

## Stochastic Cosmology and the MOND Paradigm

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### Abstract

This brief note points out that, if early cosmology drives gravitational dynamics out of equilibrium, a straightforward random walk model of orbital motion can explain away the basis of Modified Newtonian Dynamics (MOND).

**Key words:** stochastic cosmology, random walk, non-extensive statistical mechanics, MOND.

There are well-founded reasons to believe that early stages of Universe evolution drive gravitational dynamics out of thermodynamic equilibrium.

Consider this early regime and the classical Kepler problem involving a body gravitating about a central mass. The fluctuating behavior of the orbital motion may be described by a simple *random walk model*, whereby the

gravitating body drifts away from the central mass according to the standard diffusion equation. The grow over time of the mean radial distance is given by

$$\langle r \rangle \propto \sqrt{t} \quad (1)$$

and reflects the evolution towards states with maximal entropy. If the problem has spherical symmetry, the number of configurations  $\Omega$  in a shell of thickness  $dr$  is proportional to  $r^2$ . In this case, using the Boltzmann entropy

$$S = k_B \log \Omega \quad (2)$$

the expression of the mean radial force associated with the drift (1) can be shown to take the form

$$\langle F_r \rangle = T \left\langle \frac{dS}{dr} \right\rangle = \frac{2k_B T}{\langle r \rangle} \quad (3)$$

where  $T$  denotes the temperature [1-2]. Assume next that the gravitating body consists of a couple of components, each defined by entropies  $S_1$  and

$S_2$ . Because the system is out of equilibrium, its overall entropy contains a non-extensive term, and it reads

$$S = S_1 + S_2 + |q-1|S_1S_2 \quad (4)$$

where  $|q| < 1$  is the Tsallis parameter [3]. Relations (2)-(4) hint that that (4) splits in two contributions, one attributed to Newtonian dynamics ( $S_N$ ) and the other to MOND ( $S_{MOND}$ ), namely,

$$S = S_N + S_{MOND} \quad (5)$$

in which

$$S_N = S_1 + S_2 = k_B \log(\Omega_1 \Omega_2) \quad (6a)$$

$$S_{MOND} = |1-q|k_B^2 \log \Omega_1 \log \Omega_2 \quad (6b)$$

A plausible hypothesis is that the system temperature is non-uniform and it drops as reciprocal of the radial scale as in

$$T \propto \langle r \rangle^{-1} \quad (7)$$

It follows from (6a) and (7) that the Newtonian radial force (3) obeys an *inverse square law*, i.e.

$$\boxed{\langle F_r \rangle_N \propto \frac{1}{\langle r \rangle^2}} \quad (8)$$

whereas the correction to Newtonian force generated by (6b) takes the form

$$\boxed{\langle F_r \rangle_{MOND} \propto \frac{1}{\langle r \rangle} \left\langle \frac{dS_{MOND}}{dr} \right\rangle} \quad (9)$$

We close with several key observations:

- 1) Follow-up analysis must show that the sign of radial force (3) and (8) is *negative*, in sync with the attractive nature of gravitational interaction.
- 2) Dynamics induced by (8) and (9) is likely to carry over as “memory-like” effects to the current state of the Universe.
- 3) Our approach matches the idea that the value of the vacuum energy parameter and of the Hubble parameter follow from the stochastic dynamics of the early Universe [4-5]. It is also compatible with the

proposal that MOND is rooted in the *non-integrable* nature of gravitational interaction [6-8].

## **References**

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