Klein-Gordon Equation and Wave Function in Cosmological Special Theory of Relativity

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ABSTRACT

In the Cosmological Special Theory of Relativity, we study energy-momentum relations, Klein-Gordon equation and wave function.

PACS Number:03.30, 41.20

Key words: Cosmological special relativity theory;

Klein-Gordon equation;

Energy-momentum relation

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1. Introduction

Our article's aim is that we make Klein-Gordon equation and wave function in cosmological special theory of relativity.

At first, space-time relations are in cosmological special theory of relativity (CSTR).[7]

$$C t = \gamma \left(C t + \frac{V_0}{C} \Omega \left(t \right) \right) \times \mathcal{N} \Omega(t_0) = \gamma \left(\Omega(t_0) \times V + V_0 \Omega(t_0) t' \right)$$

$$\Omega(t_0) y = \Omega(t_0) y',$$

$$\Omega(t_0) z = \Omega(t_0) z',$$

$$\gamma = 1 / \sqrt{1 - \frac{V_0^2}{C^2} \Omega^2(t_0)}, \quad t_0 \text{ is cosmological time}$$
(1)

Therefore, proper time is[7]

$$d\tau^{2} = dt^{2} - \frac{1}{c^{2}} \Omega^{2}(t_{0}) [dx^{2} + dy^{2} + dz^{2}]$$

$$= dt^{2} - \frac{1}{c^{2}} \Omega^{2}(t_{0}) [dx^{2} + dy^{2} + dz^{2}]$$

$$= dt^{2} - \frac{1}{c^{2}} \Omega^{2}(t_{0}) [dx^{2} + dy^{2} + dz^{2}]$$
(2)

Hence, energy-momentum relations are by the fact that energy-momentum are 4-vector in CSTR,

$$E = \gamma(E' + V_0 \Omega^2(t_0) \rho_x'), \rho_x \Omega(t_0) = \gamma(\Omega(t_0) \rho_x' + \frac{V_0}{c^2} \Omega(t_0) E')$$

$$\Omega(t_0) p_y = \Omega(t_0) p_y',
\Omega(t_0) p_z = \Omega(t_0) p_z',
\Omega(t_0) p_z = \Omega(t_0) p_z',
\gamma = 1 / \sqrt{1 - \frac{v_0^2}{c^2} \Omega^2(t_0)}, \\
E = m_0 c^2 \frac{dt}{d\tau}, \\
\vec{p} = m_0 \frac{d\vec{x}}{d\tau} \tag{3}$$

Therefore, energy-momentum-mass relation is in CSTR,

$$m_0^2 c^4 = E^2 - \Omega^2(t_0) \rho^2 c^2 \tag{4}$$

2. Klein-Gordon Equation and Wave Fuction in CSTR

According to [7], matter wave function is in CSTR,

$$\phi = \phi_{\rm 0} \exp i \Phi = \phi_{\rm 0} \exp i [\frac{\omega t}{\sqrt{\Omega(t_{\rm 0})}} - \vec{k} \cdot \vec{x} \sqrt{\Omega(t_{\rm 0})}]$$

$$=\phi^{\scriptscriptstyle \parallel}=\phi_{\scriptscriptstyle 0}\exp{i\Phi^{\scriptscriptstyle \parallel}}=\phi_{\scriptscriptstyle 0}\exp{i[\frac{\omega^{\scriptscriptstyle \parallel}t^{\scriptscriptstyle \parallel}}{\sqrt{\Omega(t_{\scriptscriptstyle 0})}}-\vec{k}^{\scriptscriptstyle \parallel}\cdot\vec{x}^{\scriptscriptstyle \parallel}\sqrt{\Omega(t_{\scriptscriptstyle 0})}]}$$

$$\phi_0$$
 is amplitude, ω is angular frequency, $k = |\vec{k}|$ is wave number. (5)

If we use Eq(1) in Eq(5), we obtain angular frequency-wave number relation.

$$\omega' = \gamma(\omega - V_0 \Omega(t_0) k_1), \ k_1' = \gamma(k_1 - \frac{V_0}{C^2} \Omega(t_0) \omega)$$

$$k_2' = k_2, k_3' = k_3, \gamma = 1/\sqrt{1 - \frac{{V_0}^2}{C^2}\Omega^2(t_0)}$$
 (6)

In this time, if we define energy-momentum by angular frequency-wave number,

$$E = \hbar \omega, \vec{\rho} = \frac{\hbar \vec{k}}{\Omega(t_0)} \tag{7}$$

Hence, we obtain the angular frequency-wave number relation about the energy-momentum-mass relation in CSTR,

$$m_0^2 c^4 = E^2 - \Omega^2 (t_0) \rho^2 c^2 = \hbar^2 \omega^2 - \hbar^2 k^2 c^2$$
 (8)

We obtain next result by the transformation of the angular frequency-wave number relation, Eq(6) in CSTR.

$$m_0^2 c^4 = \hbar^2 \omega^2 - \hbar^2 k^2 c^2 = \hbar^2 \omega^{'2} - \hbar^2 k^{'2} c^2$$
 (9)

If we define the differential operator about energy-momentum in CSTR,

$$E = i\hbar\sqrt{\Omega(t_0)}\frac{\partial}{\partial t}, \vec{\rho} = -i\hbar\frac{1}{\Omega(t_0)\sqrt{\Omega(t_0)}}\vec{\nabla}$$
 (10)

If we apply Eq(10) to Eq(4),

$$m_0^2 c^4 = E^2 - \Omega^2(t_0) \rho^2 c^2 = \hbar^2 [-\Omega(t_0) (\frac{\partial}{\partial t})^2 + \frac{1}{\Omega(t_0)} c^2 \nabla^2]$$

We finally obtain Klein-Gordon equation in CSTR.

$$\frac{m_0^2 c^2}{\hbar^2} \phi = \left[-\Omega(t_0) \frac{1}{c^2} \left(\frac{\partial}{\partial t} \right)^2 + \frac{1}{\Omega(t_0)} \nabla^2 \right] \phi \tag{11}$$

Wave function, Eq(5) satisfy Klein-Gordon equation, Eq(11) in CSTR.

3. Conclusion

We are able to describe free particle by Klein-Gordon equation and wave function in CSTR.

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