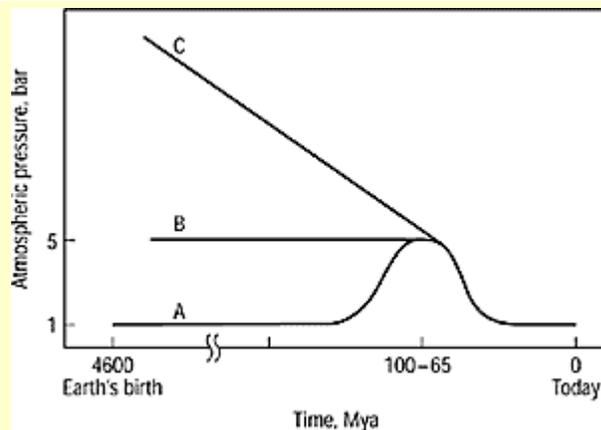


Figure 1. Three possible alternatives for the atmospheric pressure early in Earth’s lifetime, given that it was at ~5 bar, ~100 Mya.

¹ Figure 1, and the review of the thicker atmosphere theory of Levenspiel, Fitzgerald and Pettit are presented only to show that there are multiple analysts presenting theories of prehistorically thicker atmospheres. It is not used in the subsequent analysis, which focuses solely on Esker’s thicker atmosphere theory.

- *The pressure could have been ~4–5 bar from Earth’s beginning, 4600 Mya; and ~65 Mya, it could have begun to come down to today’s 1 bar (curve B).*
- *The atmosphere could have started at higher pressure and then decreased continuously through Earth’s life to ~4–5 bar ~100 Mya and down to 1 bar today (curve C).*

The third alternative seems to be the most reasonable ...Geologists believe that most of the carbon on the young, hot Earth, >4000 Mya, was in the form of gaseous carbon dioxide, carbon monoxide, and methane. With time, the CO and CH₄ reacted with oxide minerals and were transformed into CO₂. These reactions did not change the total amount of carbon in the atmosphere.

Our sister planet and nearest neighbor, Venus, has an atmosphere of 90 bar pressure, consisting of 96% CO₂ (5). Why should Earth be so different? ... [W]hy did Venus’s atmosphere remain at 90 bar while Earth’s decreased to a few bar during the age of dinosaurs and then declined to the 1 bar it is today? What happened to Earth’s CO₂ and by what mechanism did it virtually disappear? ... Being thinner, Earth’s crust was fragile and broke up under the action of the mantle’s convective forces. In contrast, Venus’s thicker crust remained rigid and did not permit the mechanisms that removed the CO₂ from its bound state. In addition, because Venus is closer to the Sun and hotter than Earth, free liquid water cannot exist on it, whereas Earth has giant oceans that cover two-thirds of the planet. The oceans played an important secondary role in removing CO₂ from the atmosphere ...

Today, vast deposits of sedimentary carbonate rocks are found on land and on ocean bottoms, >1,000,000 km³ throughout Earth’s crust. Above the continents, the CO₂ was taken up by rainwater and by groundwater. This CO₂-rich water reacted with rocks to form bicarbonates, followed by transport to the ocean and precipitation as calcium and magnesium carbonates. In the ocean, dissolved CO₂ combined with the calcium hydroxide to form deposits of chalk, or it was taken up by coral, mollusks, and other living creatures to form giant reefs. A study of the distribution through time of these deposits gives us clues to the history of CO₂ in the atmosphere ...

With time, the concentration of CO₂ steadily decreased, primarily because of the formation and deposition of limestone and other carbonaceous materials. CO₂ was also lost by photosynthesis followed by the deposition of carbonaceous substances such as coal, petroleum, peat, oil shale, and tar sands; however, this loss was quite minor. Calculations show that the deposit of what are now considered fuel reserves lowered the atmospheric CO₂ by <<1 bar. At the same time, the concentration of oxygen slowly rose. These two changes, the decrease in CO₂ and the rise in oxygen, thinned the forests and the dead material began to

be oxidized more rapidly, so that dense layers of dead organics were no longer deposited. Evidence of this change in atmospheric conditions is that we cannot find any massive coal deposits younger than 65 million years. Animal life found this changed atmosphere to its liking, so mammals and dinosaurs flourished, first as very small creatures but then increasing in size as a result of evolutionary competition. This led to the giant flying creatures close to the end of the dinosaur age. It could be that these creatures died out as the total pressure of the atmosphere dropped below their sustainable level ...

If we assume that Earth's early atmosphere was very different, both in composition (mainly CO₂) and total pressure, that would answer some puzzling questions from a variety of disciplines.

- *How did the flying creatures from the age of dinosaurs have enough energy to fly when physiology, biology, and aeronautics say that this was impossible?*
- *How could life have developed on Earth when astronomy says that Earth was too cold to sustain life?*
- *If Earth's atmosphere had stayed at ~1 bar throughout its history, where did the equivalent of 50–70 bar of CO₂ in limestone and other carbonates on Earth's surface come from?*

This picture of high CO₂ concentration and high pressure in the past also explains why most massive coal seams are older than 65 million years and why most limestone caves are younger than 100 million years. Although we do not know the values for the atmospheric pressure in those early times, and although each of the arguments in this paper only leads to suggestions, when taken together, the evidence from these various sources leads to the same conclusion: The atmospheric pressure was higher in the past than it is today and consisted primarily of CO₂. This hypothesis presents a picture of our evolving planet that should be examined and that could have interesting consequences.

2.2. Esker

In a subsequent, more comprehensive look at this topic, “Scientific Theory Solving the Dinosaur Paradox and Numerous Other Paradoxes Regarding Earth's Evolution,” Esker states: [2]

... [T]he large dinosaurs and pterosaurs of the Mesozoic era present a scientific paradox. Four areas of scientific incongruities regarding these animals' large size are identified: 1) insufficient muscle strength, 2) insufficient bone strength, 3) unacceptably high blood pressure within the tallest dinosaurs, and 4) the

paradox of pterosaurs having grossly insufficient power to fly in atmospheric conditions similar to the present ... [T]he development of airplanes has always been more of an art than a science. The absence of a theoretical understanding of flight becomes most apparent when the paleontologists make their foolish attempts trying to explain how the giant pterosaurs flew. Common sense tells everyone that a reptile the size of a horse should not be capable of flight, but until now there has not been a theoretical understanding of flight enabling us to scientifically clarify what is wrong with the paleontologists' claim that there is nothing odd about gigantic flying reptiles ... The Thick Atmosphere Solution's ability to solve the dinosaur paradox qualifies it as being a strong hypothesis, but with additional evidence it can be shown that the Thick Atmosphere Solution is actually a new scientific theory ... [T]he Thick Atmosphere Theory solves the long-standing paleoclimatologist puzzle of how the Mesozoic era Earth had the same pleasant climate over its entire surface ...

Just as the largest animals have the lowest relative bone strength, it is also true that the largest animals have the lowest relative muscle strength. Absolute strength can be defined as how much weight an animal can lift regardless of the animal's own weight, and clearly the larger animals have greater absolute strength than the smaller animals. But when we look at relative strength, the lifting ability of an animal relative to its own weight, it is the smallest animals that have the greatest relative strength ... For most physically fit human beings we have more than enough relative strength so that getting out of bed in the morning is not outside our physical capacity. But the larger animals that have lower relative strength lifting their body off the ground can be a serious issue. Large farm animals such as cattle or horses exert all the strength that they have when they pick themselves up off the ground. Likewise the large wild animals such as elephants and giraffes need all their strength to perform this task that is not challenging for the smaller animals. As a consequence of these difficulties, it is not surprising that many of these larger animals evolved the behavior of sleeping while standing up. Yet numerous dinosaurs were much larger than these animals. Their greater size would mean that their relative strength would be substantially less than that of the large animals of today. It is not realistic to imagine that the large dinosaurs never fell or otherwise found themselves on the ground throughout their entire lives. If a Jurassic Park was actually created, any sauropod or other large dinosaur would be stuck lying on the ground much like a helpless whale stranded on a beach ...

The buoyancy force is best described by Archimedes' principle that states that when an object is partially or fully submerged in a fluid, an upward buoyancy force lifts up on the submerged object that exactly equals the weight of the fluid displaced. ... [B]uoyancy ... is what gives a lifting force to hot air balloons. The main

difference in the buoyancy effect provided by these two fluids [air vs. water] is the amount of fluid volume that needs to be displaced to achieve flotation. For terrestrial vertebrates, it is the net force produced by their weight that often limits their size. But this is not true for species that exist in the water. For the latter species it is not their weight but rather other factors, such as the availability of food that might limit the size of these species. Without the weight limitation some of these aquatic species grow to display gigantism. It is the buoyancy of water that allows the whales, the largest animals of today, to grow so large ... Without this buoyancy to counteract gravity, the poor whale that finds itself stuck on a beach is soon having its bones broken from its own weight. To produce an effective buoyancy force on dinosaurs the Earth's atmosphere would have to be thick enough to have a density comparable to the density of water. By summing the forces acting on a typical dinosaur such as a Brachiosaurus the density of the necessary atmosphere is calculated ... to be 670 kg/m³. This says that to produce the necessary buoyancy so that the dinosaurs could grow to their exceptional size,² the density of the Earth's air near the Earth's surface would need to be 2/3's of the density of water ...

It may be hard to imagine that the Earth's air could be so thick that its density would be comparable to water. Nevertheless, there is no reason why a gas cannot be compressed so much that it has properties similar to that of a liquid, and in fact compressing a gas into a liquid is a common industrial process ... 150 million years ago the Earth's atmospheric pressure near the surface was about 370 atmospheres ... 370 times thicker than what it is today ... [C]onsider the pressure that currently exist at the deepest depths of the oceans. The average ocean depth is 3790 m and at this depth the pressure is 380 atmospheres. So for all practical purposes, the present day pressure at the average depth of the ocean is the same as the pressure at the bottom of the Mesozoic atmosphere. Yet there are numerous species that live at this depth and many more that live much deeper. Extremely high absolute pressure has no ill effect on our present creatures of the deep that have evolved in these environments; likewise, the extremely high pressure of the Mesozoic era had no ill effect on the terrestrial species of the Mesozoic era ... If both the inside and outside of an enclosed container are at the same absolute pressure, no matter what the absolute pressure might be, there will be no net force on the sides of the container ...

² Esker's discussion makes it clear that the buoyancy provided by a thicker atmosphere benefitted not only flying dinosaurs (pterosaurs) but also those that walked on land.

Within the Phanerozoic eon [current geologic eon ... during which abundant animal and plant life has existed – 541 million years to the present] we can identify two thick atmosphere eras and two thin atmosphere eras ... Twice during the Carboniferous and the Cretaceous/Paleogene periods, the atmosphere transitioned from being extremely thick to being relatively thin ... With a massive amount of CO₂ being removed from the atmosphere we would expect to see large carbon deposits during these times and indeed that is the case ... [T]he only time that the atmosphere transitioned from being relatively thin to being extremely thick was when the earth was void of most life ... around the time of the P-T [Permian-Triassic] mass extinctions and continuing into the Triassic period ...

Figure 2 is a linearized approximation of Esker's graph of "Atmospheric Levels during the Last 350 Million Years," on which I have arbitrarily drawn transition times between the two Thick and Thin Atmosphere Eras using an arbitrary transition atmosphere of 200 atm. Starting around 350 million years ago with an atmospheric thickness of nearly 500 atm, he presents alternating periods of decreasing and increasing atmospheric pressure up to today's present "Thin" atmosphere, which I have assumed to be "Thick" and "Thin" as shown in my approximation of Esker's figure. This results in two Thin and Thick Eras, as shown. They transition at approximately 340, 230 and 53 million years ago, with the Thick1 Era assumed to begin 2.4 billion years ago, since this is the reported age of the oldest recorded Earth crater, the 16-km Suavjärvi crater in Asia (see Table I).

3. Thick Atmosphere Theory and Earth Cratering Rates

The previous discussions by Levenspiel, et al., and Esker supporting a Thick Atmosphere Theory focus on mainly biological and aerodynamic arguments. After reading these discussions, I seek an independent means by which to examine this theory at least for plausibility, as anything definitive is currently beyond achieving. Reasoning that a thicker atmosphere should "burn up" more incoming meteors than a thinner one, I examine the cratering rate for impacting meteors on the Earth, based on the Earth Impact Database

(see Table I).³ From the Earth Impact Database I compile a list of all Earth craters from meteoric impacts that have been recorded (including some still cited as “unconfirmed” [*red italics*]). For reasons that will become evident, only craters at least 4 km in size are counted. To the present 163 such craters have been identified, which reduces to 111 if only those at least 10,000 years old are counted (roughly up to the end of the last Ice Age). Note that this affects only the last Esker Era, labelled as Thin2. This somewhat arbitrary truncation results from the preponderance of North American craters of most recent age relative to similar craters worldwide. The intent is to remove possible bias from more extensive crater identification having been performed on our continent.

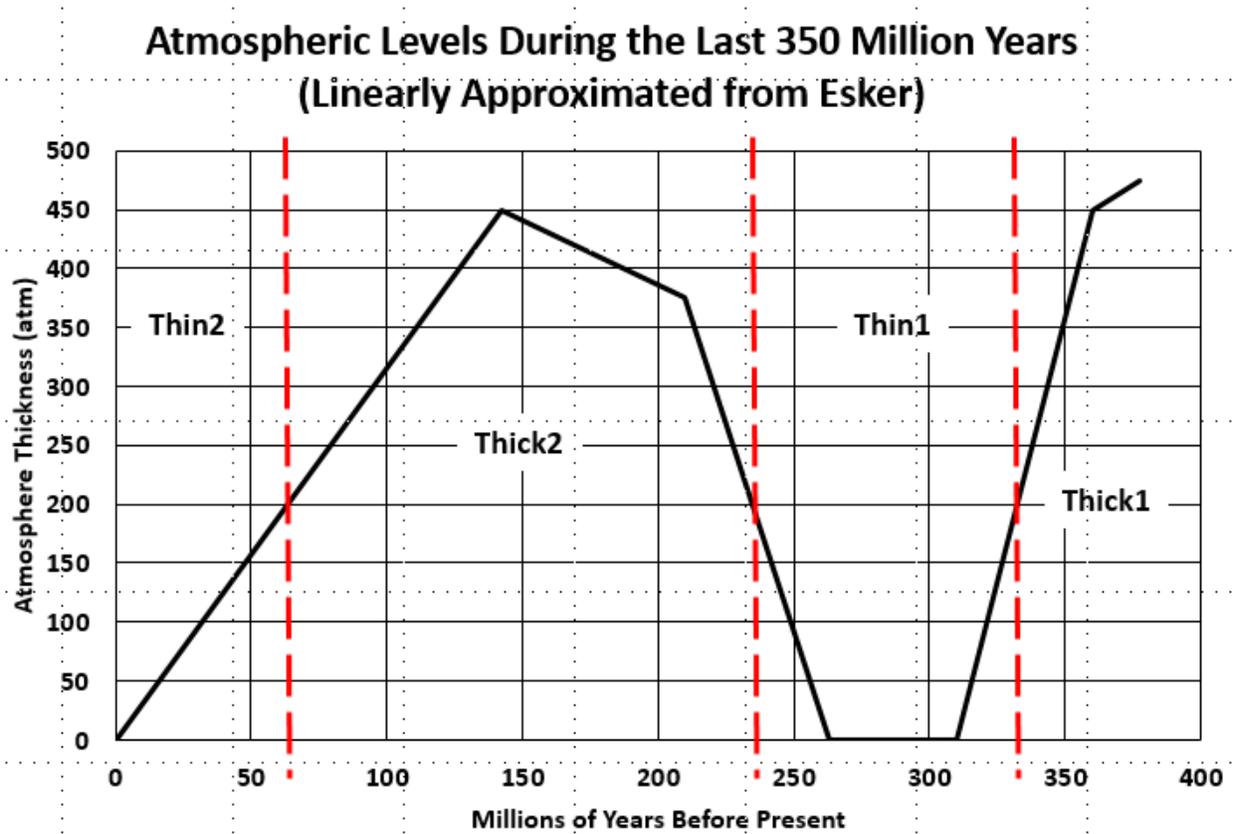


Figure 2. Atmospheric Levels during the Last 350 Million Years with Assumed Transition Times

³ “List of Impact Craters on Earth,” [Earth Impact Database](https://en.wikipedia.org/wiki/List_of_impact_craters_on_earth) (available at https://en.wikipedia.org/wiki/List_of_impact_craters_on_earth).

Before proceeding, it is important to ascertain the time history of what the cratering rate would have been for the Earth in the absence of an atmosphere, its geologic activity, etc. This may be possible by assuming the time history of the Moon's cratering rate would be closely representative, on a per unit area, given its proximity to the Earth. Figure 3 presents an estimate of the lunar cratering rate over the assumed roughly five-billion-year lifespan of the Moon. [3] Corresponding to the four Esker Atmospheric Eras is this figure showing the estimated rate of cratering on the Moon since its alleged birth in terms of the rate per unit surface area (km^2) for craters > 4 km in size. Table II shows the starting and finishing times for each of the Esker Eras, with the corresponding cratering rates at the start and finish of each based on the "constant production rate" curve (dashed). For each Era, the geometric mean (given the logarithmic plot) between the starting and finishing rates is assumed to be characteristic for that Era. For example, for Thick1, the geometric mean is just the square root of the cratering rates at the start and finish, i.e., $\sqrt{(8.5 \times 10^{-5} \text{km}^{-2})(1.1 \times 10^{-5} \text{km}^{-2})} = 3.06 \times 10^{-5} \text{km}^{-2}$. Consistent with the curve, this decreases with time, from $\sim 3\text{E-}5/\text{km}^2$ during the earlier Thick1 Era down to $\sim 2\text{E-}6/\text{km}^2$ for the present Thin 2 Era, slightly over a factor of 10. When the time-weighted rates for both Thick and Thin Eras are calculated, we see that the weighted cratering rate for the Thin Eras is about one-quarter of that for the Thick ones.⁴ This is expected given the Thick Eras always precede the Thin ones, such that their cratering rates are relatively higher, and the cumulative time periods for the Thick Eras (~ 2.2 billion years) is over 10 times longer than for the Thin ones (~ 160 million years).

For each of the Esker Eras, I estimate the cratering rate on Earth (for craters at least 4 km in size, to place on an equivalent basis for comparison with the Moon) as the number of craters identified for that Era divided by the length of the Era and the $\sim 29\%$ of the surface area of the Earth that is land ($[0.29 \times 4\pi \times 6371 \text{km}]^2 = 1.5 \times 10^8 \text{km}^2$). This is evaluated on an annual basis, e.g., for Thick1 to the end of the Ice Age:

⁴ Weighting over the two Thick and two Thin Atmospheric Eras is accomplished as follows (shown for the Thick Eras – it is analogous for the Thin Eras):

$$\frac{(3.06 \times 10^{-5} \text{km}^{-2})(2.4 \times 10^9 \text{y} - 3.4 \times 10^8 \text{y}) + (5.05 \times 10^{-6} \text{km}^{-2})(2.3 \times 10^8 \text{y} - 5.3 \times 10^7 \text{y})}{(2.4 \times 10^9 \text{y} - 3.4 \times 10^8 \text{y} + 2.3 \times 10^8 \text{y} - 5.3 \times 10^7 \text{y})} = 2.85 \times 10^{-5} \text{km}^{-2}$$

$$^{43}/(1.5 \times 10^8 \text{ km}^2)(2.4 \times 10^9 \text{ y} - 3.4 \times 10^8 \text{ y}) = 5.63 \times 10^{-16} \text{ y}^{-1} \text{ km}^{-2}. \quad [1]$$

Then I weight over the two Thick and Thin Eras, as shown.⁵ Table III presents two sets of estimates, one where I truncate the counting of Earth craters at the end of the last Ice Age (10,000 years ago) and one without truncation (i.e., counting all craters to present time). This has no effect on the cratering rate for the Thick Eras ($4.87 \times 10^{-7} \text{ km}^{-2}$), which is a factor of 59 lower than the corresponding lunar cratering rate ($2.85 \times 10^{-5} \text{ km}^{-2}$), to be expected given Earth's active climate and geology. The cratering rates for the Thin Eras vary by about a factor of 2.5, being lower when truncated at the end of the Ice Age ($2.64 \times 10^{-7} \text{ km}^{-2}$ vs. $6.15 \times 10^{-7} \text{ km}^{-2}$). Both are lower than the lunar cratering rate for the corresponding Thin Eras ($6.99 \times 10^{-6} \text{ km}^{-2}$), as would be expected, but notably not lower by as high a factor when compared to the Thick Eras (~59 for the Thick Eras, but around 26 and 11 for the Thin Eras).

What is of particular interest is the ratio of the weighted cratering rates (*red italics* in Table III). When truncated at the end of the Ice Age, the cratering rate during the Thin Eras is reduced by nearly a factor of two relative to that for the Thick Era, somewhat to be expected given the lunar result which showed roughly a factor of four reduction. The fact that the Earth cratering rate during the Thin Eras is reduced by less compared to the Moon rate may be indicative of the effect of atmospheric thickness. That is, the thinner Earth atmosphere allowed more cratering during the Thin Eras than would be expected relative to the cratering rate during the Thick Eras when compared to the ratio for the Moon which is climatically and geologically dead (compare ratios of 0.542 to 0.245 [*red italics* in Table II]). If the Earth crater counting is not truncated, i.e., counted to present time, this difference is much more pronounced. In fact, the cratering rate during the Thin Eras now is slightly higher than during the Thick Eras, by about one quarter (ratio = 1.26, in red italics). Figure 4 shows this graphically by the three different trend lines (solid red for the

⁵ This weighting is slightly different from that used for the lunar rates, as follows, e.g., for the two Thick Eras:

$$(1.41 \times 10^{-16} \text{ y}^{-1} \text{ km}^{-2})(2.4 \times 10^9 \text{ y} - 3.4 \times 10^8 \text{ y}) + (1.10 \times 10^{-15} \text{ y}^{-1} \text{ km}^{-2})(2.3 \times 10^8 \text{ y} - 5.3 \times 10^7 \text{ y}) = 4.87 \times 10^{-7} \text{ km}^{-2}.$$

Note that this is the same as combining the two Thick Eras initially:

$$^{(43 + 29)}/(1.5 \times 10^8 \text{ km}^2) = 4.87 \times 10^{-7} \text{ km}^{-2}.$$

Moon; dashed green for the Earth to the Ice Age; and dotted blue for the Earth to Present). The lunar trend line is the steepest downward. That for the Earth to the Ice Age is also downward, but not as steep, while the trend line for the Earth to Present is slightly upward.

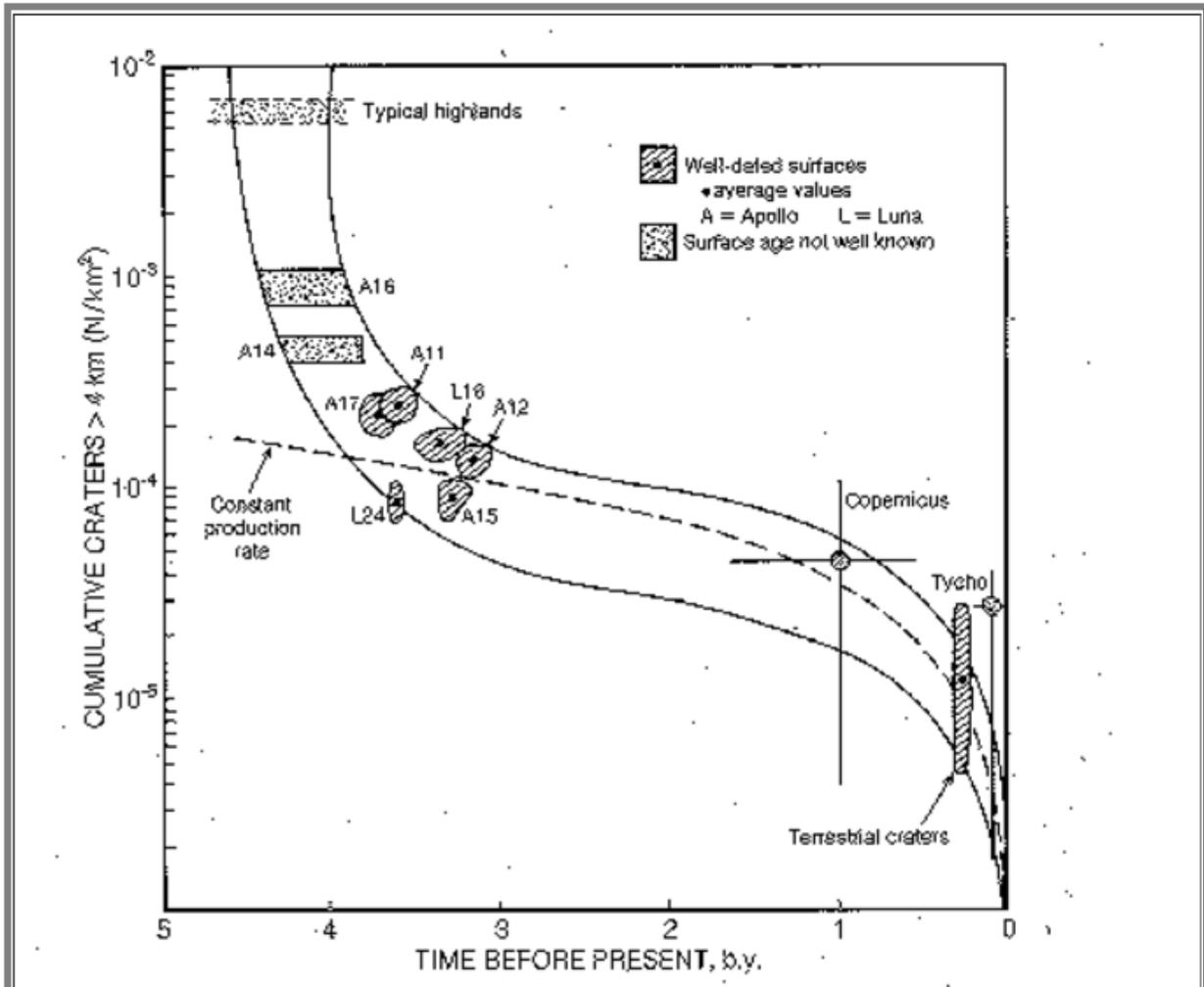


Figure 1. Lunar crater production rates through geologic time as reconstructed from the measurement of crater densities on the lunar surface and from absolute age dating of returned lunar rocks. Firm correlations can only be reconstructed for (1) the well-characterized basalt surfaces (3.8-3.2 Gyr) and (2) the contemporary meteorite flux based on current astronomical observations ($t = 0$). The ages of Tycho and Copernicus are inferred from indirect evidence. From F. Horz et al. p.84.

Figure 3. Lunar Crater Production Rates through Geologic Time

The Earth to Present trend is completely different from the lunar, which saw a reduction by about a factor of four rather than this increase by one quarter. This may be indicative even more so of the atmospheric thickness effect, although the caveat previously mentioned about the preponderance of the most recent craters having been identified in North America somewhat tempers it. Nonetheless, even the comparison for truncation at the end of the Ice Age shows a noticeable difference relative to what would be expected for a body without an atmosphere subjected to the same meteor influx, represented by the Moon. Another factor, though likely not as dominant as the potential atmospheric effect, could be a decreasing geologic activity on Earth with time, since the Thick Eras each preceded the Thin Eras. However, given Earth is still quite geologically active, likely not much less so than around two billion years ago, this effect is expected to be dwarfed by the atmospheric thickness difference.

4. Summary

Given all the assumptions and approximations employed, and the fidelity of cratering data for both the Earth and Moon, no definitive conclusion can be drawn. However, at least this cratering rate analysis does not contradict the postulate that Earth's atmosphere has varied substantially in thickness as per Esker and offers an independent means to test the hypothesis to supplement the more biological and aerodynamic ones that both he and Levenspiel, et al., provide. During the Thick Atmosphere Era, meteor impact on the Earth would be decreased by a relatively greater degree vs. the Thin Atmosphere Era when compared to what would be expected on a per unit surface area for the geologically and climatologically dead Moon. Given two meteors of comparable size, speed and entry angle, the one hitting the thick atmosphere would be less likely to survive to impact than the one hitting the thin atmosphere on Earth.

This has been demonstrated by the analysis presented here, which considers two scenarios, varying with the truncation time for the cratering rates. The first truncates at the end of the Ice Age; the second does not truncate, but extends to the present.

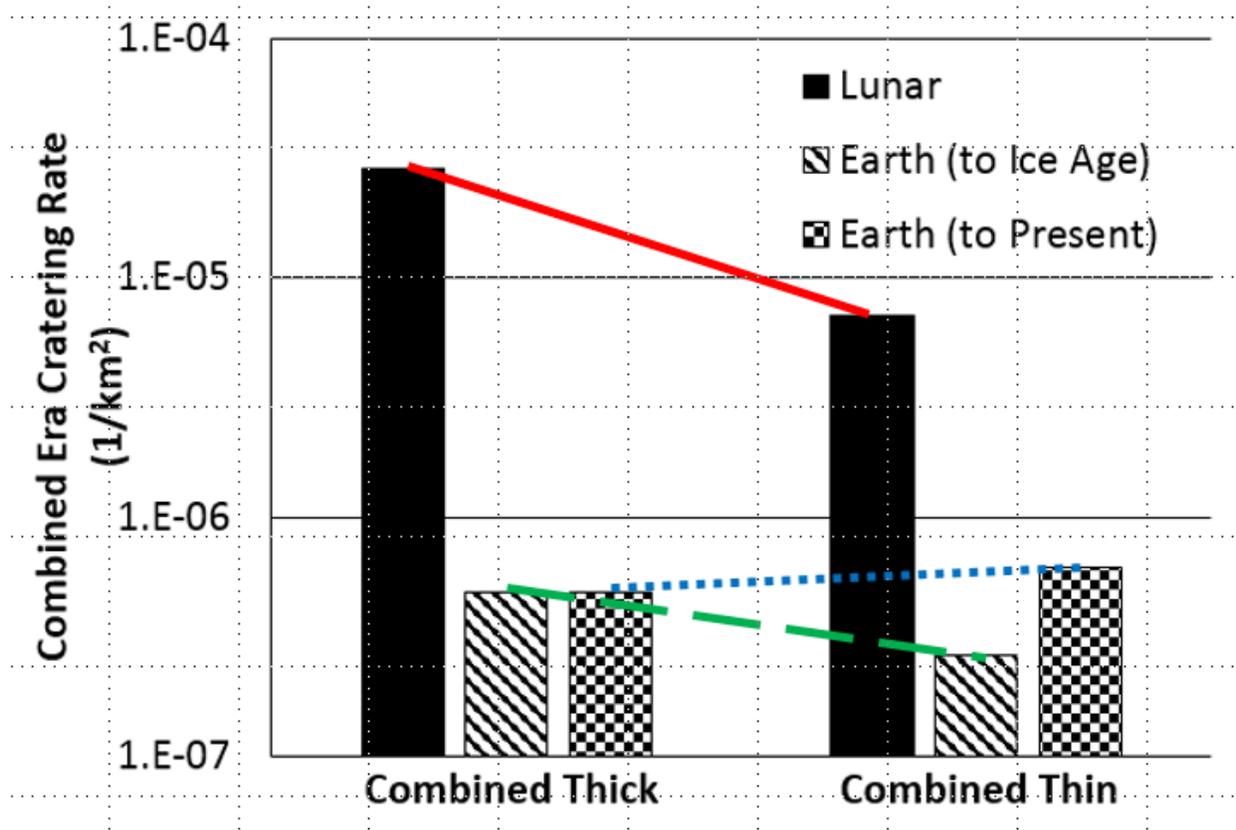


Figure 4. Comparing Trends in Cratering Rates for the Earth and Moon over the Combined Thick and Thin Atmosphere Eras

For the first scenario, the ratio of lunar cratering rate between the Thin and Thick Eras (as defined for Earth) is 0.245, indicating that the Moon, without an atmosphere, experienced roughly four times the cratering rate during the Thick Eras vs. the Thin Eras ($1/0.245 \approx 4$). For truncation at the end of the Ice Age, the corresponding ratio for the Earth during these same Eras is 0.542, slightly more than twice as high ($1/0.542 \approx 2$). This indicates that, on Earth, the cratering rate during the Thick Eras was slightly less than twice that during the Thin Eras. Therefore, the effect of Earth's Thicker vs. Thinner Atmosphere Eras was to reduce the cratering rate more during the Thick Eras than the Thin Eras relative to what the reduction would have been without an atmosphere, as evidenced by the Earth's higher Thin vs. Thick ratio relative to that for the Moon (0.542 vs. 0.245). That is, instead of exhibiting a Thick Era cratering rate four times as high as that for the Thin Era, as per the Moon without an atmosphere, the Earth exhibited a rate

only twice as high during the Thick vs. Thin Era. This ratio difference supports the conjecture that thicker atmosphere reduces cratering rate.

For the second scenario, this tendency is even more pronounced for the second scenario when the cratering rate is not truncated at the end of the Ice Age but extended to the present. Now the ratio between the Thin and Thick Eras on Earth is 1.26, indicating a Thin Era cratering rate 26% higher than during the Thick Era. Contrasting against the Moon's ratio of 0.245, one sees a pronounced decrease during the Thick Era relative to the Thin Era on Earth vs. what would have been experienced without an atmosphere, as evidenced by the Moon. Again, this ratio difference supports the conjecture that a thicker atmosphere reduces cratering rate, aligning with Esker's conjecture, which is based on biological/aerodynamical arguments.

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3. “Lunar Crater Production Rates through Geologic Time” (available at <http://muller.lbl.gov/pages/crateringrates.htm>).

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Table I. Complete List of Identified Earth Craters

#	Name/Location	Continent	Size (km)	Age (yr)
1	Suavjärvi	Asia	16	2.4E+09
2	Vredefort	Africa	300	2.0E+09
3	Yarrabubba	Australia	30	2.0E+09
4	Dhala	Asia	11	1.9E+09
5	Keuruselkä	Europe	30	1.8E+09
6	Paasselkä	Europe	10	1.8E+09
7	Shoemaker (was Teague)	Australia	30	1.5E+09
8	Matt Wilson	Australia	7.5	1.4E+09
9	<i>Ullapool</i>	<i>Europe</i>	<i>50</i>	<i>1.2E+09</i>
10	Amelia Creek	Australia	20	1.1E+09
11	Lumparn	Europe	9	1.0E+09
12	Suvasvesi North	Europe	4	1.0E+09
13	Jämsjärvi	Asia	14	7.0E+08
14	Strangways	Australia	25	6.5E+08
15	Söderfjärden	Europe	6.6	6.0E+08
16	Acraman	Australia	90	5.9E+08
17	Luizi	Africa	17	5.8E+08
18	Spider	Australia	13	5.7E+08
19	Sääksjärvi	Europe	6	5.6E+08
20	Kelly West	Australia	10	5.5E+08
21	<i>Massive Australian Precambrian/Cambrian Impact Structure, MAPCIS</i>	<i>Australia</i>	<i>1250</i>	<i>5.5E+08</i>
22	Foelsche	Australia	6	5.5E+08
23	Lawn Hill	Australia	18	5.2E+08
24	Glikson	Australia	19	5.1E+08
25	<i>Wilkes Land</i>	<i>Antarctica</i>	<i>485</i>	<i>5.0E+08</i>
26	Gardnos	Europe	5	5.0E+08
27	Mizarai	Europe	5	5.0E+08
28	Neugrund	Europe	8	4.7E+08
29	<i>Decorah crater</i>	<i>North America</i>	<i>5.6</i>	<i>4.7E+08</i>
30	Lockne	Europe	7.5	4.6E+08
31	Käröla	Europe	4	4.6E+08
32	<i>Ishim</i>	<i>Asia</i>	<i>350</i>	<i>4.5E+08</i>
33	<i>Tokraakaya</i>	<i>Asia</i>	<i>220</i>	<i>4.5E+08</i>
34	<i>Jeptha Knob</i>	<i>North America</i>	<i>4.3</i>	<i>4.3E+08</i>
35	Kaluga	Asia	15	3.8E+08
36	Iltvintets	Europe	8.5	3.8E+08

37	Siljan	Europe	52	3.8E+08
38	<i>Panther Mountain</i>	<i>North America</i>	<i>10</i>	<i>3.8E+08</i>
39	<i>Alamo bolide impact</i>	<i>North America</i>	<i>assumed > 4 km given age</i>	<i>3.7E+08</i>
40	Woodleigh	Australia	90	3.6E+08
41	Piccarinry	Australia	7	3.6E+08
42	Gweni-Fada	Africa	14	3.5E+08
43	Aorounga	Africa	12.6	3.5E+08
44	<i>East Warburton Basin</i>	<i>Australia</i>	<i>200</i>	<i>3.3E+08</i>
45	<i>West Warburton Basin</i>	<i>Australia</i>	<i>200</i>	<i>3.3E+08</i>
46	<i>Weaubleau-Osceola</i>	<i>North America</i>	<i>17.5</i>	<i>3.3E+08</i>
47	<i>Unnamed impact</i>	<i>Australia</i>	<i>130</i>	<i>3.0E+08</i>
48	<i>Unnamed impact</i>	<i>Australia</i>	<i>120</i>	<i>3.0E+08</i>
49	Serra da Cangalha	South America	12	3.0E+08
50	Dobeles	Europe	4.5	2.9E+08
51	Temovka	Europe	11	2.8E+08
52	Kursk	Asia	6	2.5E+08
53	<i>Bedout</i>	<i>Australia</i>	<i>200</i>	<i>2.5E+08</i>
54	Araguaína	South America	40	2.4E+08
55	Saqqaq	Asia	34	2.4E+08
56	<i>Saqqaq*</i>	<i>Asia</i>	<i>34</i>	<i>2.4E+08</i>
57	Rochechouart	Europe	23	2.1E+08

58	<i>Guarás</i>	<i>Europe</i>	<i>30</i>	<i>2.0E+08</i>
59	Riachão Ring	South America	4.5	2.0E+08
60	Obolon	Europe	20	1.7E+08
61	Puchezh-Katunki	Asia	80	1.7E+08
62	Vepriai	Europe	8	1.6E+08
63	Morokweng	Africa	70	1.5E+08
64	Gosses Bluff	Australia	22	1.4E+08
65	Mjotrir	Europe	40	1.4E+08
66	Tookoonooka	Australia	55	1.3E+08
67	Miten	Europe	9	1.2E+08
68	Oasis	Africa	18	1.2E+08
69	Mount Toondina	Australia	4	1.1E+08
70	<i>Kebra</i>	<i>Africa</i>	<i>31</i>	<i>1.0E+08</i>
71	Dellen	Europe	19	8.9E+07
72	<i>Praia Grande</i>	<i>South America</i>	<i>20</i>	<i>8.4E+07</i>
73	Lappajärvi	Europe	23	7.3E+07
74	Kara	Asia	65	7.0E+07
75	Tin Bider	Africa	6	7.0E+07
76	Chukcha	Asia	6	7.0E+07
77	<i>Bow City</i>	<i>North America</i>	<i>8</i>	<i>7.0E+07</i>
78	Vargeão Dome	South America	12	7.0E+07
79	Boltysh	Europe	24	6.5E+07
80	<i>Shiva</i>	<i>Asia</i>	<i>600</i>	<i>6.5E+07</i>
81	Vista Alegre	South America	9.5	6.5E+07
82	Chicxulub	North America	170	6.5E+07
83	<i>Wambo-Nyamaring structure</i>	<i>Africa</i>	<i>41</i>	<i>6.0E+07</i>
84	Connolly Basin	Australia	9	6.0E+07
85	<i>Silverpit</i>	<i>Europe</i>	<i>20</i>	<i>6.0E+07</i>
86	Goat Paddock	Australia	5	5.0E+07
87	Kamensk	Asia	25	4.9E+07
88	Jabal Waqf es Swwan	Asia	5.5	4.7E+07
89	Ragozinka	Asia	9	4.6E+07
90	Chiyli	Asia	5.5	4.6E+07
91	<i>Victoria Island structure</i>	<i>North America</i>	<i>5.5</i>	<i>4.3E+07</i>
92	Lогоisk	Europe	15	4.2E+07
93	Logancha	Asia	20	4.0E+07
94	Beyenchim e-Salaatin	Asia	8	4.0E+07

95	Popigai	Asia	100	3.6E+07
96	Flaxman	Australia	10	3.5E+07
97	Crawford	Australia	8.5	3.5E+07
98	<i>Toms Canyon</i>	<i>North America</i>	<i>22</i>	<i>3.5E+07</i>
99	<i>Vichada Structure</i>	<i>South America</i>	<i>50</i>	<i>3.0E+07</i>
100	<i>Ross</i>	<i>Antarctica</i>	<i>550</i>	<i>2.8E+07</i>
101	Nördlinger Ries	Europe	25	1.5E+07
102	Karakul	Asia	52	5.0E+06
103	Karla	Asia	10	5.0E+06
104	Bigach	Asia	8	5.0E+06
105	El'gygytyn	Asia	18	3.5E+06
106	<i>Corossil</i>	<i>North America</i>	<i>4</i>	<i>2.6E+06</i>
107	Bosumtwi	Africa	10.5	1.1E+06
108	<i>Pantasma</i>	<i>North America</i>	<i>10</i>	<i>1.0E+06</i>
109	Zhamanshin	Asia	14	9.0E+05
110	Rio Cuarto	South America	4.5	1.0E+05
111	<i>Iturralde</i>	<i>South America</i>	<i>8</i>	<i>2.1E+04</i>
112	<i>Zerelia West</i>	<i>Europe</i>	<i>250</i>	<i>7.0E+03</i>
113	<i>Zerelia East</i>	<i>Europe</i>	<i>150</i>	<i>7.0E+03</i>
114	Sudbury	North America	250	1.9E+03
115	<i>Spence</i>	<i>Europe</i>	<i>127.5</i>	<i>1.7E+03</i>
116	Santa Fe	North America	9.5	1.2E+03
117	Beaverhead	North America	60	6.0E+02
118	Rock Elm	North America	6	5.1E+02
119	Presqu'île	North America	24	5.0E+02

#	Name/Location	Continent	Size (km)	Age (yr)
120	Glover Bluff	North America	8	5.0E+02
121	Ames	North America	16	4.7E+02
122	Slate Islands	North America	30	4.5E+02
123	Calvin	North America	8.5	4.5E+02
124	Pilot	North America	6	4.5E+02
125	Couture	North America	8	4.3E+02
126	Glasford	North America	4	4.3E+02
127	Nicholson	North America	12.5	4.0E+02
128	La Moirerie	North America	8	4.0E+02
129	Elbow	North America	8	4.0E+02
130	Charlevoix	North America	54	3.4E+02
131	Serpent Mound	North America	8	3.2E+02
132	Crooked Creek	North America	7	3.2E+02
133	Decaturville	North America	6	3.0E+02
134	Middlesboro	North America	6	3.0E+02
135	Île Rouleau	North America	4	3.0E+02
136	Lac à l'Eau Claire Ouest	North America	36	2.9E+02
137	Lac à l'Eau Claire Est	North America	26	2.9E+02
138	Des Plaines	North America	8	2.8E+02
139	Gow	North America	4	2.5E+02
140	Tunnunik	North America	25	2.4E+02
141	Saint Martin	North America	40	2.2E+02
142	Manicouagan	North America	100	2.1E+02
143	Wells Creek	North America	12	2.0E+02
144	Red Wing	North America	9	2.0E+02
145	Cloud Creek	North America	7	1.9E+02
146	Upheaval Dome	North America	10	1.7E+02
147	Carswell	North America	39	1.2E+02
148	Sierra Madera	North America	13	1.0E+02
149	Deep Bay	North America	13	9.9E+01
150	Kentland	North America	13	9.7E+01
151	Avak	North America	12	9.5E+01
152	Steen River	North America	25	9.1E+01
153	Wetumpka	North America	7.6	8.3E+01
154	Santa Marta	South America	10	8.3E+01
155	Maple Creek	North America	6	7.5E+01
156	Manson	North America	35	7.4E+01
157	Eagle Butte	North America	10	6.5E+01

158	Marquez	North America	12.7	5.8E+01
159	Montagnais	North America	45	5.1E+01
160	Haughton	North America	23	3.9E+01
161	Wanapitei	North America	7.5	3.7E+01
162	Mistastin	North America	28	3.6E+01
163	Chesapeake Bay	North America	90	3.6E+01
164	Bowers	Antarctica	100	Unknown
165	Snows Island	North America	11	Unknown
166	Cerro Jarca	South America	10	Unknown
167	Brent	North America	3.8	3.1E+08
168	Suvasvesi South	Europe	3.8	2.5E+08
169	Steinheim	Europe	3.8	1.5E+07
170	Flynn Creek	North America	3.8	3.6E+02
171	Colônia	South America	3.6	2.1E+01
172	Kgagodi	Africa	3.5	1.8E+08
173	Zeleny Gai	Europe	3.5	8.0E+07
174	Quarkiz	Africa	3.5	7.0E+07
175	Pingualut	North America	3.44	1.4E+00
176	Zapadnaya	Europe	3.2	1.7E+08
177	Newport	North America	3.2	5.0E+02
178	Goyder	Australia	3	1.4E+09
179	Iso-Naakkima	Europe	3	1.0E+09
180	Granby	Europe	3	4.7E+08
181	Gusev	Asia	3	4.9E+07
182	Agoudal	Africa	3	1.1E+05

#	Name/Location	Continent	Size (km)	Age (yr)
183	Ramgarh	Asia	3	Unknown
184	Gatun structure	North America	2.85	2.0E+07
185	Mahas	Africa	2.85	Unknown
186	Shanak	Asia	2.8	4.5E+07
187	Rotmistrovka	Europe	2.7	1.2E+08
188	Ritland crater	Europe	2.7	5.2E+02
189	Mishina Gora	Asia	2.5	3.0E+08
190	Roter Kamm	Africa	2.5	3.7E+06
191	Viewfield	North America	2.5	1.9E+02
192	West Hawk	North America	2.44	3.5E+02
193	Holleford	North America	2.35	5.5E+02
194	Tvären	Europe	2	4.6E+08
195	BP Structure	Africa	2	1.2E+08
196	Brushy Creek Feature	North America	2	2.1E+04
197	Tenoumer	Africa	1.9	2.1E+04
198	Lonar	Asia	1.83	5.2E+04
199	Xiuyan crater	Asia	1.8	5.0E+04
200	Talen zone	Africa	1.75	3.0E+06
201	Liverpool	Australia	1.6	7.7E+08
202	Saanjärvi	Europe	1.5	6.0E+08
203	Karikkoselkä	Europe	1.4	2.3E+08
204	Tabun-Khara-Obo	Asia	1.3	1.5E+08
205	Hummeln structure	Europe	1.2	4.6E+08
206	Darwin Crater	Australia	1.2	8.0E+05
207	Barringer	North America	1.19	4.9E-02
208	Tswaing (was Pretoria Saltpan)	Africa	1.13	2.2E+05
209	Malingen	Europe	1	4.6E+08
210	Wolfe Creek	Australia	0.87	3.0E+05
211	Temimichka	Africa	0.75	Unknown
212	Kalkkop	Africa	0.64	2.5E+05
213	Chesko	Asia	0.5	1.0E+02
214	Monturaqui	South America	0.46	1.0E+06
215	Amguid	Africa	0.45	1.0E+05
216	Aouelloul	Africa	0.39	3.0E+06
217	Macha	Asia	0.3	7.0E+03
218	Hicoman Crater	Australia	0.27	3.2E+04

219	Boxhole	Australia	0.17	5.4E+03
220	Odessa	North America	0.168	5.0E-02
221	Henbury	Australia	0.16	4.2E+03
222	Patonsky	Asia	0.16	3.0E+02
223	Wabar	Asia	0.116	1.5E+02
224	Kaali	Europe	0.11	4.0E+03
225	Morasko	Europe	0.1	1.0E+04
226	Veevers	Australia	0.08	2.0E+04
227	Ilumetsa	Europe	0.08	6.6E+03
228	Sobolev	Asia	0.053	1.0E+03
229	Campo del Cielo	South America	0.05	4.0E+03
230	Kamil	Africa	0.045	2.0E+03
231	Whitecourt	North America	0.04	1.1E-03
232	Sikhote-Alin	Asia	0.026	7.0E+01
233	Dalgaranga	Australia	0.02	3.0E+03
234	Haviland	North America	0.015	1.0E-03
235	Carancas	South America	0.0135	7.0E-06

Italicized entries (*red*) indicate "unconfirmed" craters. For size and age, if listing indicates finite range, midpoint value is assumed (geometric if range is order of magnitude or more); if open-ended range (i.e., > x or < x), minimum or maximum cited value (x) is assumed. Unknown entries excluded except for Alamo Bolidé Impact which, given age, is assumed originally > 4km in size to be detectable today.

**Table II. Estimating the Rate of Lunar Cratering over the Corresponding Esker Atmospheric Eras
for Craters at Least 4 km in Size**

<u>Atmospheric Eras (based on Esker [2])</u>			<u>Lunar Rate (/km²)</u>				
<u>Name</u>	<u>Start (y)</u>	<u>Finish (y)</u>	<u>At Start</u>	<u>At Finish</u>	<u>Geo Mean</u>	<u>Combined</u>	
<u>Thick1</u>	2.4E+09	3.4E+08	8.5E-05	1.1E-05	3.06E-05	Both Thick	2.85E-05
<u>Thin1</u>	3.4E+08	2.3E+08	1.1E-05	8.5E-06	9.67E-06	Both Thin	6.99E-06
<u>Thick2</u>	2.3E+08	5.3E+07	8.5E-06	3.0E-06	5.05E-06	Ratio	
<u>Thin2</u>	5.3E+07	0.0E+00	3.0E-06	1.0E-06	1.73E-06	Thin/Thick	<i>2.45E-01</i>

Table III. Estimating the Rate of Earth Cratering for the Esker Atmospheric Eras for Craters at Least 4 km in Size

<u>Atmospheric Eras (based on Esker [2])</u>			<u>Earth Rate - to End of Ice Age</u>				<u>Earth Rate - To Present</u>			
<u>Name</u>	<u>Start (y)</u>	<u>Finish (y)</u>	<u># of Craters</u>	<u>1/y-km²</u>	<u>Combined</u>	<u>1/km²</u>	<u># of Craters</u>	<u>1/y-km²</u>	<u>Combined</u>	<u>1/km²</u>
Thick1	2.4E+09	3.4E+08	43	1.41E-16	Both Thick	4.87E-07	43	1.41E-16	Both Thick	4.87E-07
Thin1	3.4E+08	2.3E+08	13	8.44E-16	Both Thin	2.64E-07	13	8.44E-16	Both Thin	6.15E-07
Thick2	2.3E+08	5.3E+07	29	1.10E-15	Ratio		29	1.10E-15	Ratio	
Thin2	5.3E+07	0.0E+00	26	3.31E-15	Thin/Thick	5.42E-01	78	9.94E-15	Thin/Thick	1.26E+00